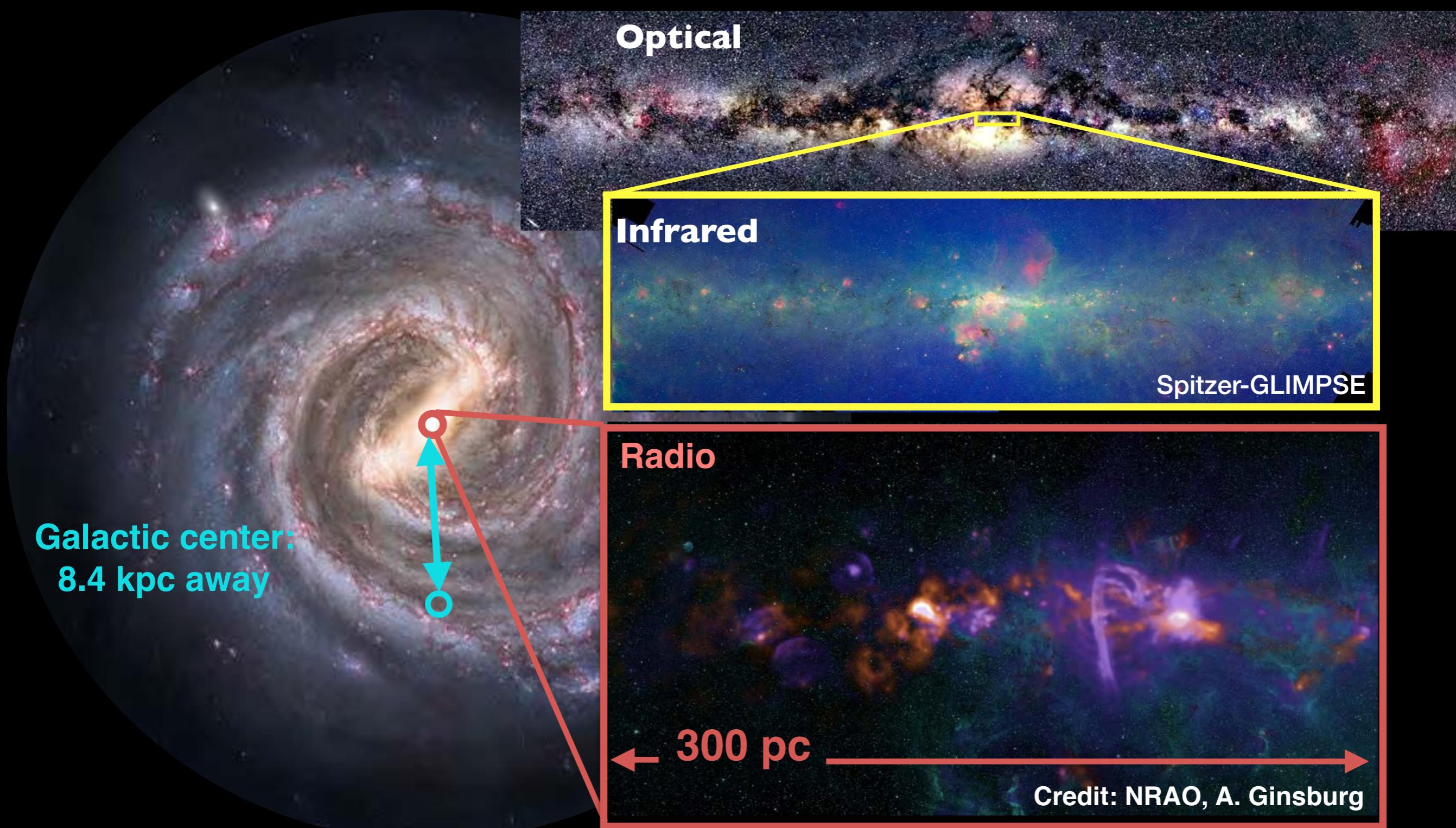


Apples to Apples: Comparing the Cores of Nearby Galaxies



Elisabeth A.C. Mills (BU),
Aditya Togi (UTSA), Diederik Kruijssen (U. Heidelberg),
Michael Kaufman (SJSU), Alberto Bolatto (Maryland), Adam Leroy (OSU),
David Meier (New Mexico Tech)

