

VLF to HF Ground and Spaceborne Instrumentation

for Active Radiowave Plasma Experiments



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Northeast Radio Observatory Corporation (NEROC) Symposium Series

NEROC Symposium • MIT Haystack Observatory • Nov 4, 2016

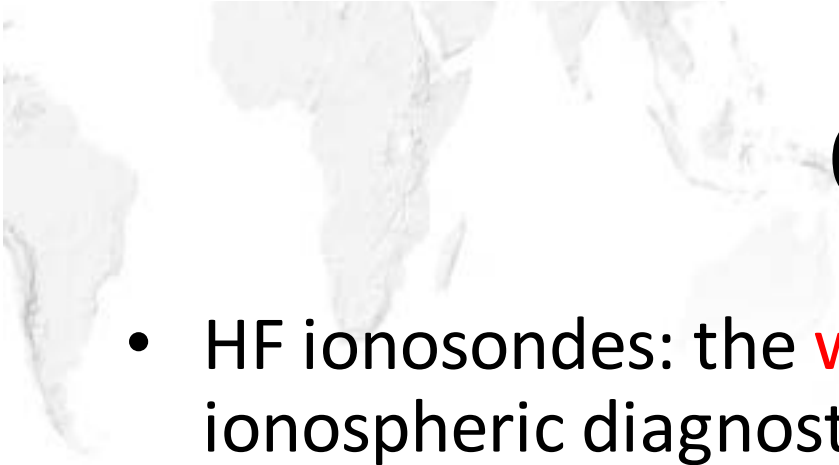
Acknowledgements

Lowell Team (UML, LDI)



Other teams:

- SwRI
- NASA Goddard SFC
- AFRL
- Able Engineering / ATK
- Net-TIDE Europe
- HAARP
- Kansas University
- Stanford University
- Rice University

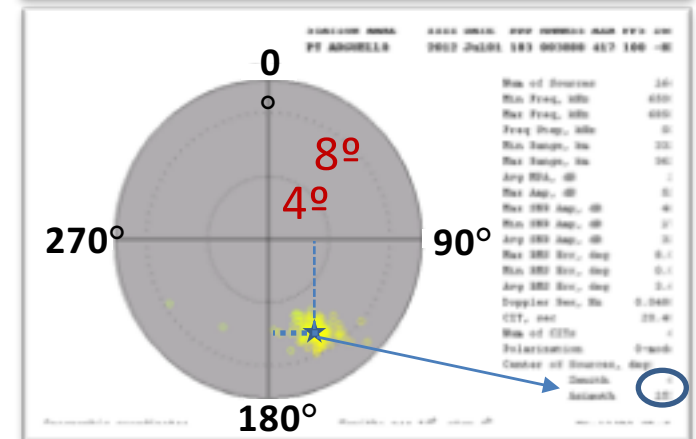
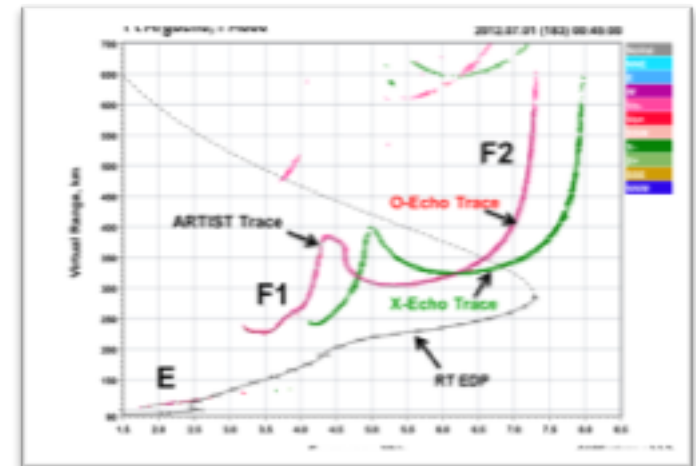
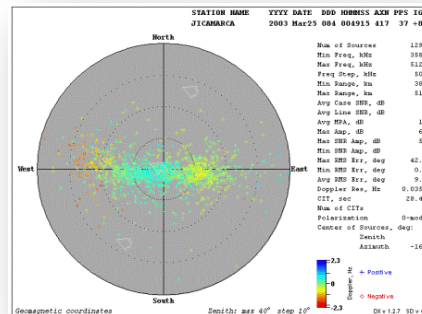
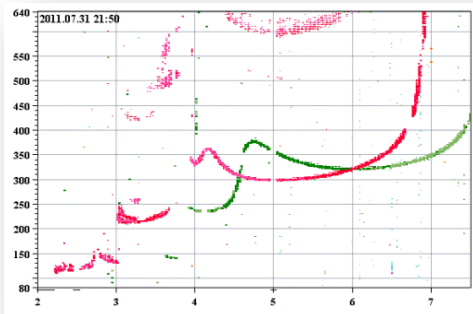


Outline

- HF ionosondes: the **workhorse** and the **ground truth** of ionospheric diagnostics with [near] **real-time** capability
 - Is there life after 85?
- **VLF transmission in space**
 - Radio Plasma Imager on NASA IMAGE
 - VLF Transmitter on USAF DSX
- Where do we go from here?
 - HF **Topside Sounder** / CubeSat challenge
 - Dual-band **Subsurface Radar** (GPR/IPR)

HF Ionosonde

- First ionogram: Jan 11, 1931
- 1936: five ionosondes in the world
- 1957: **150 ionosondes** in the world
- 1969: First Lowell Digisonde built
- 2016: <unknown> ionosondes in the world, but
 - 231 ionosonde locations registered in SPIDR (1942..current)
 - **164 Lowell digisondes**



Global Ionosphere Radio Observatory (GIRO)



GIRO Wall and Tower
in Lowell

Quiet Revolution in Sensor Networks

[NSF, March 2005]

- Access to data without time or distance barriers
- Remote control by operator
- Data interpretation without human attention
- Self-tested, “glass-room” operations
- Superior data management
- Automated, intelligent adjustment of sensor operation modes
- Data science (knowledge discovery)

DIGISONDE:

mid 1980s

ARTIST (1982)

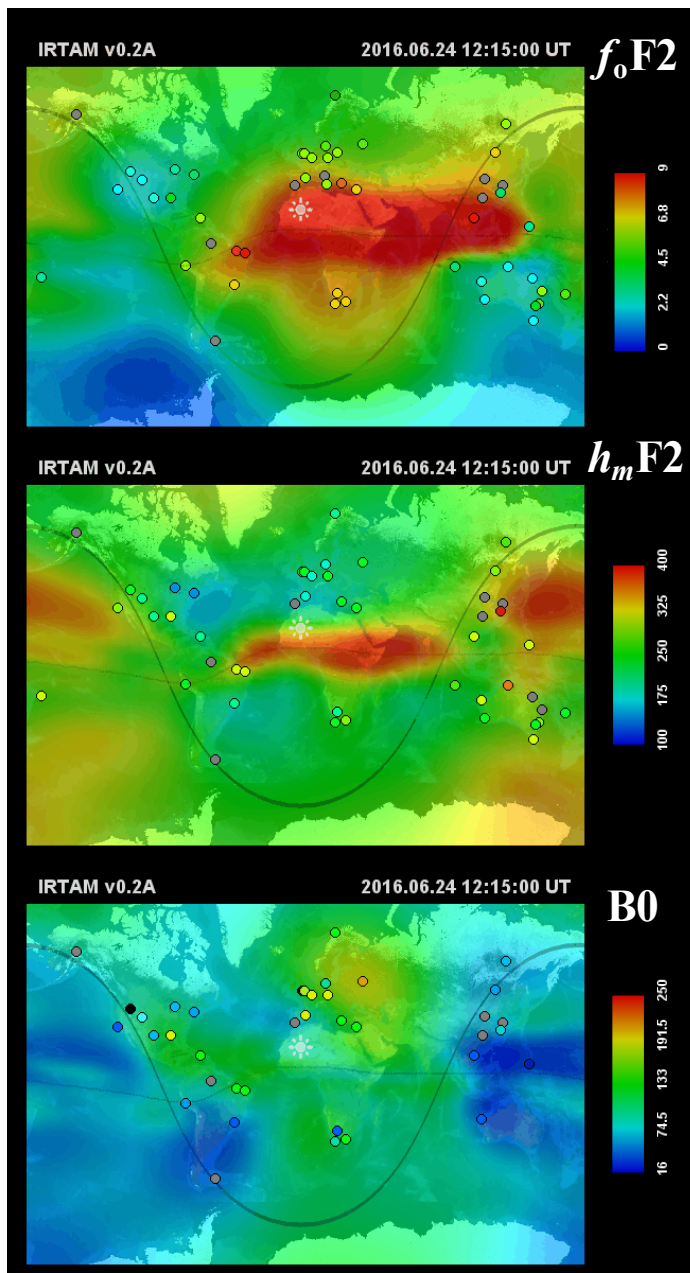
BIT (1995)

DIDBase (2002)

Drift-DB (2005)

ADRES (2004)

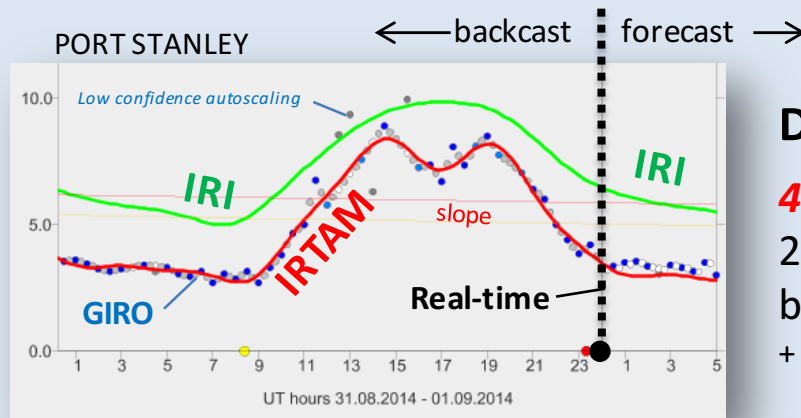
Ongoing...