The Milky Way center: the nearest galaxy nucleus



Artist's conception of the Milky Way. Credit: Nick Risinger



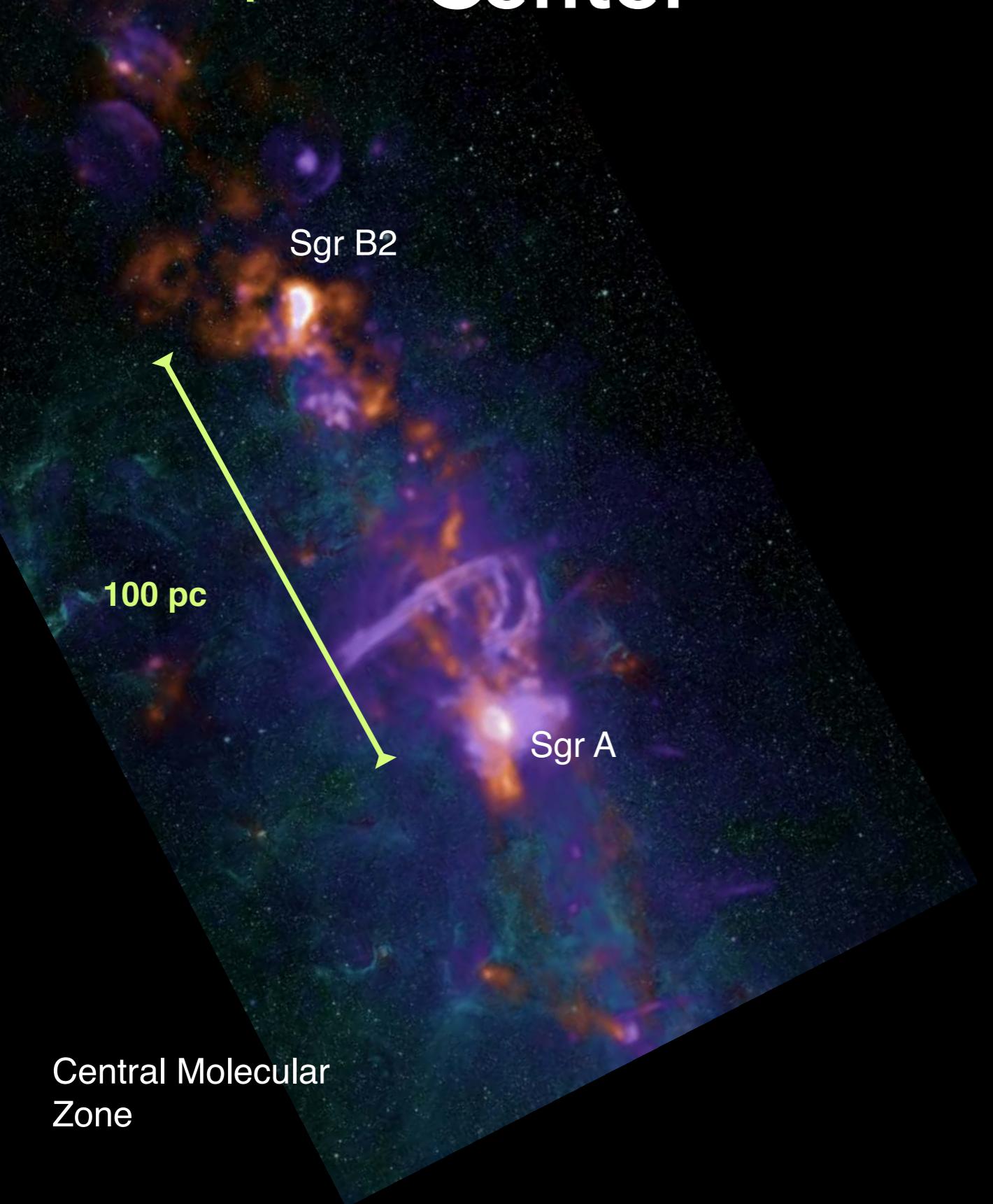
Millimeter : cold dust/gas

Infrared : hot dust, stars

Radio : hot gas, plasma

The Galactic Center

d=8.0 kpc



The Galactic Center

d=8.0 kpc

Sgr B2

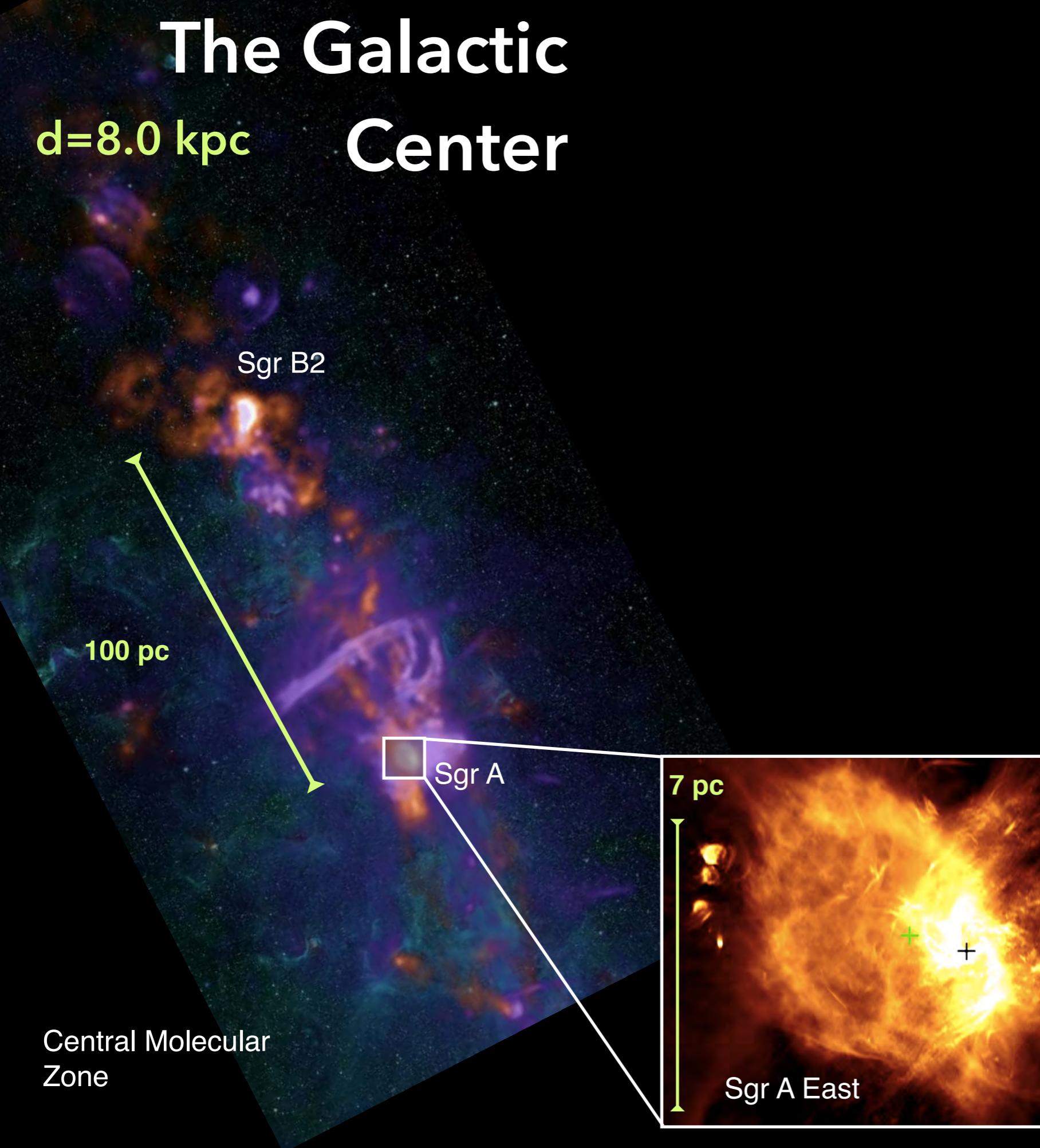
100 pc

Sgr A

Central Molecular Zone

7 pc

Sgr A East



The Galactic Center

$d=8.0 \text{ kpc}$

Sgr B2

100 pc

Sgr A

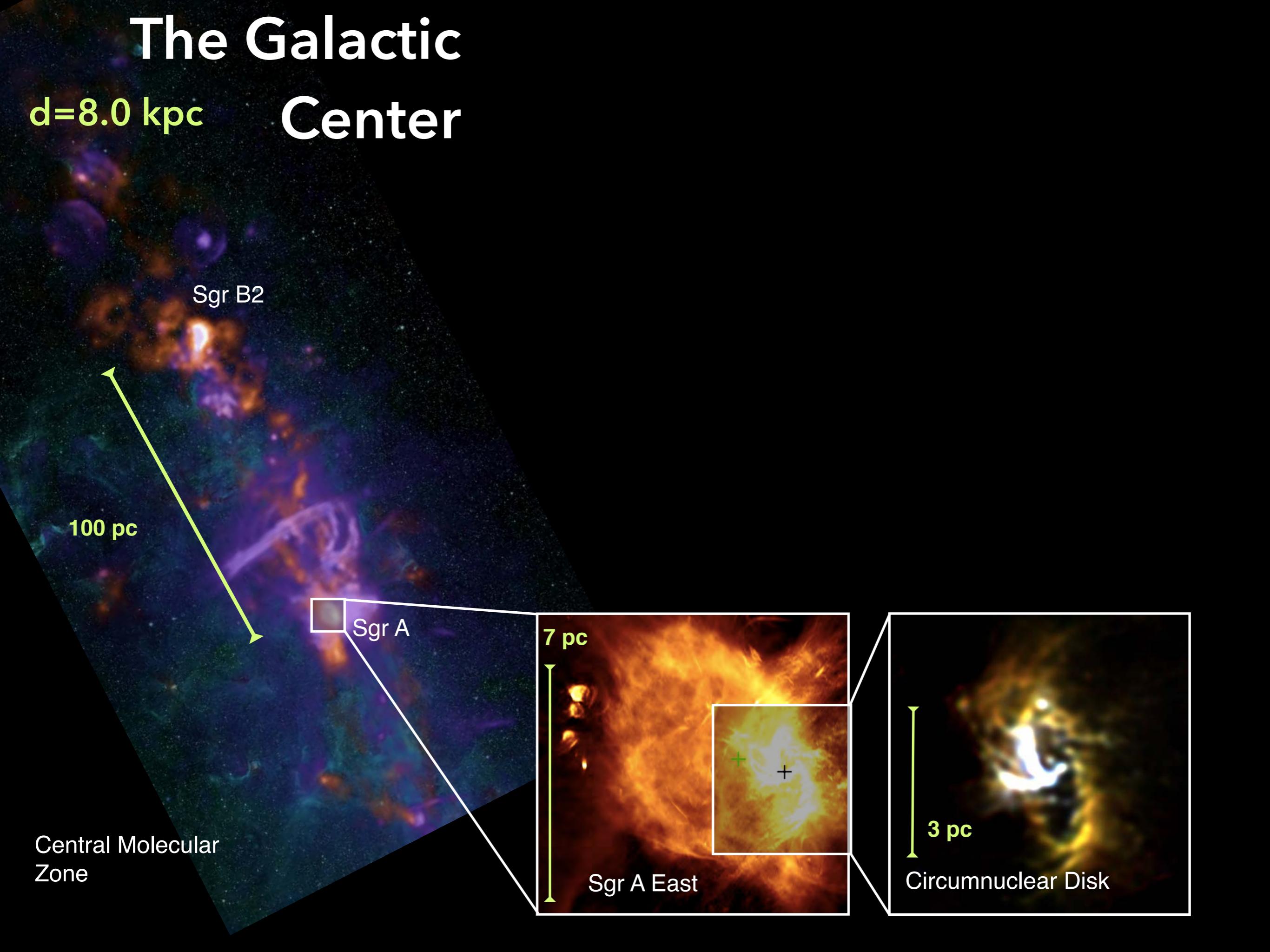
Central Molecular Zone

7 pc

Sgr A East

3 pc

Circumnuclear Disk



The Galactic Center

$d=8.0 \text{ kpc}$

Sgr B2

100 pc

Sgr A

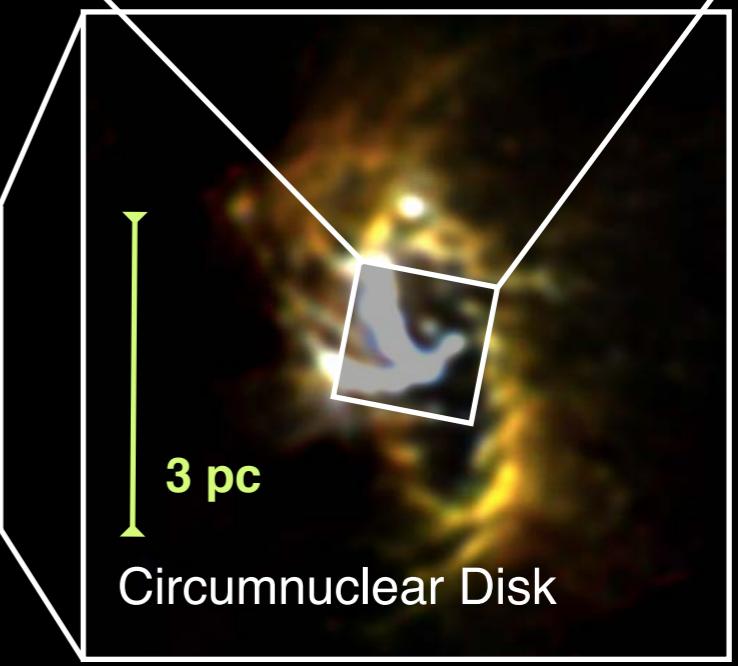
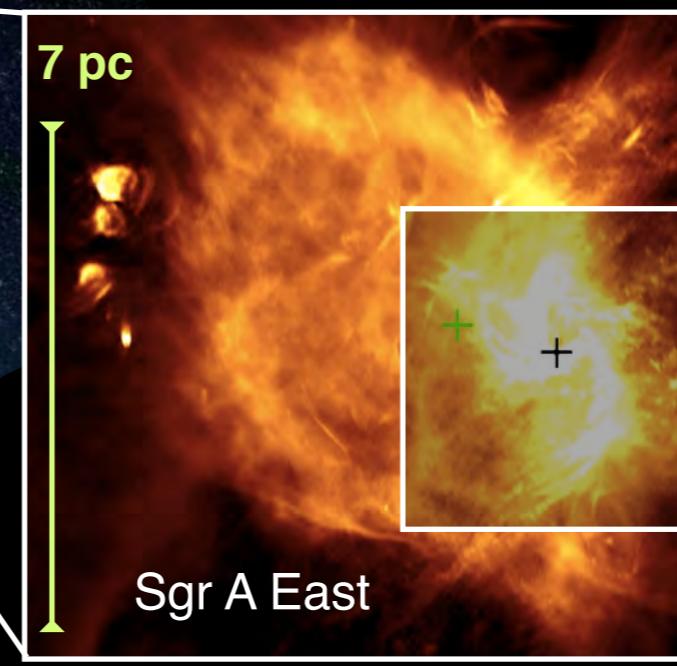
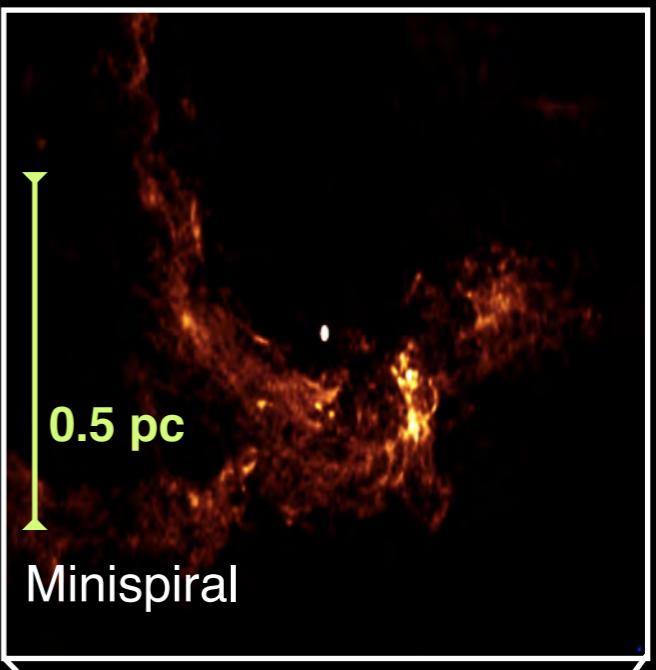
Central Molecular Zone

7 pc

Sgr A East

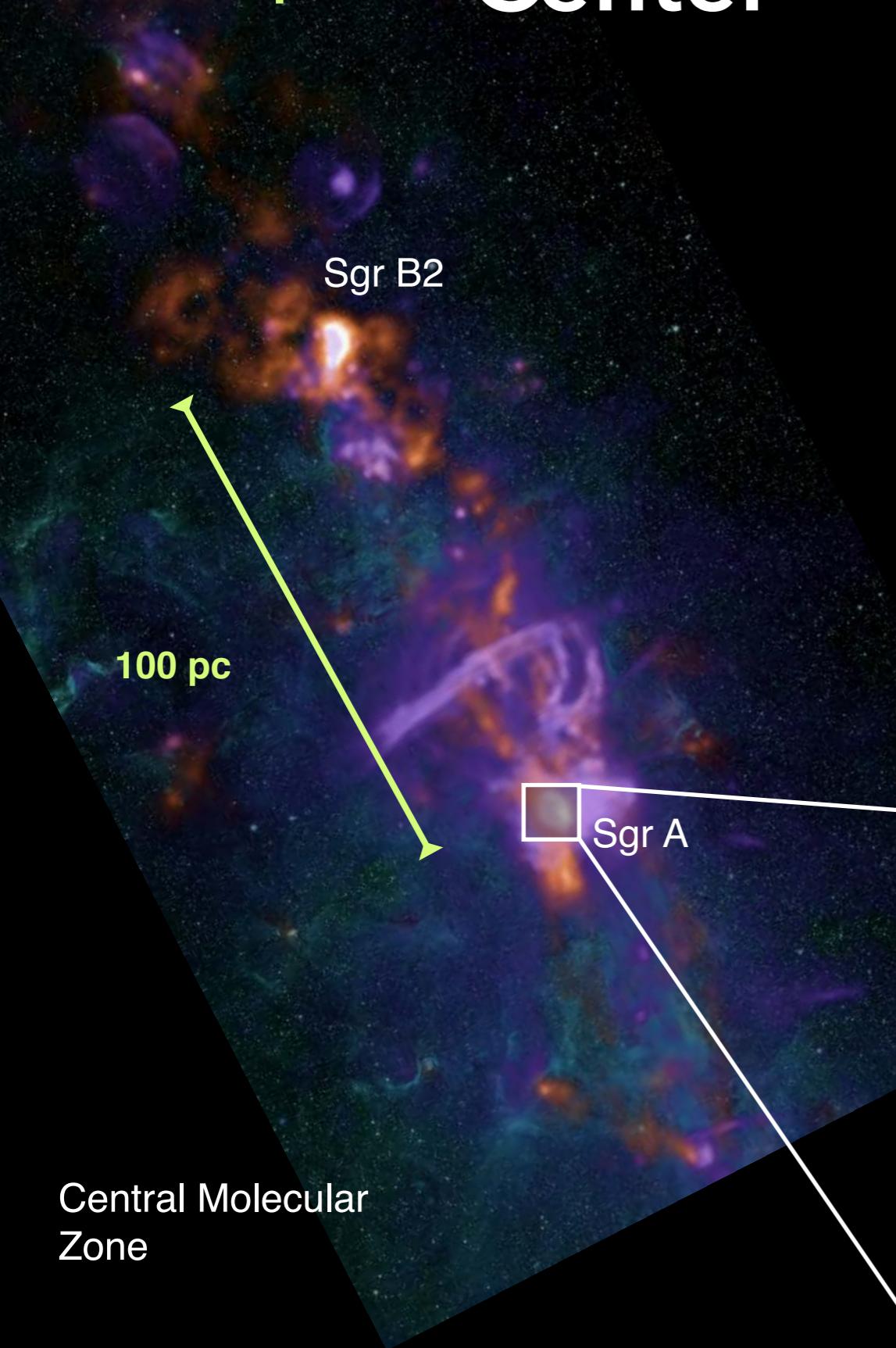
3 pc

Circumnuclear Disk

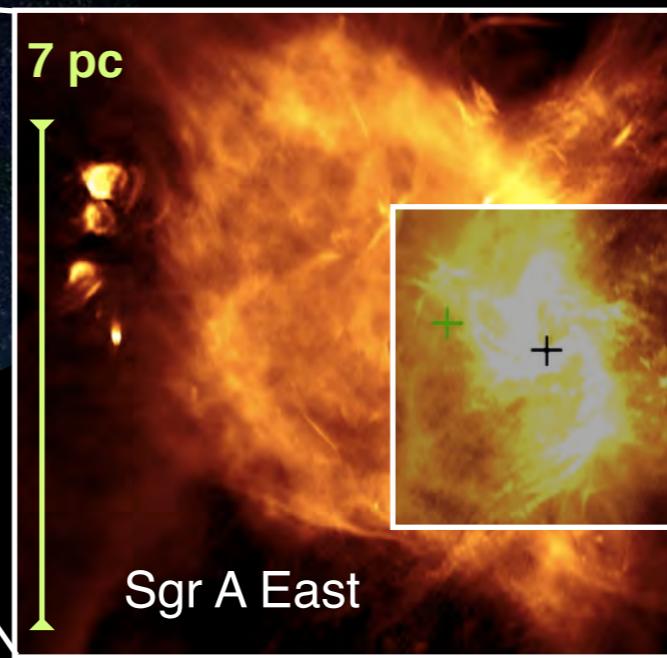


The Galactic Center

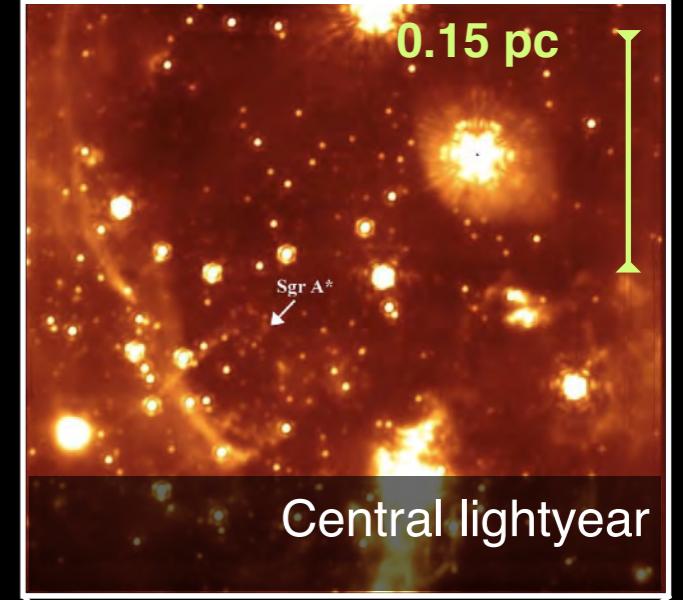
$d=8.0 \text{ kpc}$



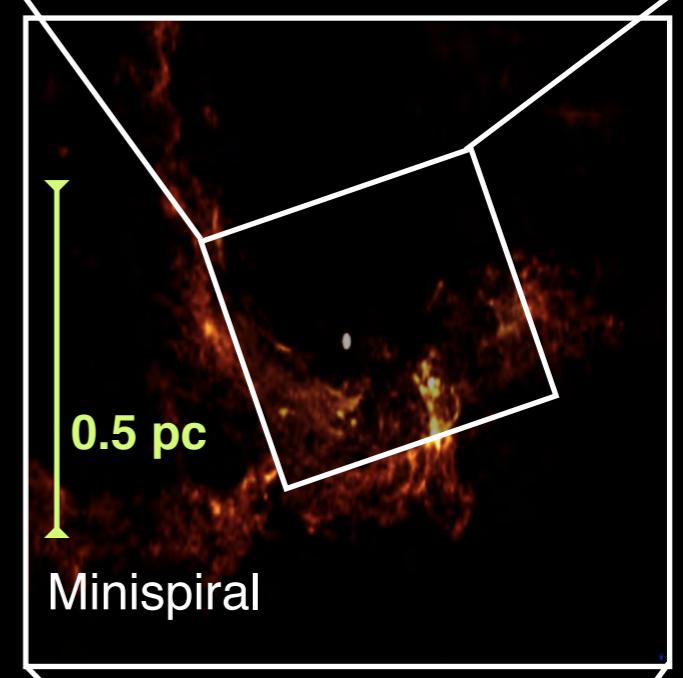
Central Molecular Zone



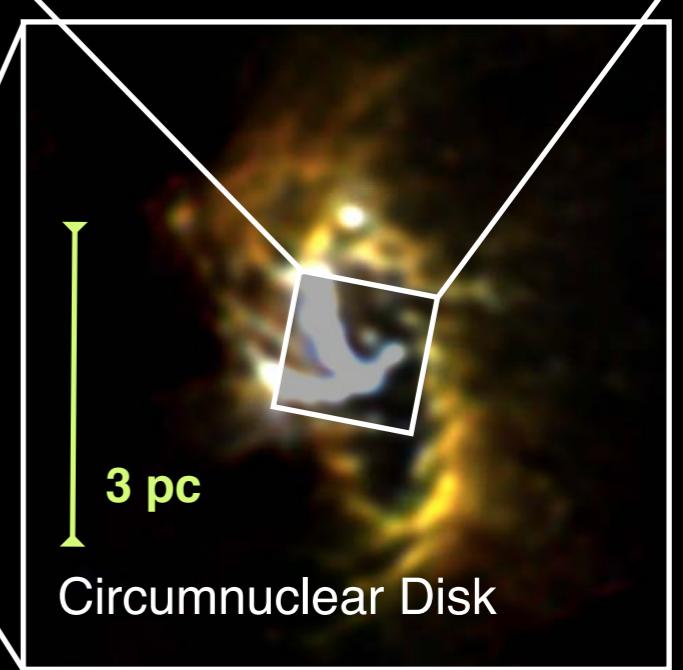
Sgr A East



Central lightyear



Minispiral



Circumnuclear Disk

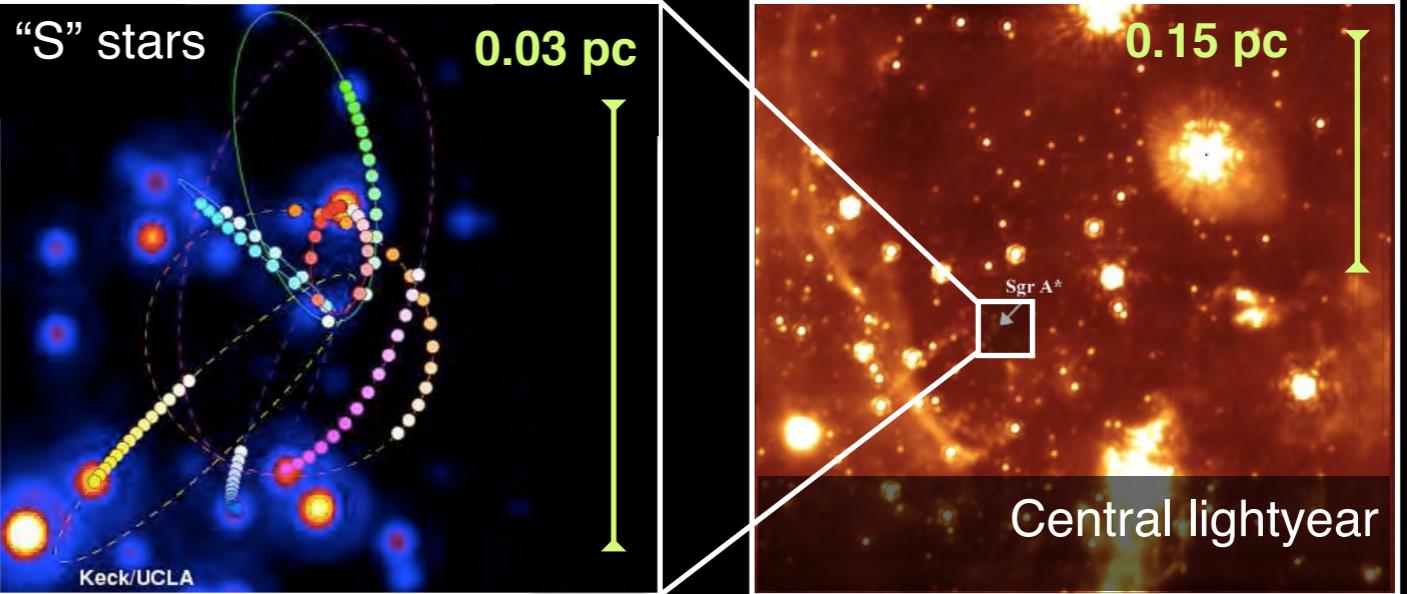
The Galactic Center

$d=8.0 \text{ kpc}$

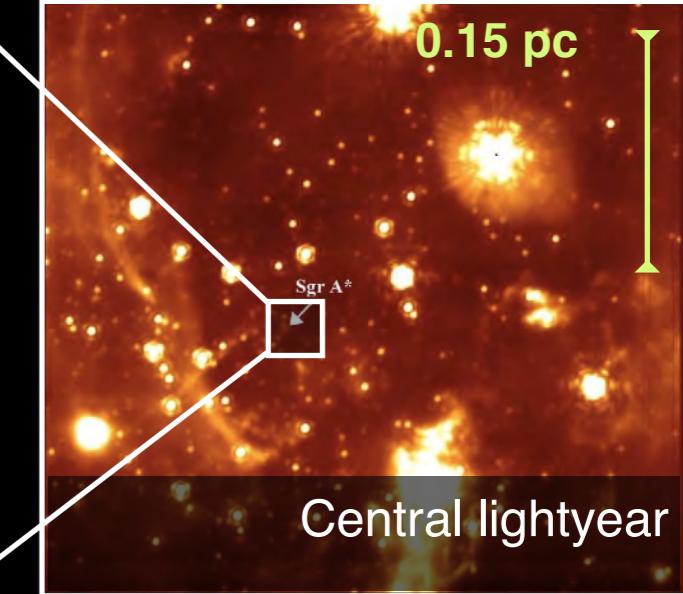
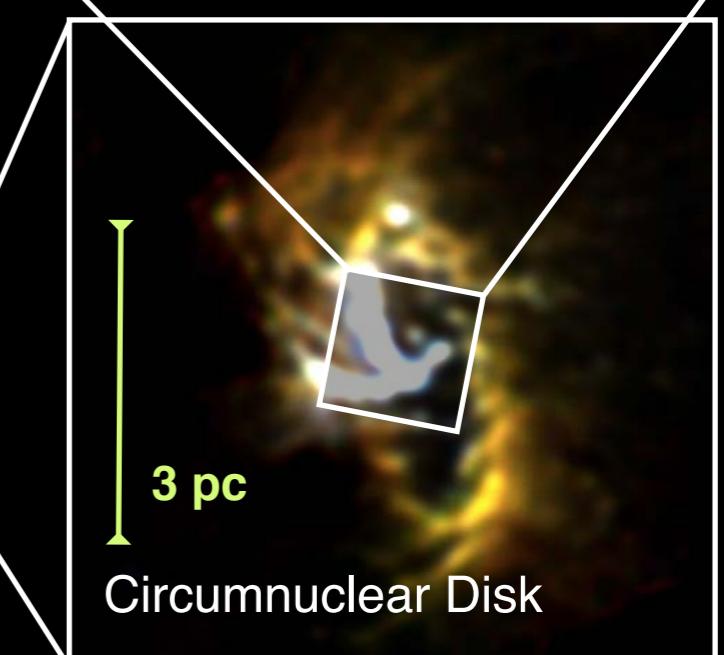
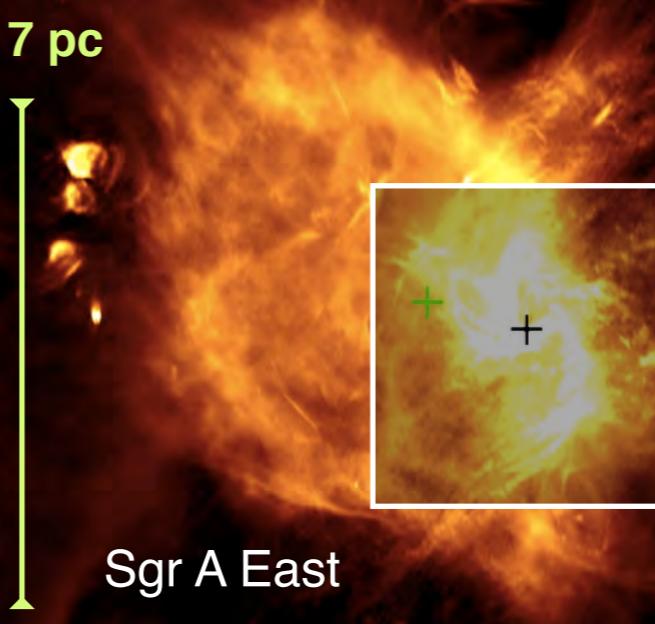
Sgr B2