Real-Time IRI

15 minute cadence

3D global bottomside ionosphere in real-time



Diurnal fit to GIRO data

4D data assimilation:
24-hour fit of differences
between IRI and GIRO
+ diurnal trend analysis

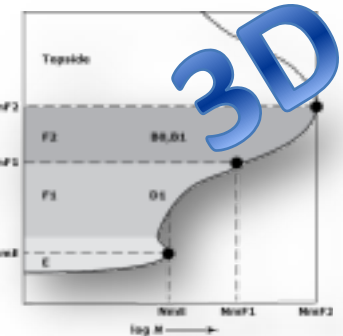
Port Stanley data courtesy of Sarah James, RAL UK

x 52 GIRO stations

Global Spatial fit

Jones-Gallet Gk basis (76)
= total map coefficients: **1064**
[24-hour global weather of C]

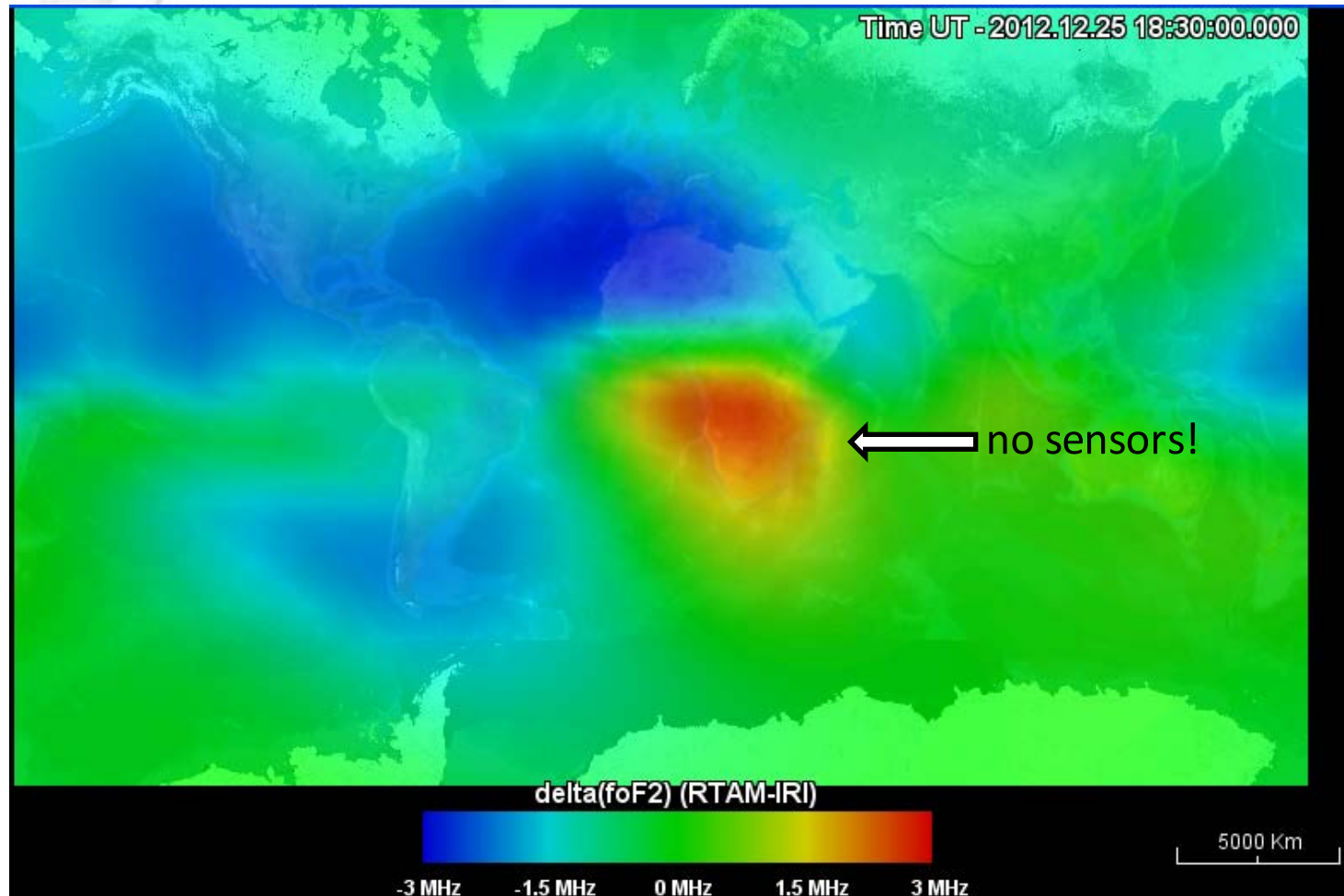
IRI profile formalism



NmF2, hmF2, B0, B1 anchors

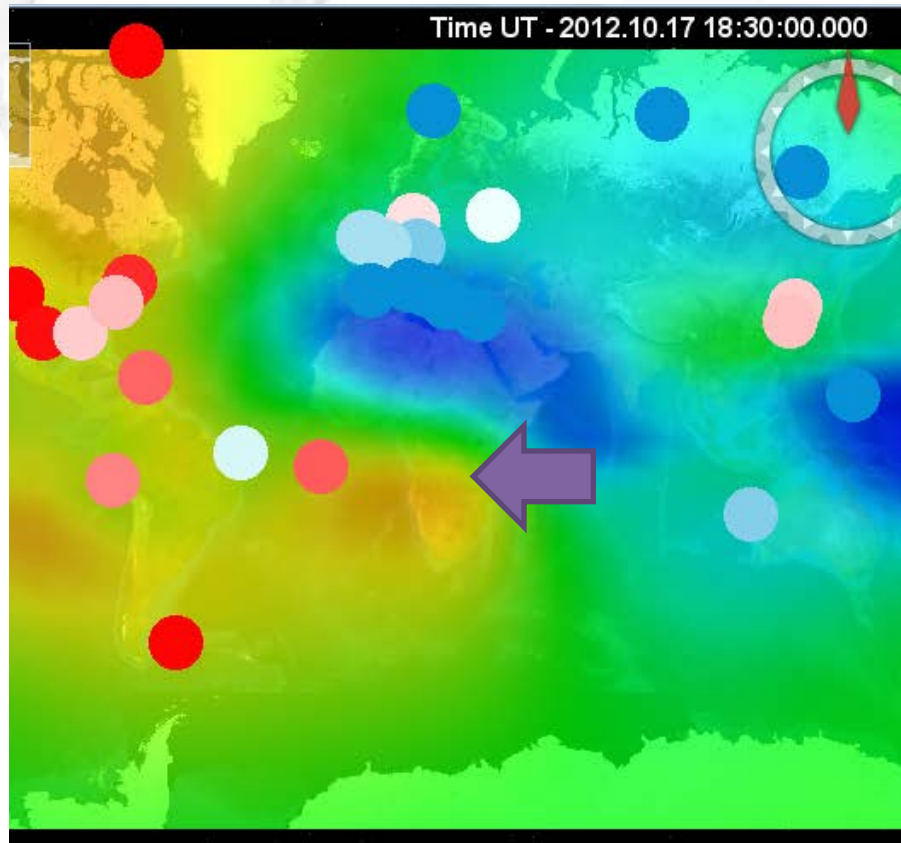
“Red Spot over Africa”

found on foF2 deviation map

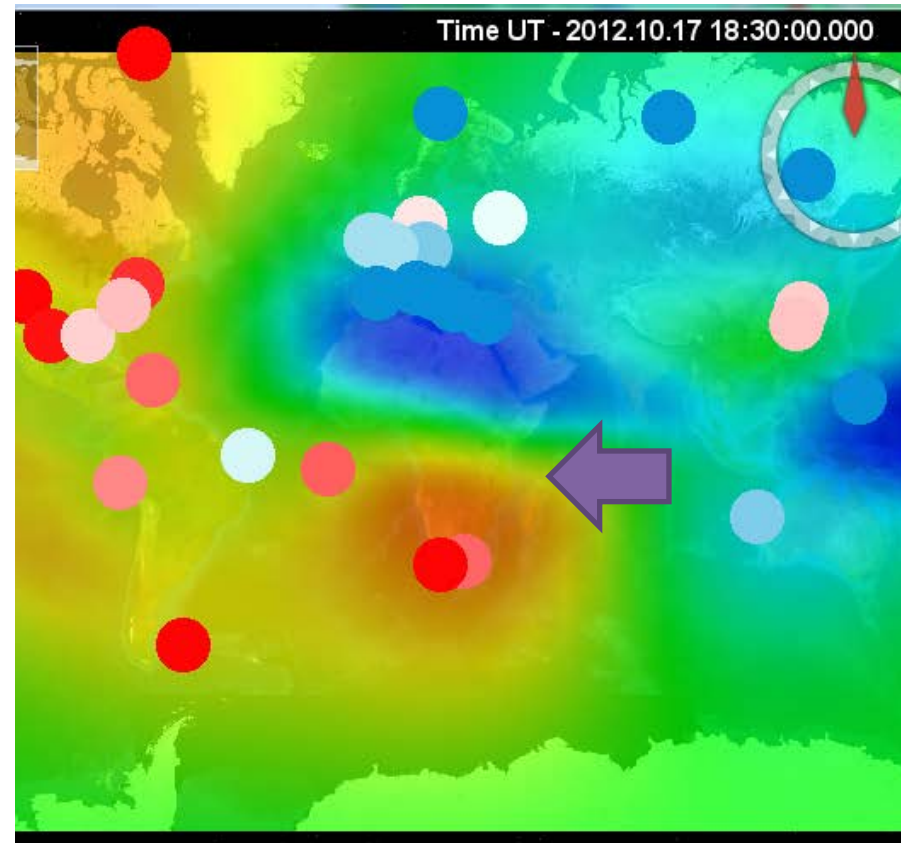


Control site analysis with RSA Data

Spatial prediction?