100 pc

Central Molecular Zone

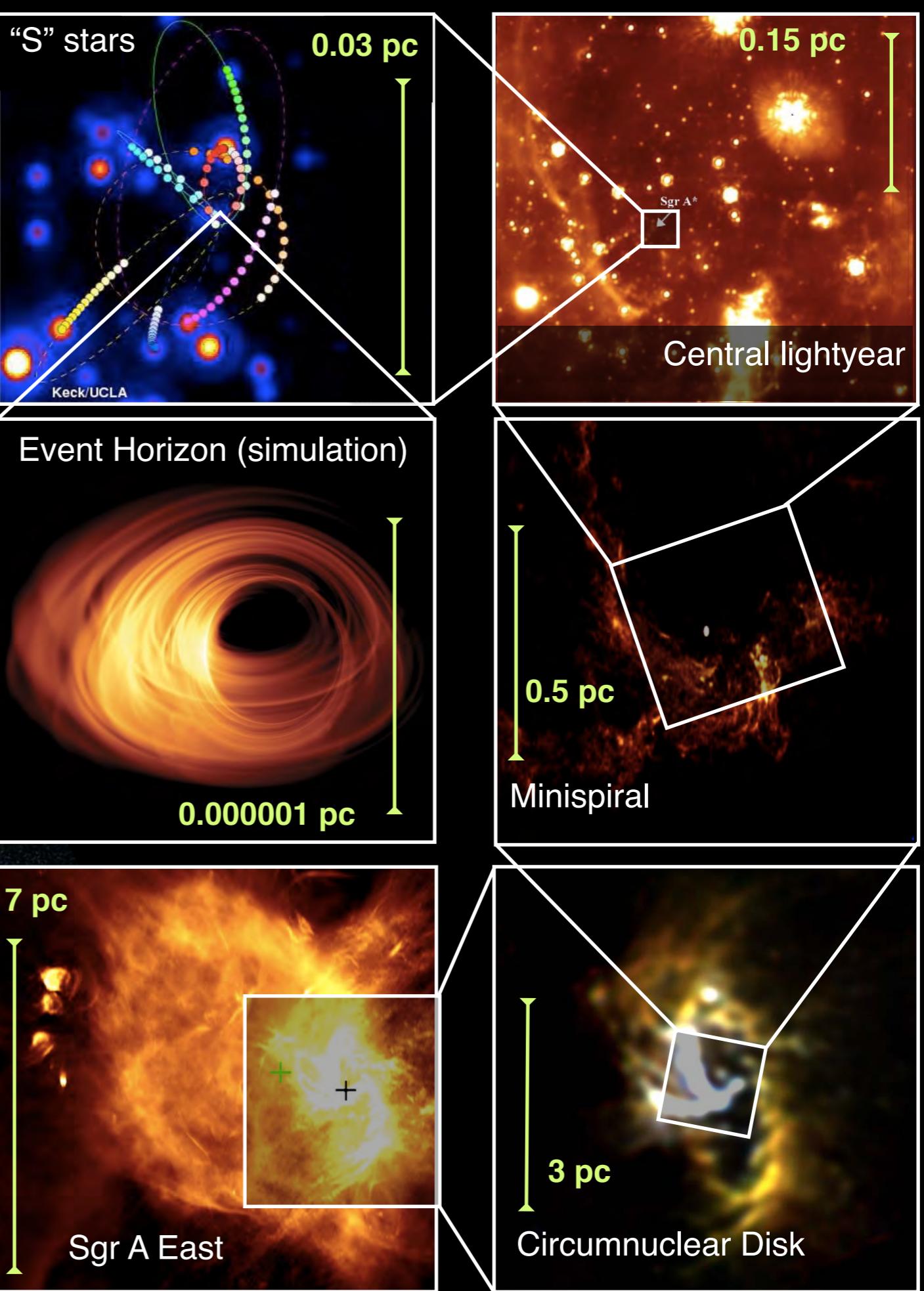
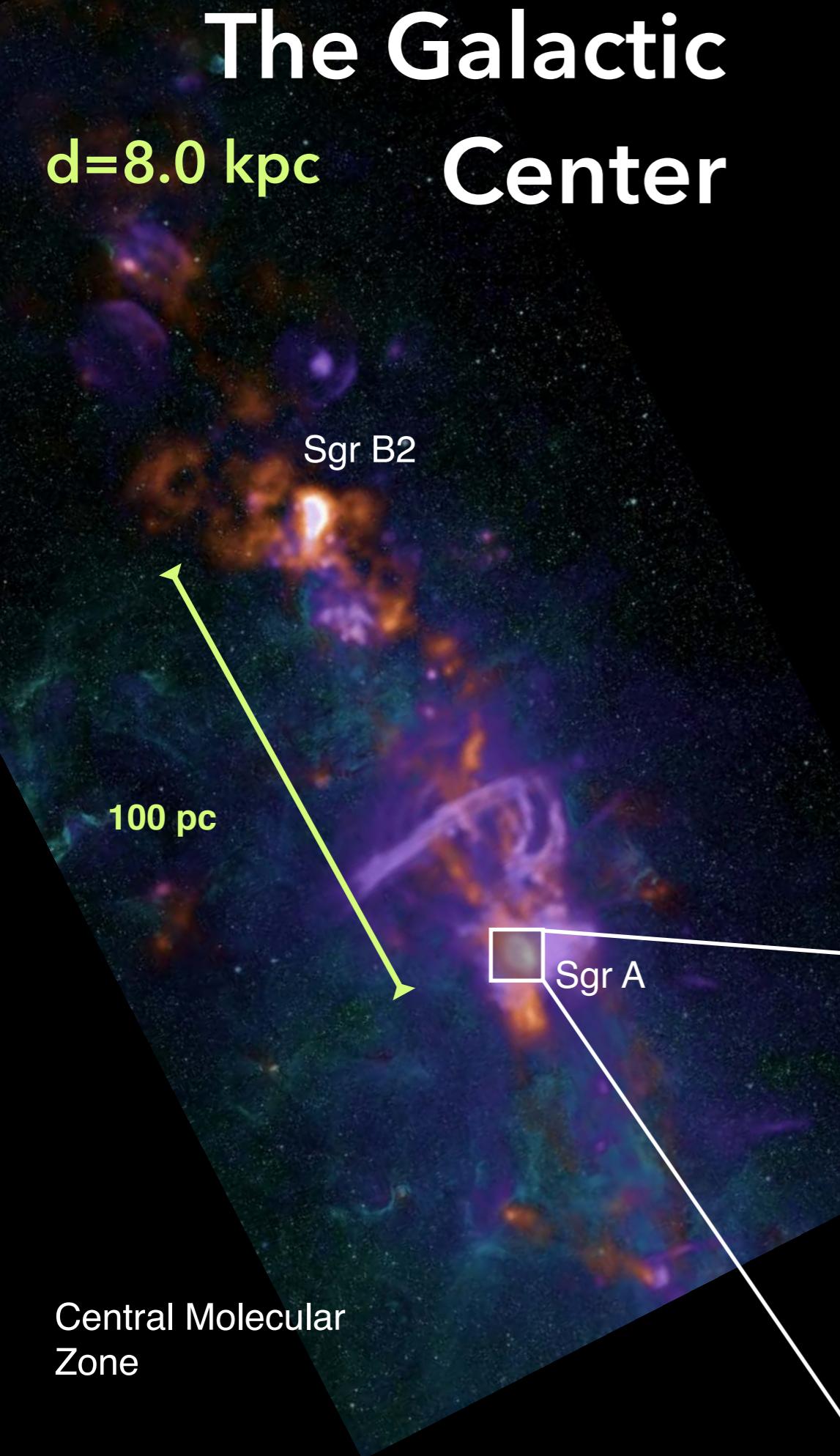