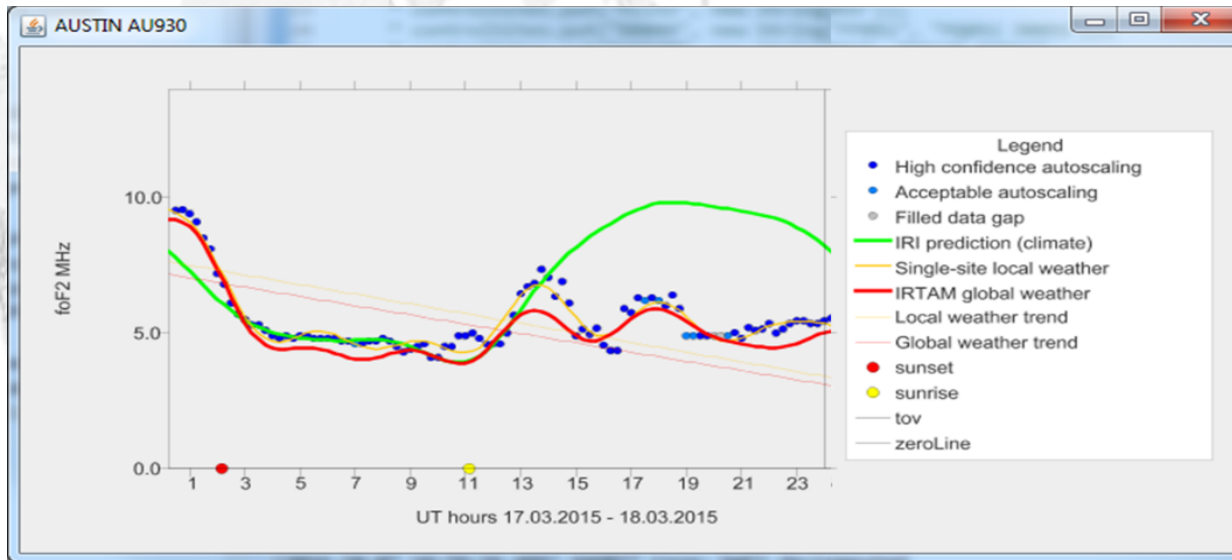
RSA data excluded



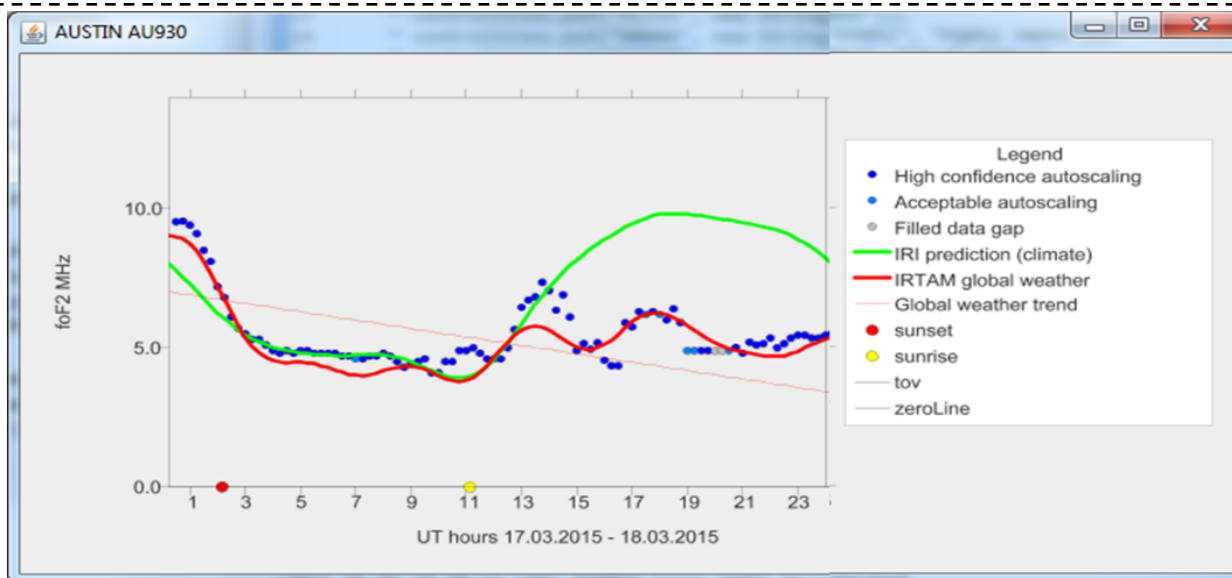
RSA data included

Control site comparison
of IRTAM assimilation

AUSTIN SITE INCLUDED



AUSTIN SITE EXCLUDED

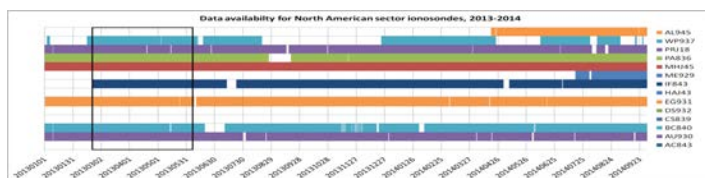


But neighboring stations help:

Eglin AFB, FL (1073 km)
Boulder, CO (1270 km)
Melrose, FL (1512 km)
Vandenberg AFB (2186 km)

Real-Time IRI Improvement of Climatology

Average improvement $\sim x2$ (15 million comparisons) in terms of $\langle \text{Obs-Model} \rangle$ error reduction



Test Period: 2013.02.19 to 2013.06.11

foF2 Improvement factor

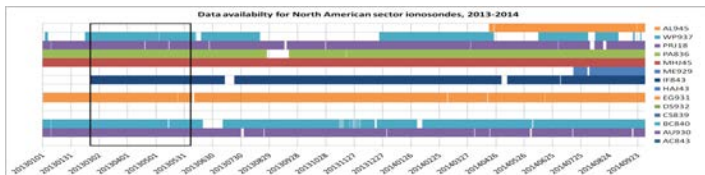
Comparisons with control site included vs excluded

Cntl Site	Included vs excluded	Lost Q
BC840	2.20 vs 1.75	20 %
IF843	1.92 vs 1.39	28 %
WP937	2.14 vs 1.46	32 %
AU930	2.10 vs 1.34	36 %
EG931	2.00 vs 1.28	36 %

IRTAM foF2 = typical $x2$ improvement

Green sites: loss of quality < 40%

Not all GIRO sites are equal



Test Period: 2013.02.19 to 2013.06.11

foF2 improvement factor

Comparisons with control site included vs excluded

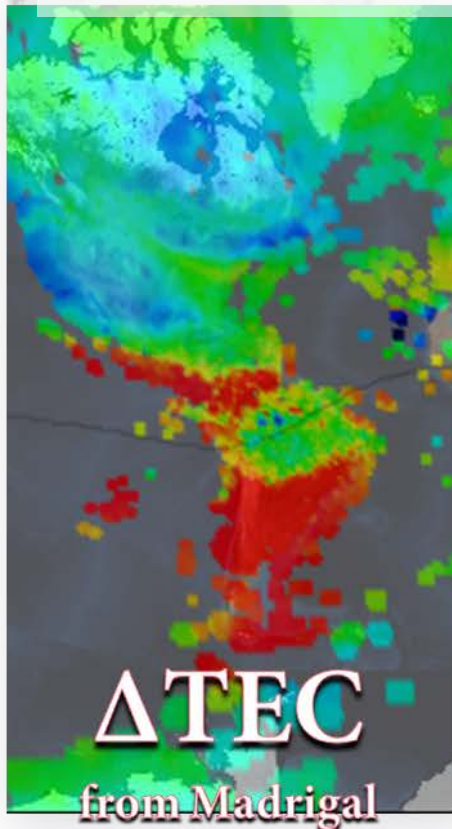
Cntl Site	Included vs excluded	Lost Q
MHJ45	2.11 vs 1.24	41 %
PA836	2.61 vs 1.29	51 %
PRJ18	2.86 vs 1.05	63 %

No improvement to IRI climatology when PRJ18 is omitted

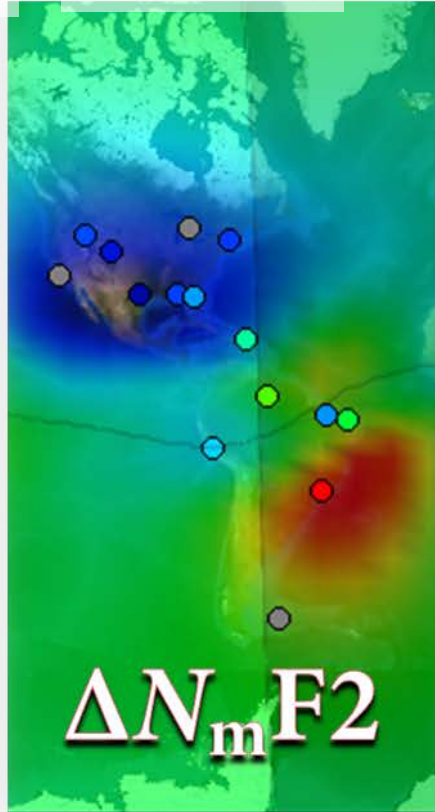
Red sites: loss of quality > 40%

In Combination with Madrigal TEC:

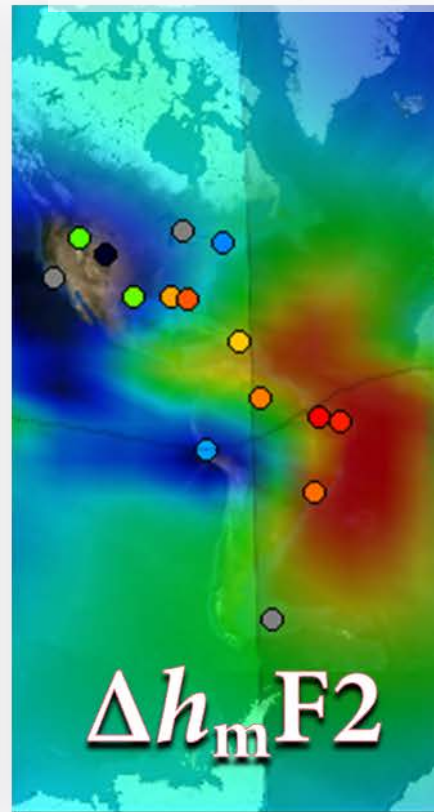
Total Electron Content



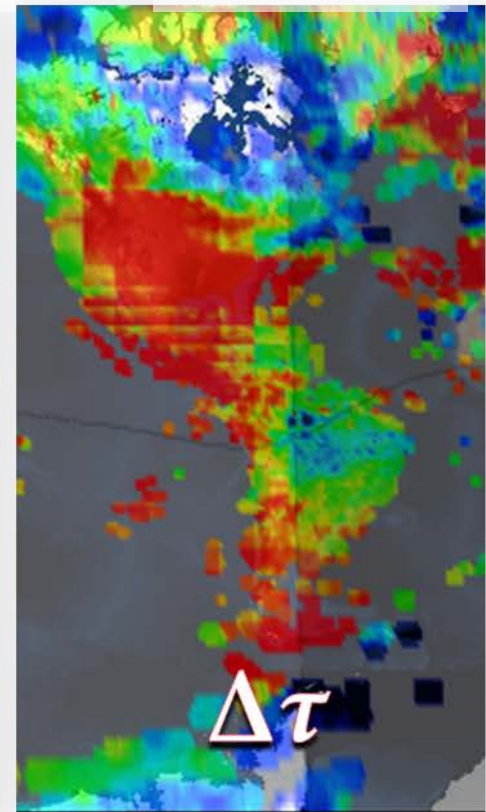
Peak Density



Peak Density Height



Slab Thickness



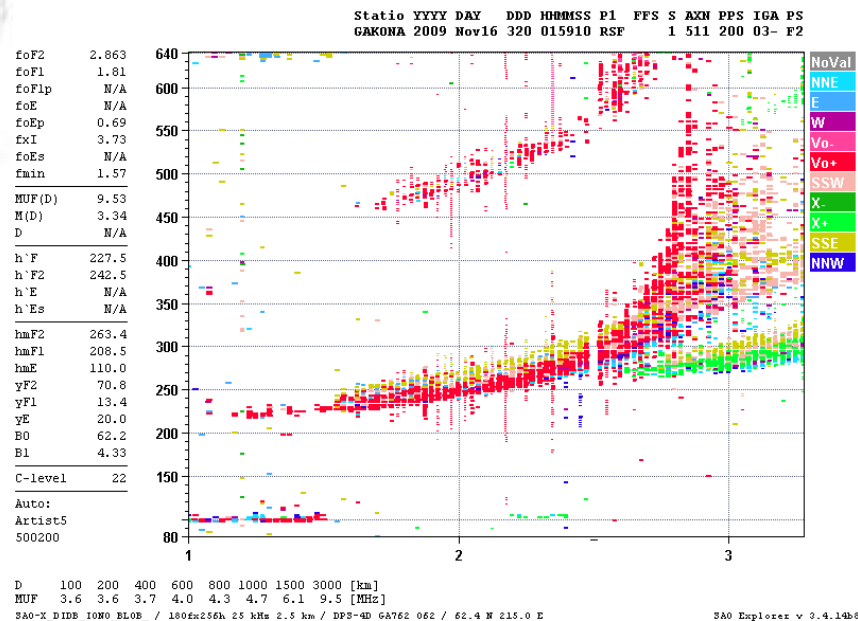
Deviation from expected quiet-time behavior

Red: larger than model Blue: smaller than model

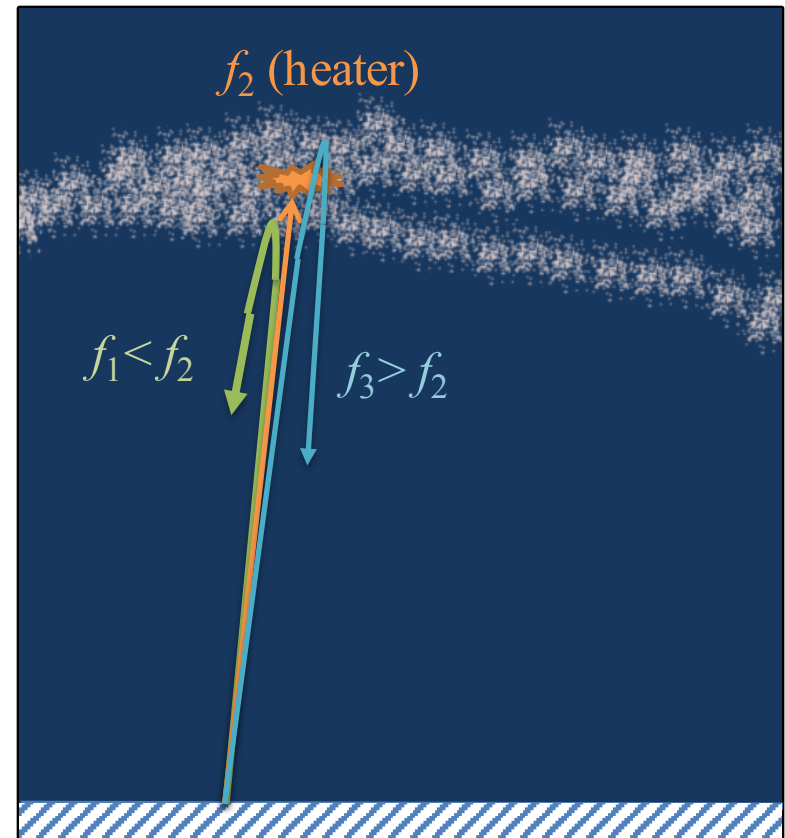


SINGLE-SITE HF IONOSONDE SCIENCE

HAARP Heating Experiments



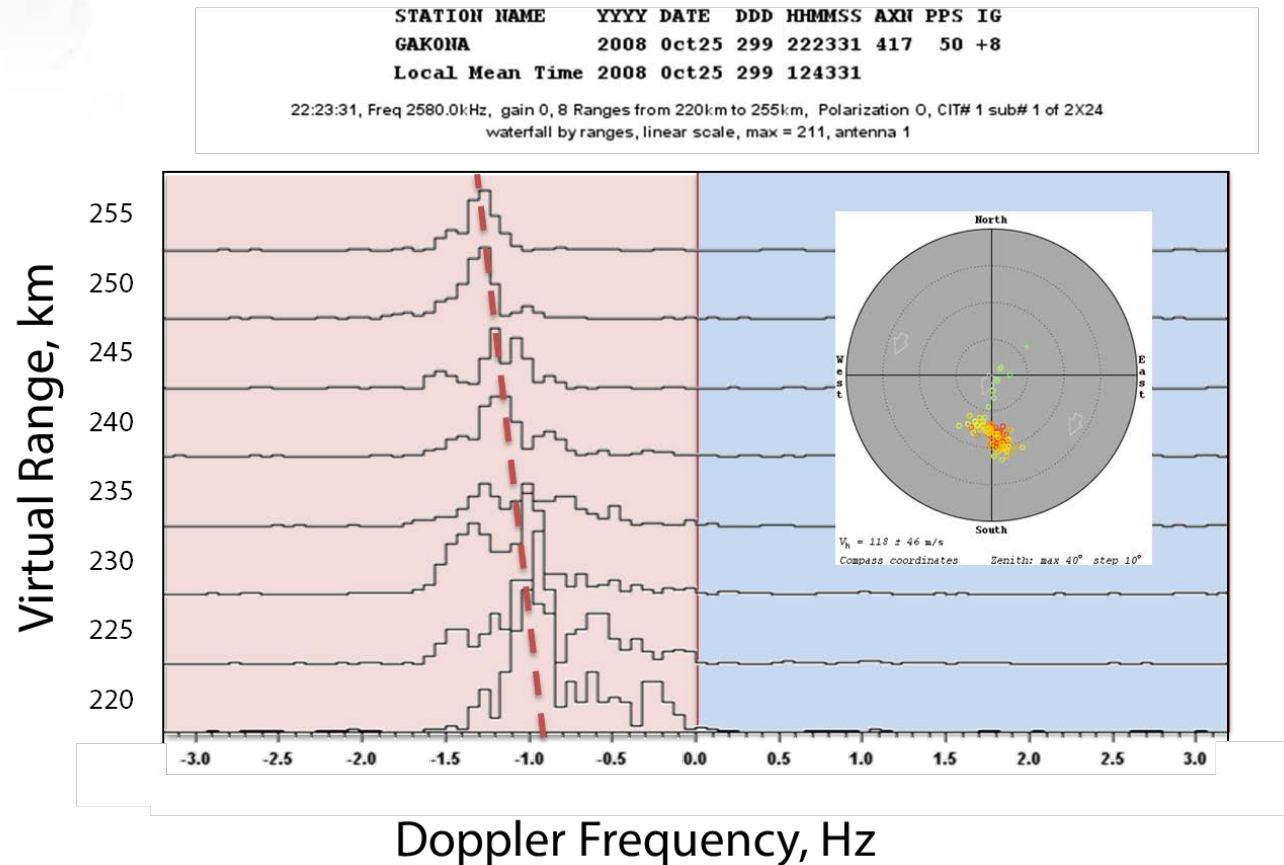
Ionograms show the bottomside of expanding plasma ball appearing as an artificial ionization layer