Sgr A



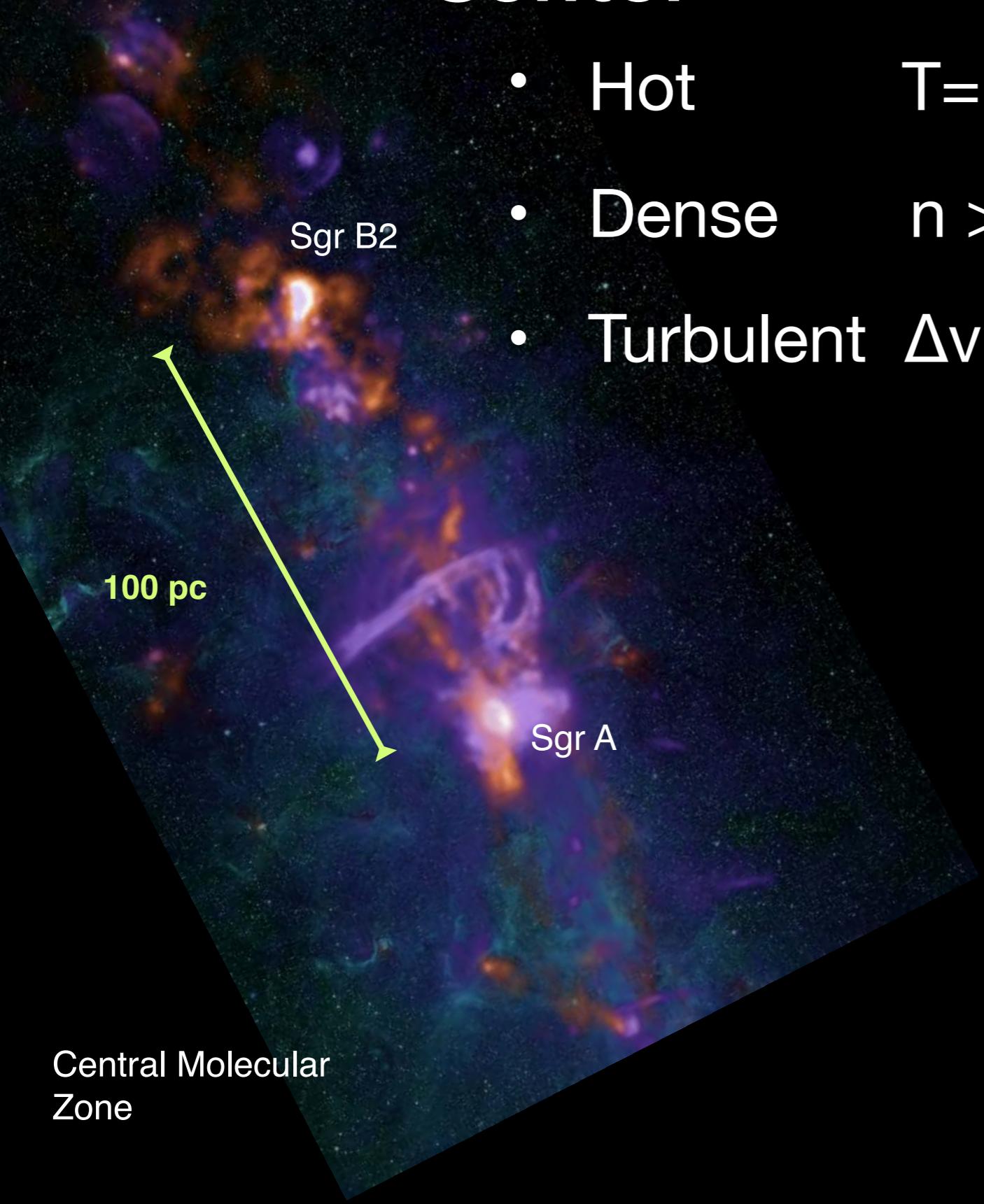
The Galactic Center

$d=8.0 \text{ kpc}$



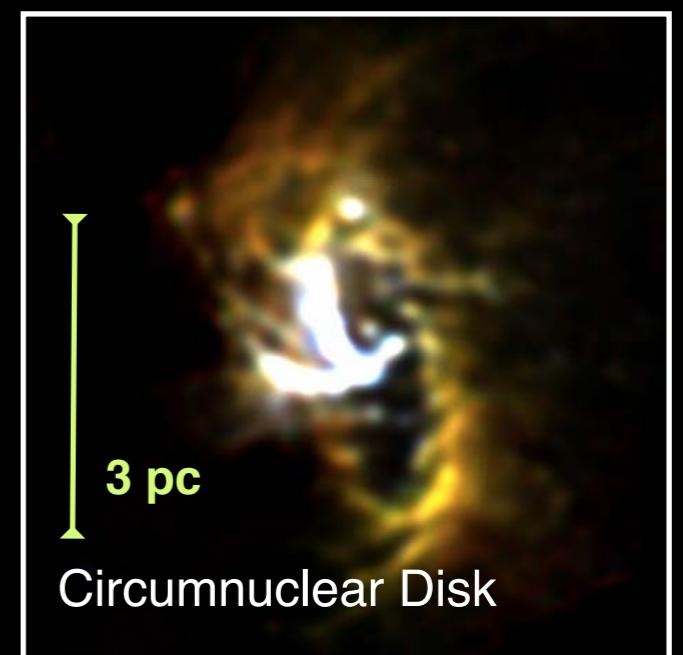
The Galactic Center

d=8.0 kpc

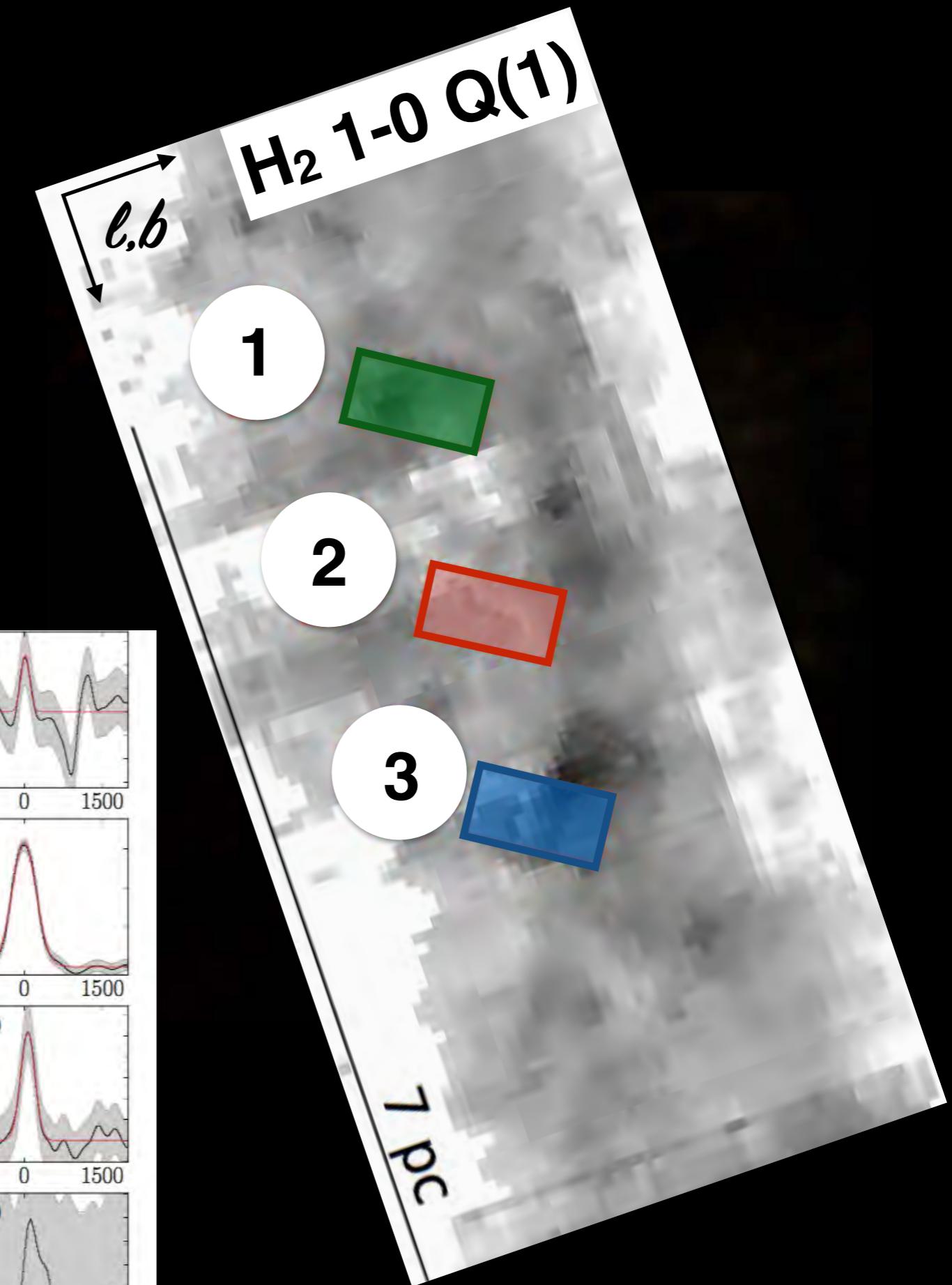
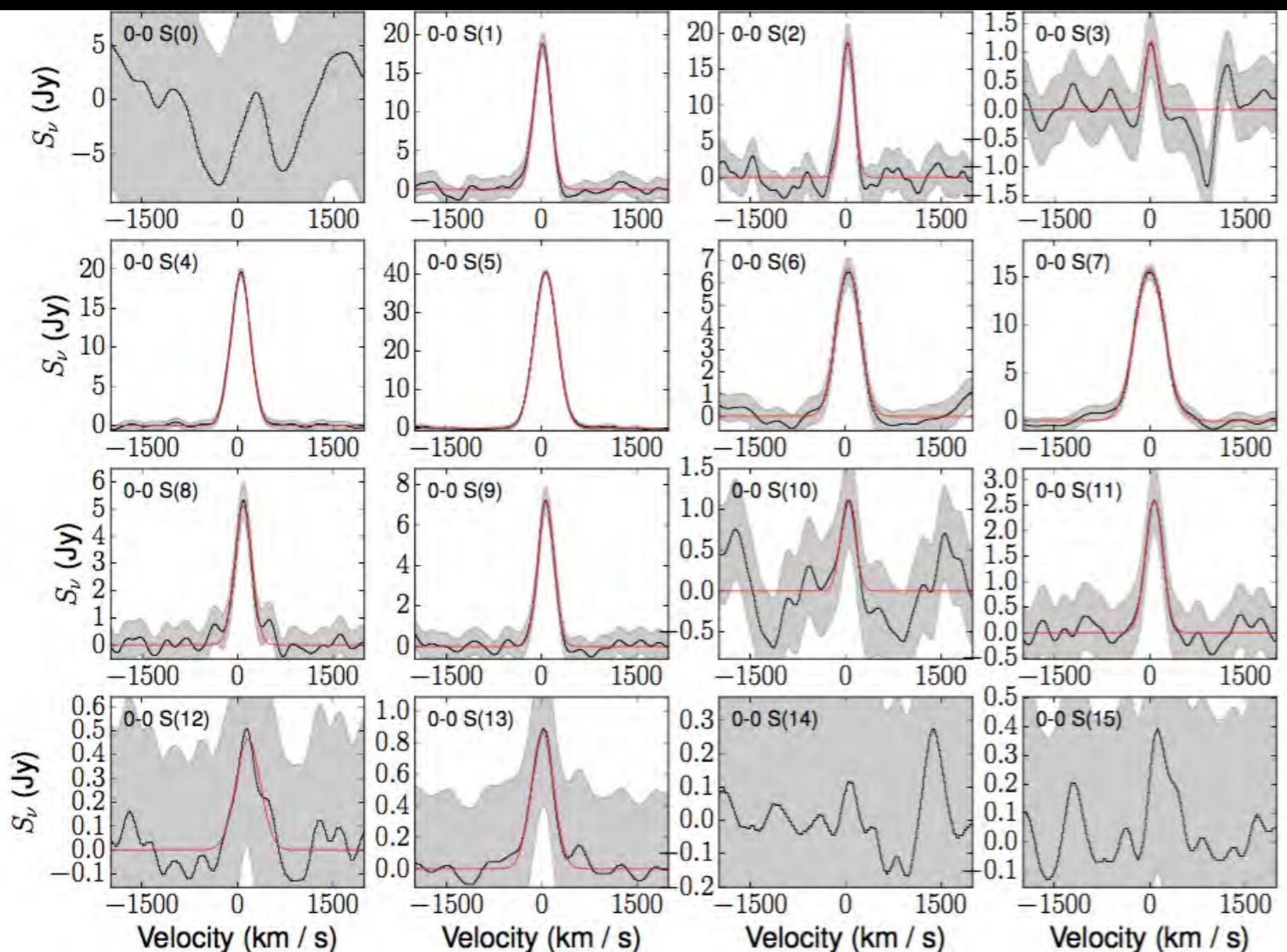


- Hot $T = 50 - 150 \text{ K}$ (Guesten et al. 1985, Huettemeister et al. 1993)
- Dense $n > 10^4 \text{ cm}^{-3}$ (Guesten et al. 1982)
- Turbulent $\Delta v \sim 15-50 \text{ km s}^{-1}$ (Bally et al. 1987)

**What about gas closer
to the black hole?**



Analyze archival
ISO spectra of H₂
toward 3 positions
in the CND to
measure
temperatures

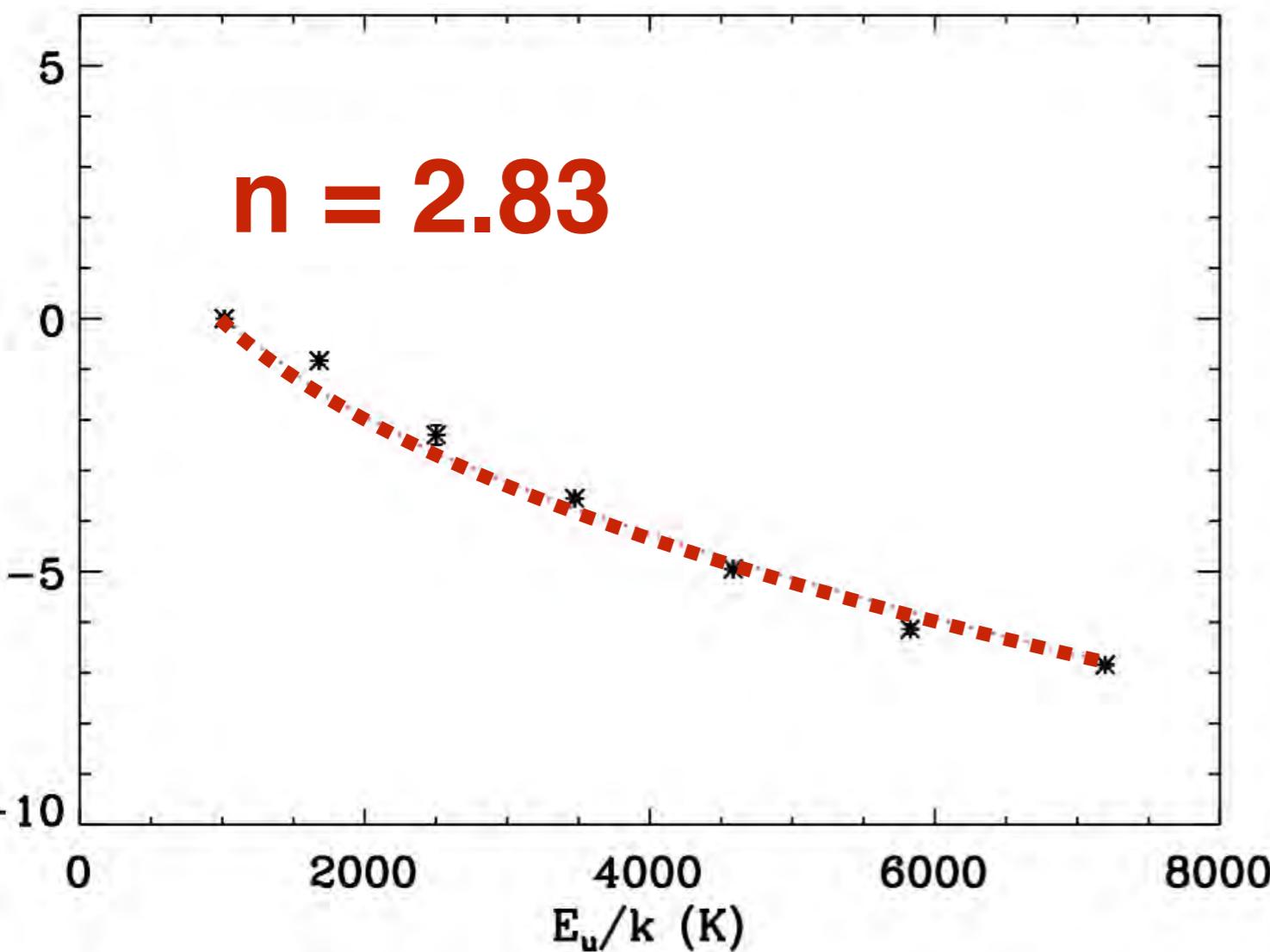


Mills, Togi & Kaufman 2017

Fit a power-law distribution of temperatures

$$dN \propto T^{-n} dT$$

Region 3 (Southwest region)



Other Galactic center clouds:

$n = 4.7 - 5.0$

Gas within a parsec of the supermassive black hole is significantly hotter than gas > 50 pc from the center