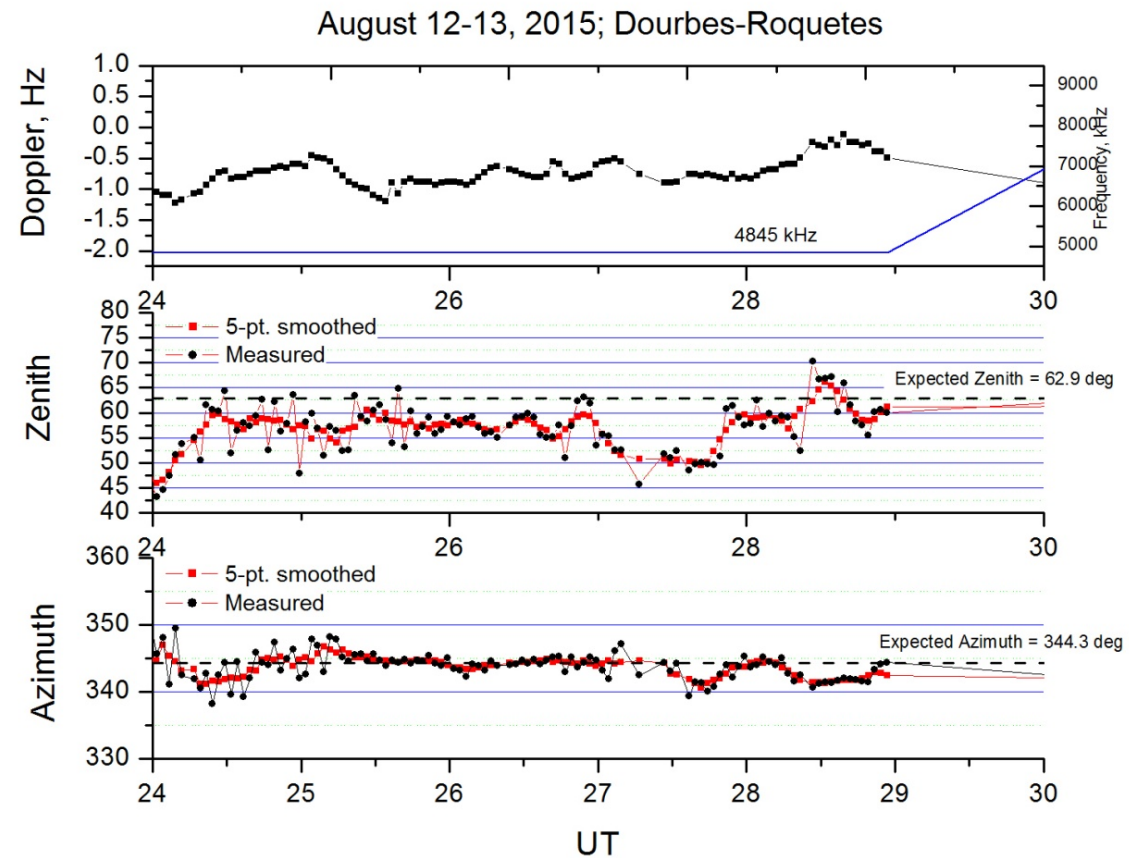
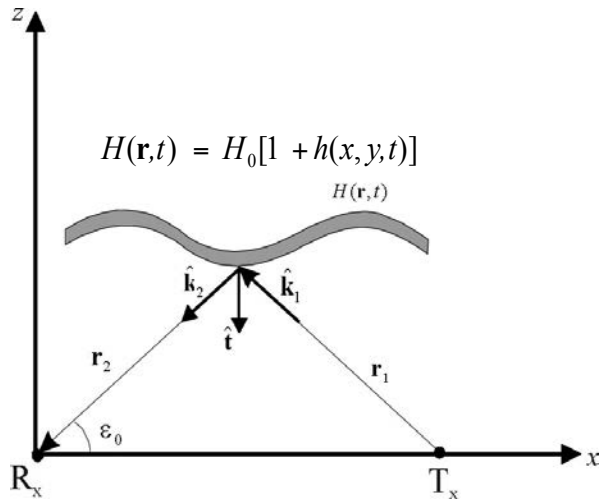
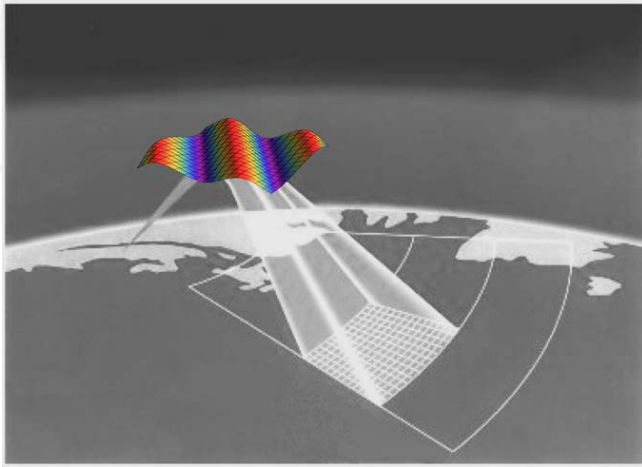
credits to Todd Pedersen and HAARP/AFRL team

Digisonde Skymap and Waterfall Display

Topside of the expanding plasma structure ~50 km wide, accelerating upward along the magnetic field line at 80-100 m/s

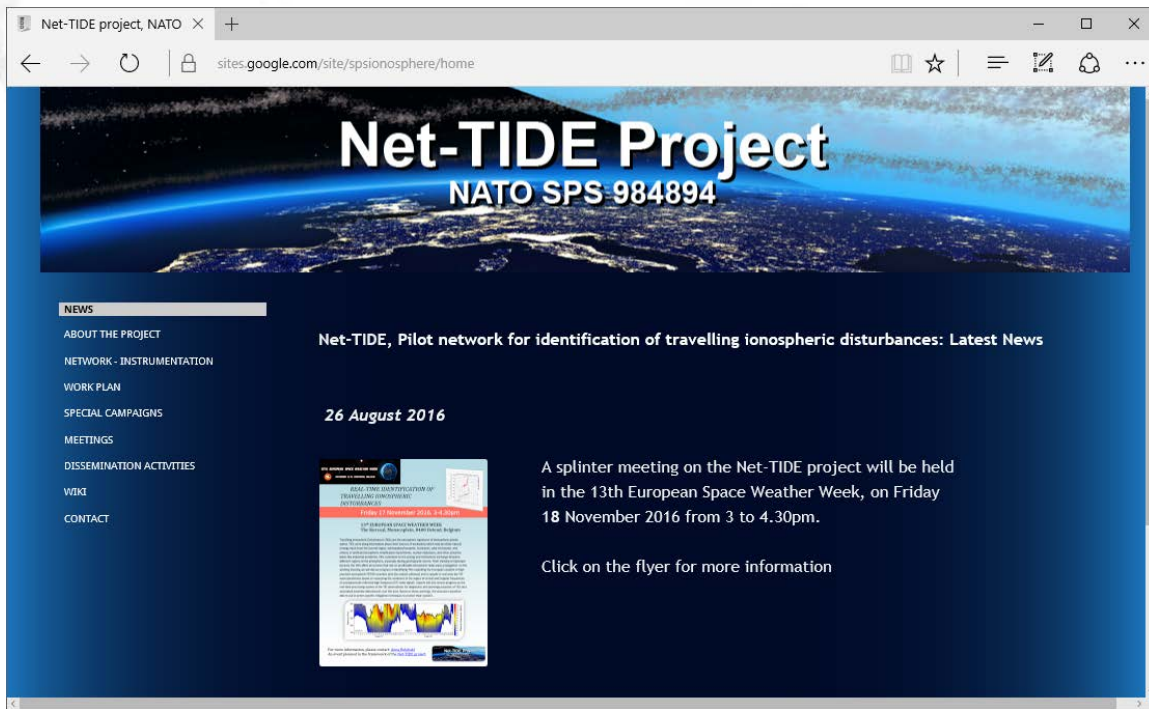


TID Detection and Evaluation



Net-TIDE Europe

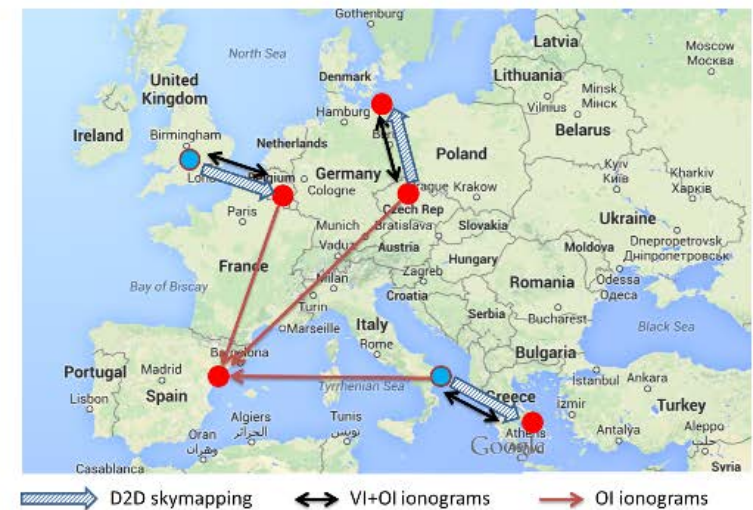
<https://sites.google.com/site/spsionosphere/>



The screenshot shows the Net-TIDE Project website. The header features a space-themed image with the text "Net-TIDE Project" and "NATO SPS 984894". A left sidebar contains a "NEWS" section and a list of project topics: ABOUT THE PROJECT, NETWORK - INSTRUMENTATION, WORK PLAN, SPECIAL CAMPAIGNS, MEETINGS, DISSEMINATION ACTIVITIES, WIKI, and CONTACT. The main content area has a sub-header "Net-TIDE, Pilot network for identification of travelling ionospheric disturbances: Latest News" and a date "26 August 2016". It includes a flyer for a "splinter meeting" on 18 November 2016 and a link to click for more information.

The DPS4D network in Europe

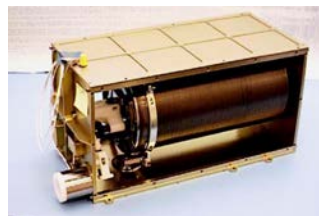
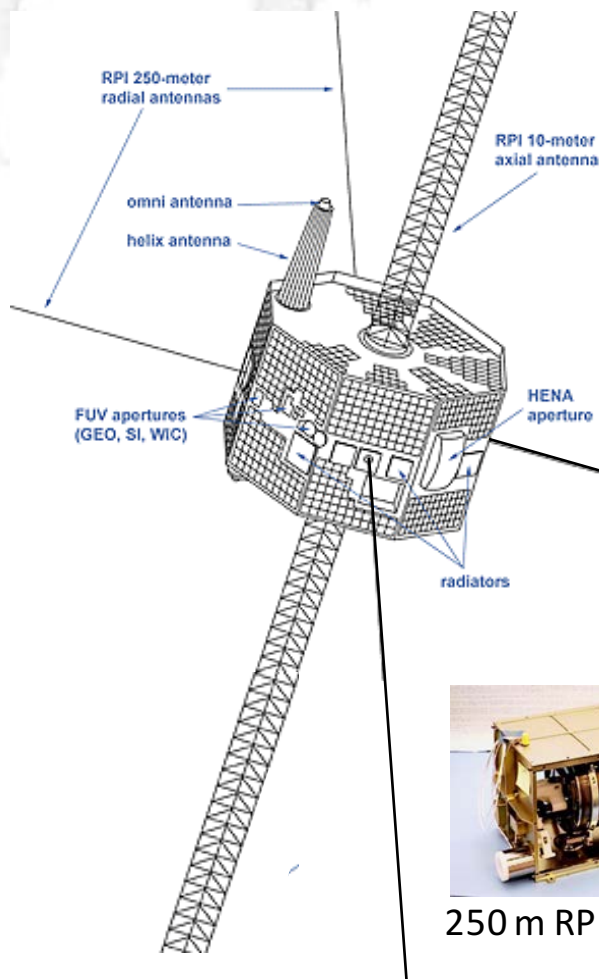
First phase of TID Experiments





VLF TRANSMISSION IN SPACE

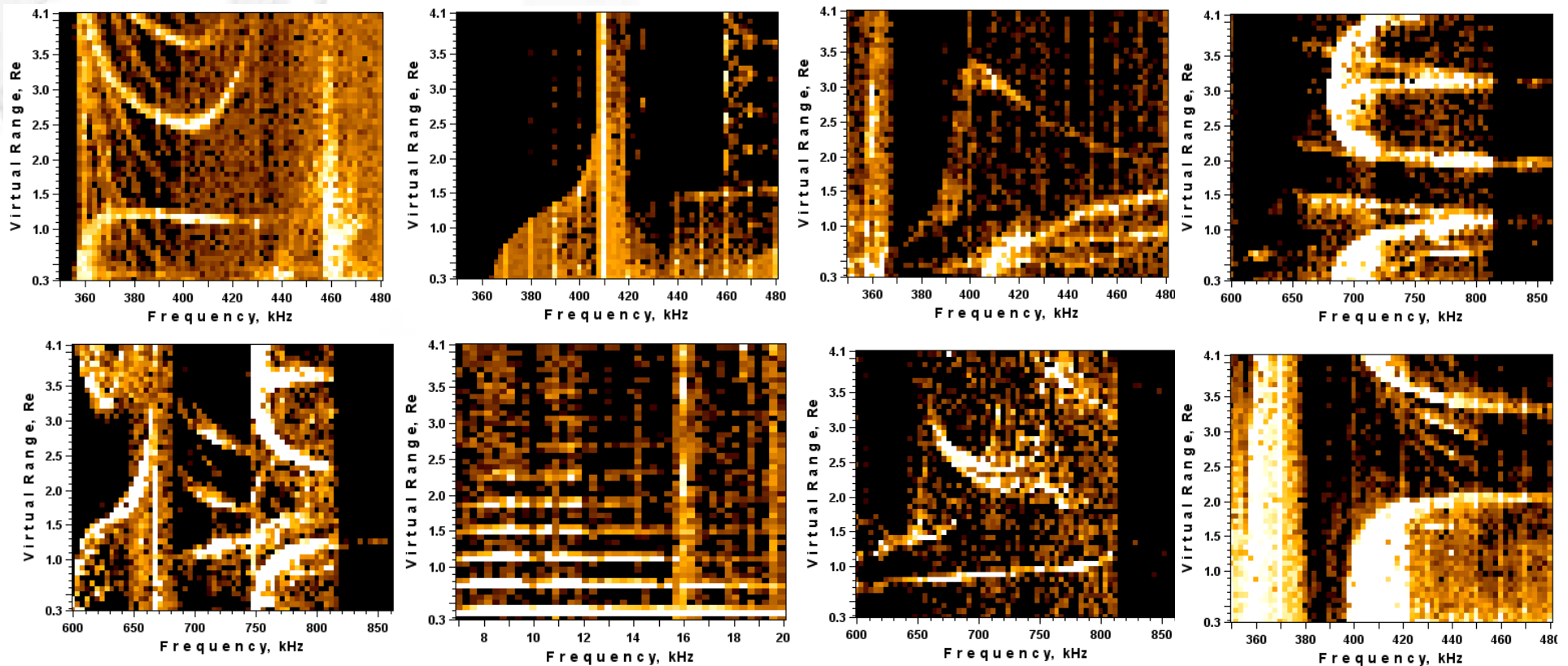
Radio Plasma Imager on IMAGE



250 m RPI Monopole

- **Low-power VLF to MF magnetospheric sounder**
- **Active mode:**
 - Remote electron density profiles from traces of echoes
 - Local plasma sensing from resonance signatures
- **Passive mode (spectrograms)**

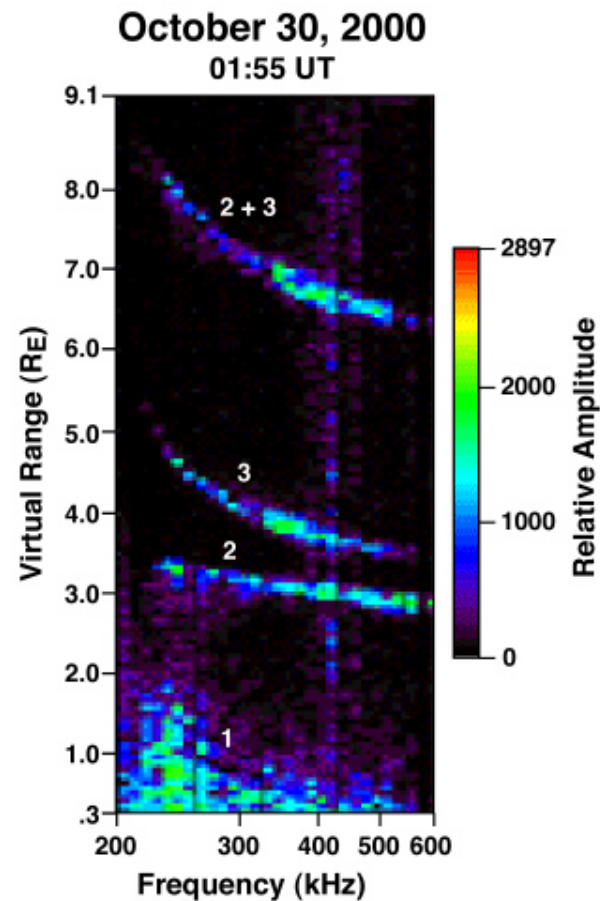
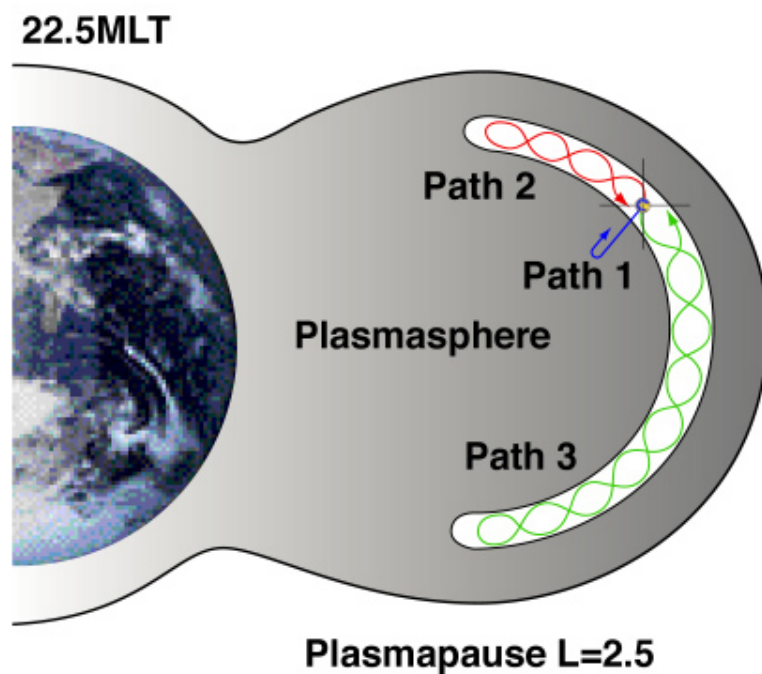
RPI Sounding Mode (Plasmagram)



Contain signatures of active sensing remote plasma regions

Recorded at opportune times and locations on orbit

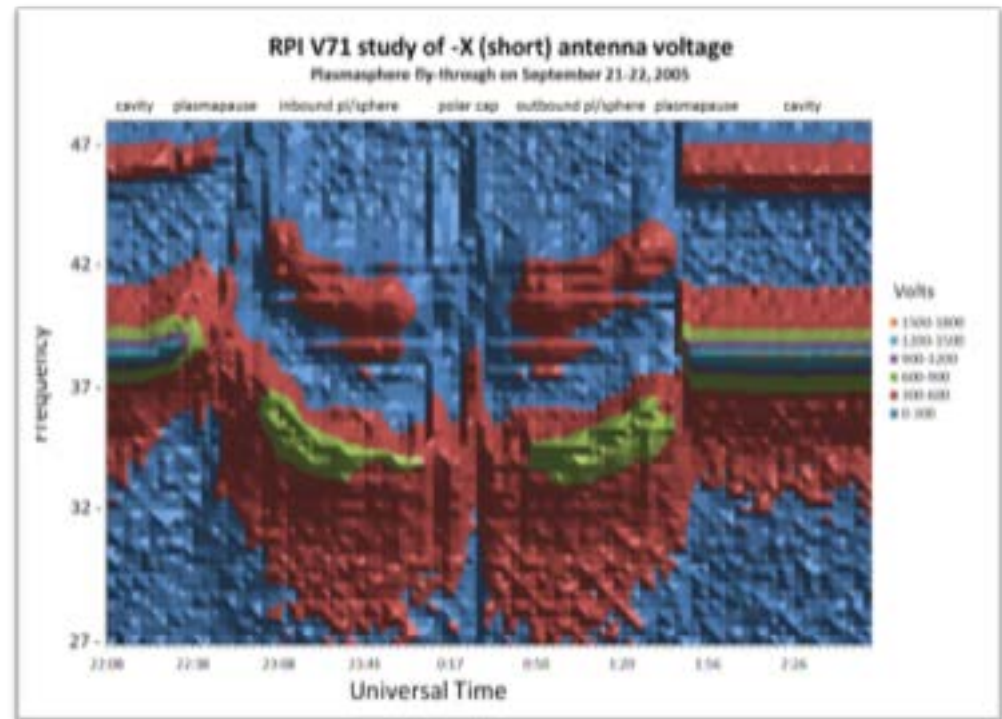
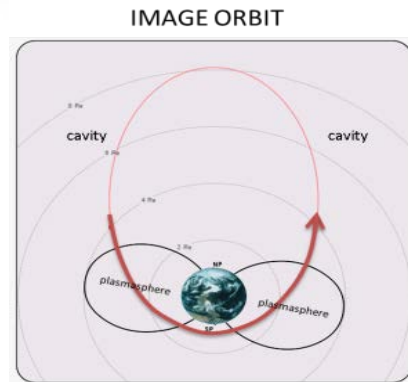
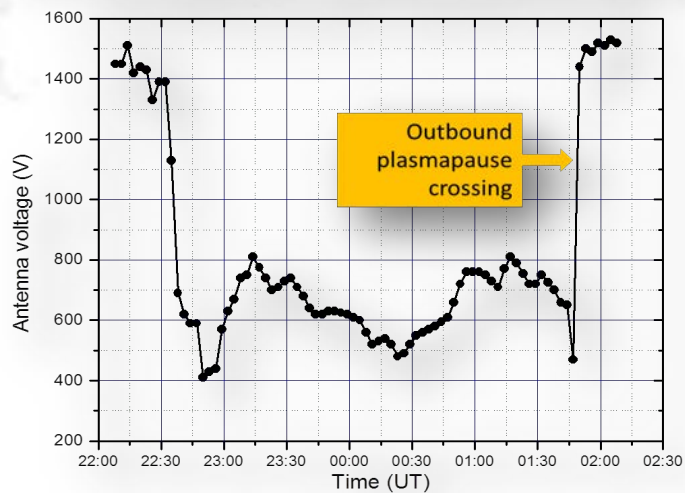
Guided/Field-aligned/Ducted propagation