NGC 253

Barred Spiral

Total mass: 10^{11} M_{sun}

Black hole mass: 5×10^6 M_{sun}

Nuclear Starburst

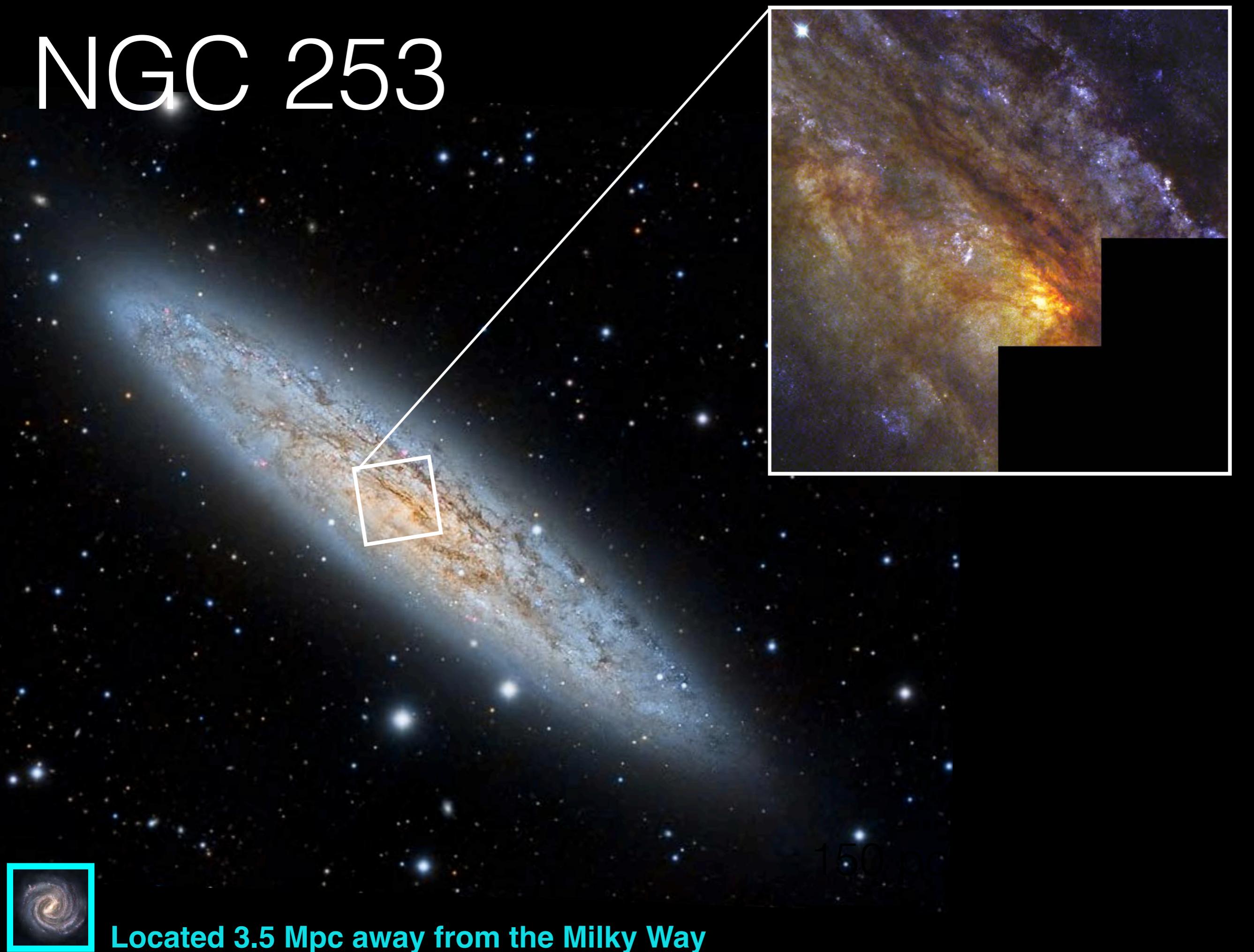
Star Formation Rate: 2.8 M_{sun}/yr

150 pc



Located 3.5 Mpc away from the Milky Way

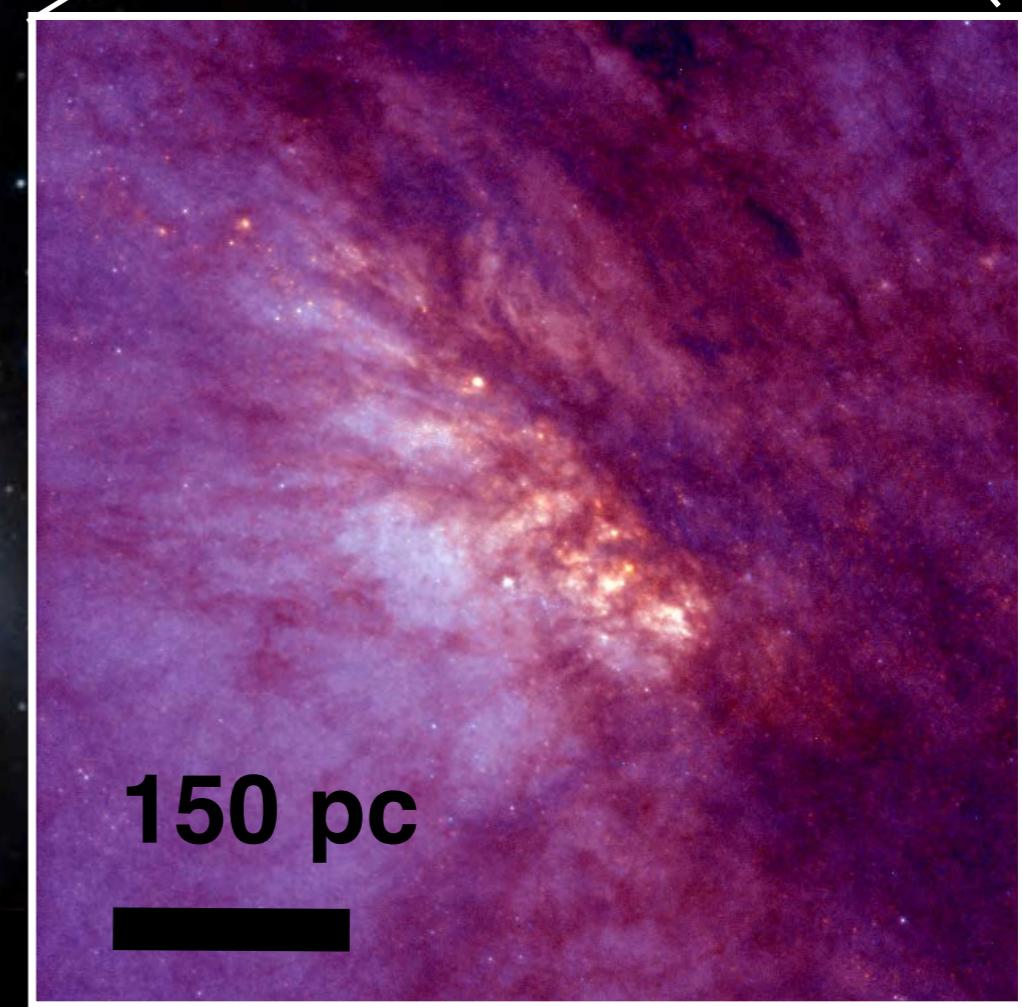