VLF Transmission in Space Plasmas

Dynamics of RPI Antenna Voltage Resonance

RPI V71 study of -X (short) antenna voltage
Plasmasphere fly-through on September 21-22, 2005



Circuit resonance frequency varies
as RPI moves into denser plasma

VLF Transmission in Space Plasmas

RPI V71 study of X-axis power supply current

Plasmasphere fly-through on September 21-22, 2005

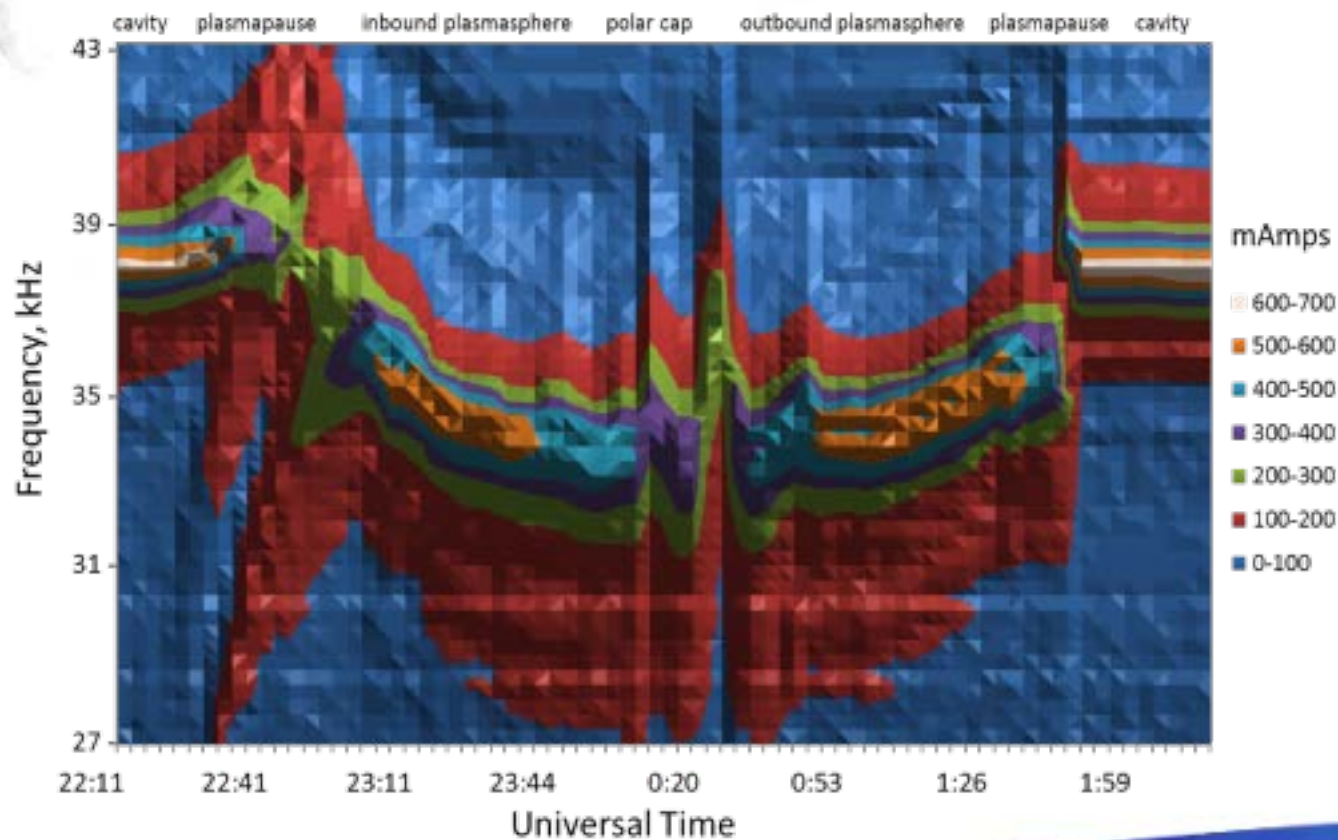
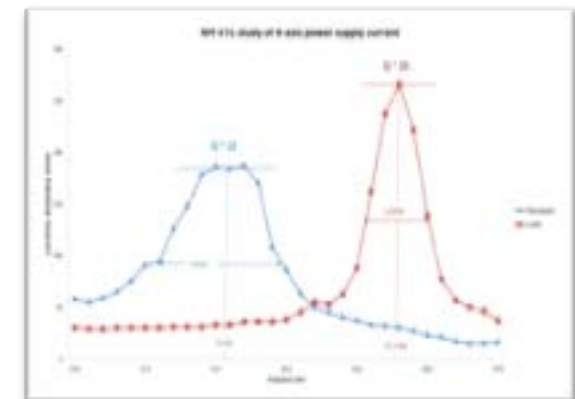
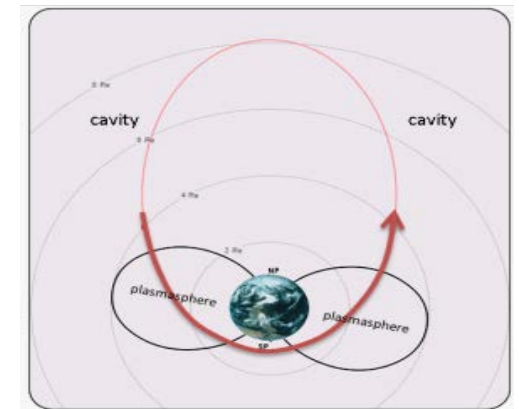


IMAGE ORBIT

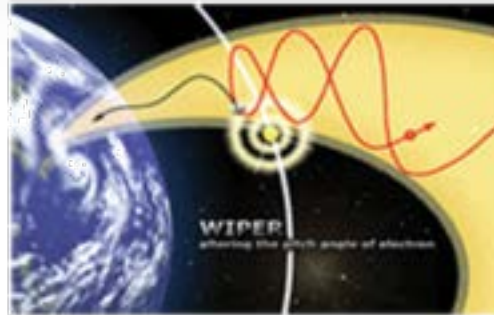
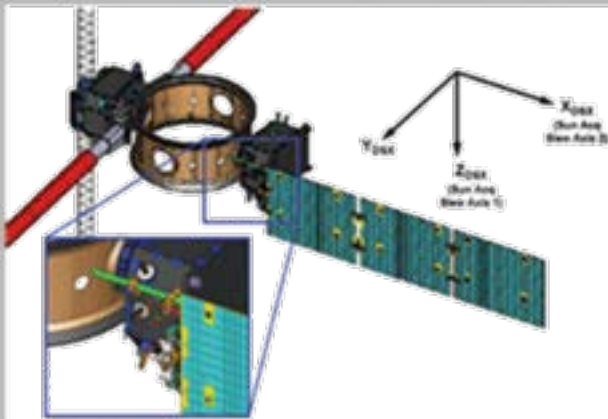


Adaptive Tuning of VLF Transmitter

DSX

Abstract—The Air Force Research Laboratory (AFRL) Space Vehicles Directorate has developed the Demonstration and Sciences Experiment (DSX) to research the technologies needed to deploy space assets in the harsh MEO radiation environment. The DSX comprises three basic research experiments:

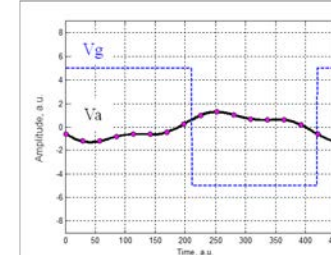
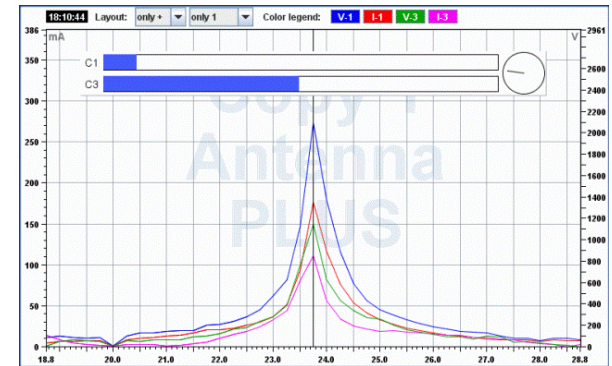
- 1) Wave Particle Interaction Experiment
- 2) Space Weather Experiment
- 3) Space Environmental Effects Experiment



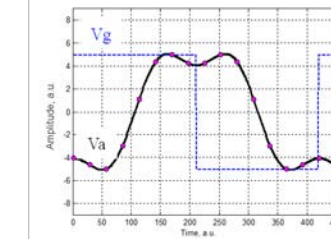
WPIx

The Wave Particle Interaction Experiment will explore the transmission, propagation, and amplification of very low frequency (VLF) waves in the magnetosphere. The experiment will be performed using a mission architecture that includes space transmitters and receivers on DSX, ground transmitters such as High frequency Active Auroral Research Program (HAARP) and VLF broadcast stations.

<http://ulcar.uml.edu/tntdsx.html>



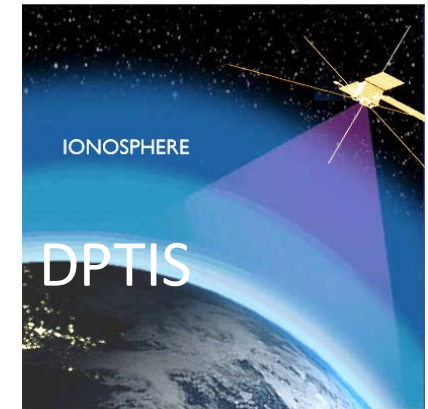
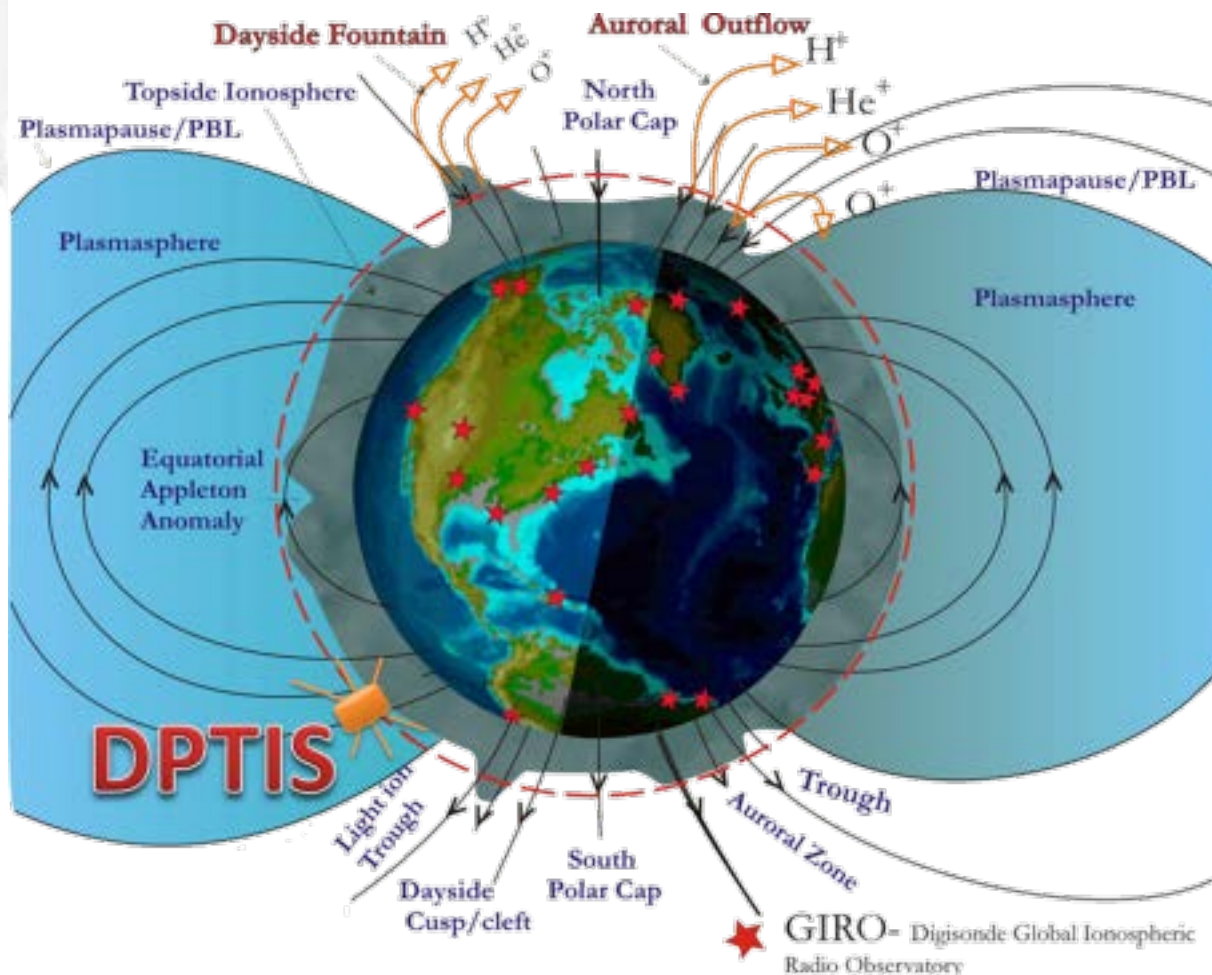
Out of tune



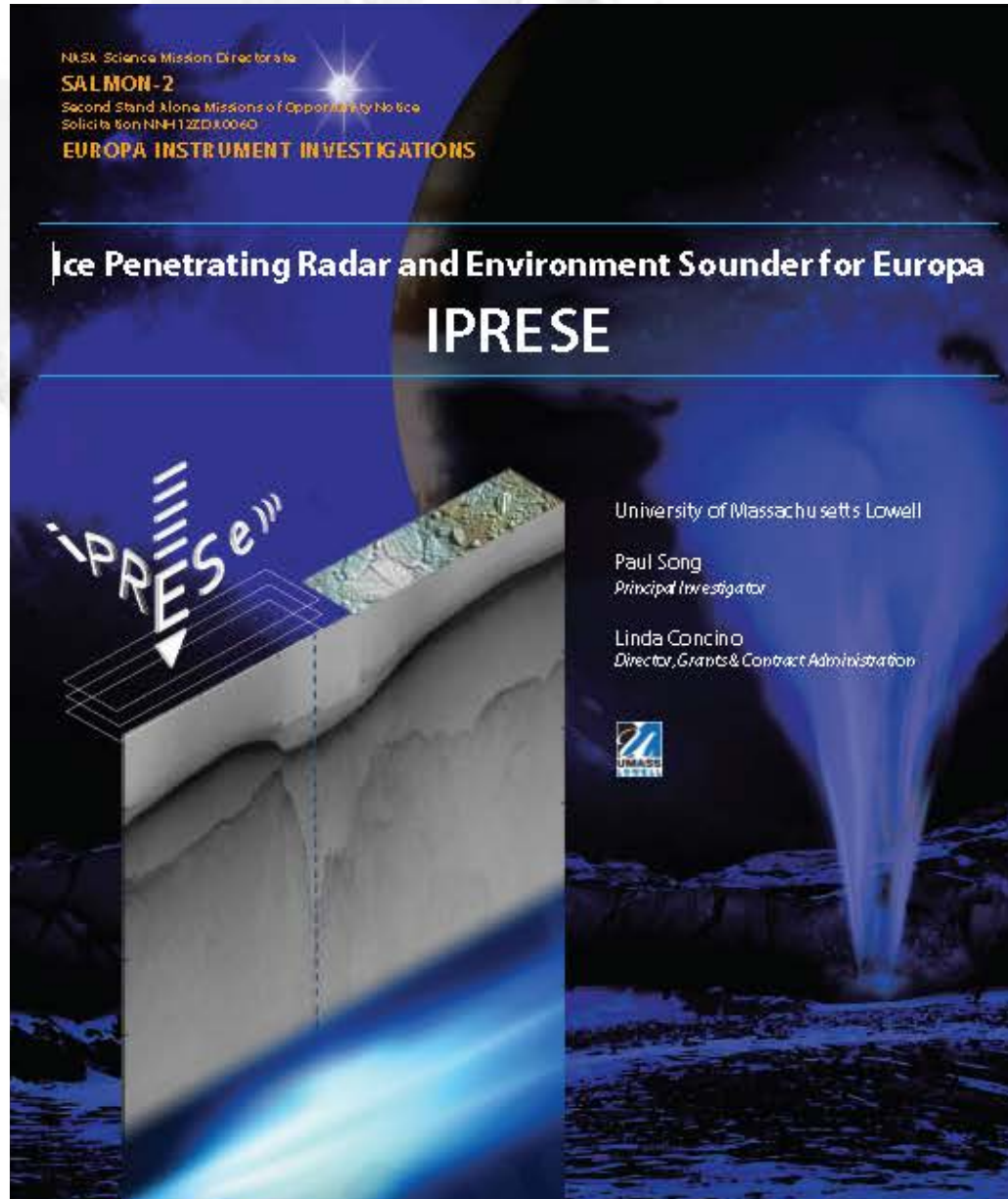
Tuned

Topside Ionospheric Sounder

800 – 1,000 km altitude



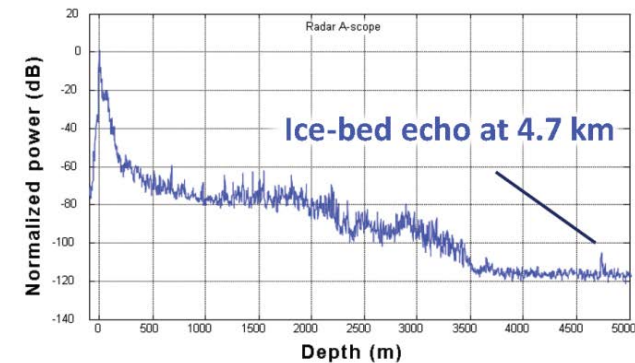
- 4-7 s ionograms
- Doppler skymaps for irregularities
- 0.5 s relaxation sounding sweep
- 3-axis antenna with Tx beamforming and Rx direction and polarization
- 3D Double probe E-field sensor



GPR

• Concept Instrument

- VHF 50 MHz Ice Penetrating Radar
 - Kansas U. Team
 - 120+ dB dynamic range
 - Chirp waveform



- MF-HF IPR/Sounder
 - UMass Lowell Team
 - Narrowband signal with quiet frequency search