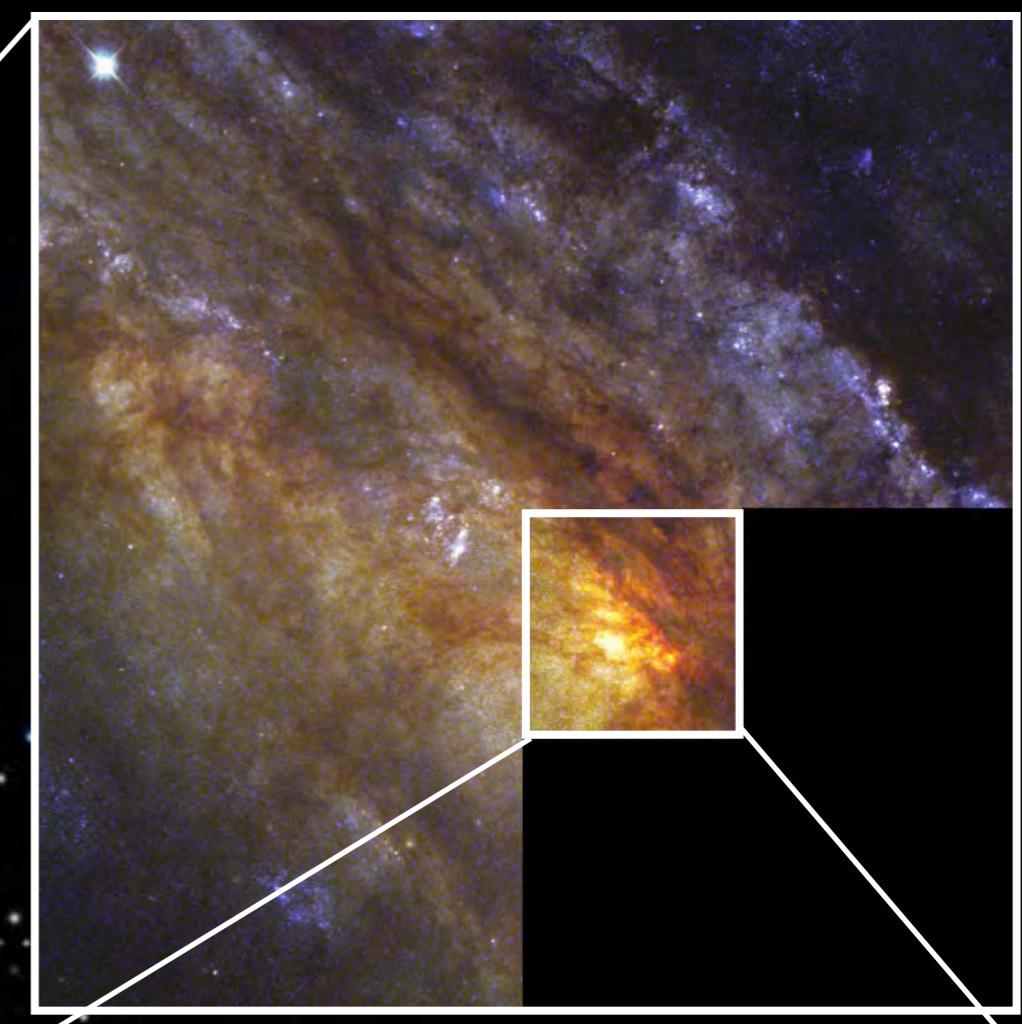
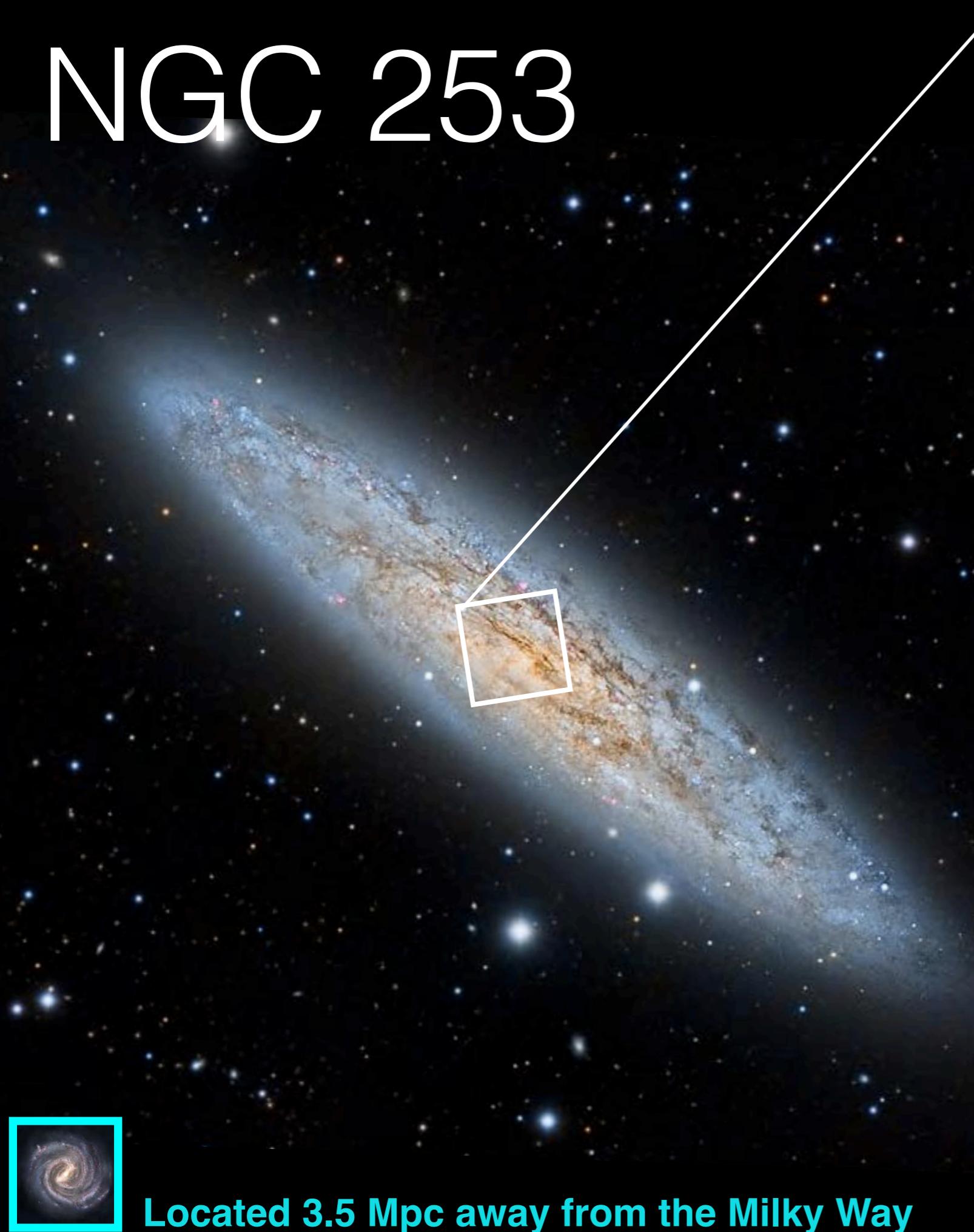
NGC 253



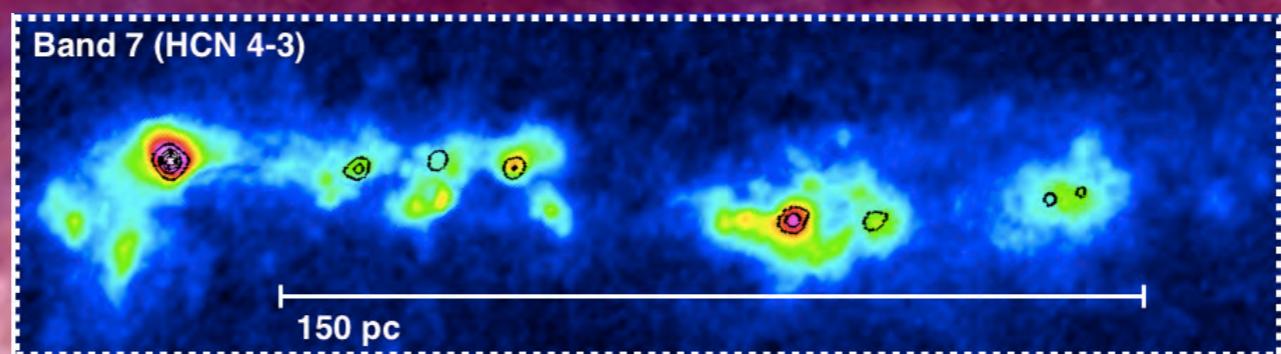
Located 3.5 Mpc away from the Milky Way

150 pc

NGC 253

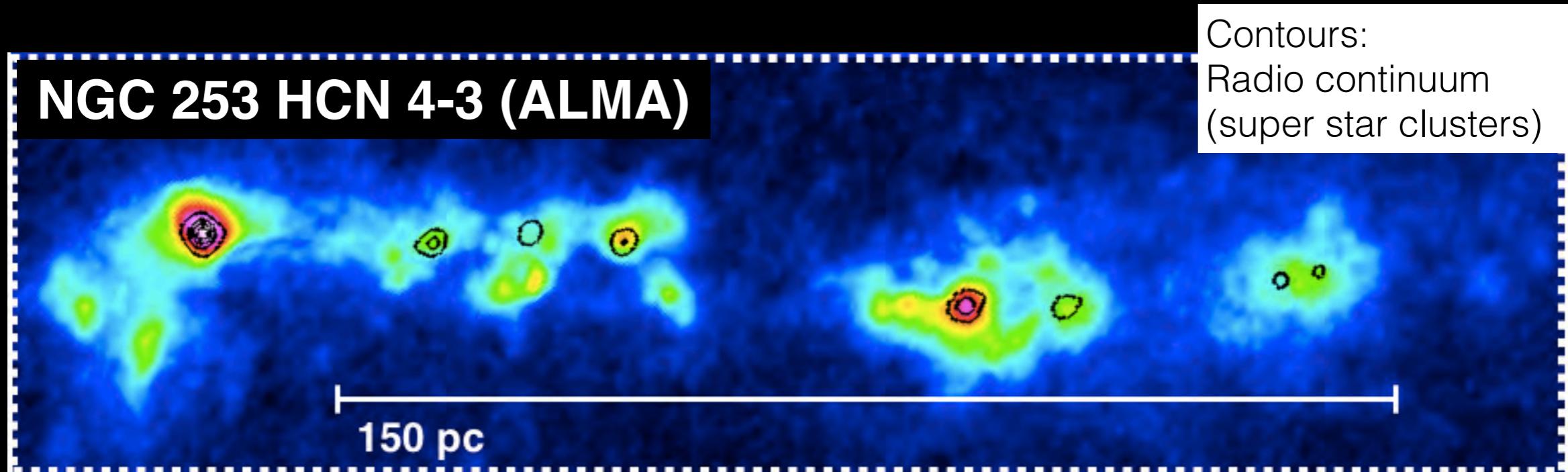


Located 3.5 Mpc away from the Milky Way

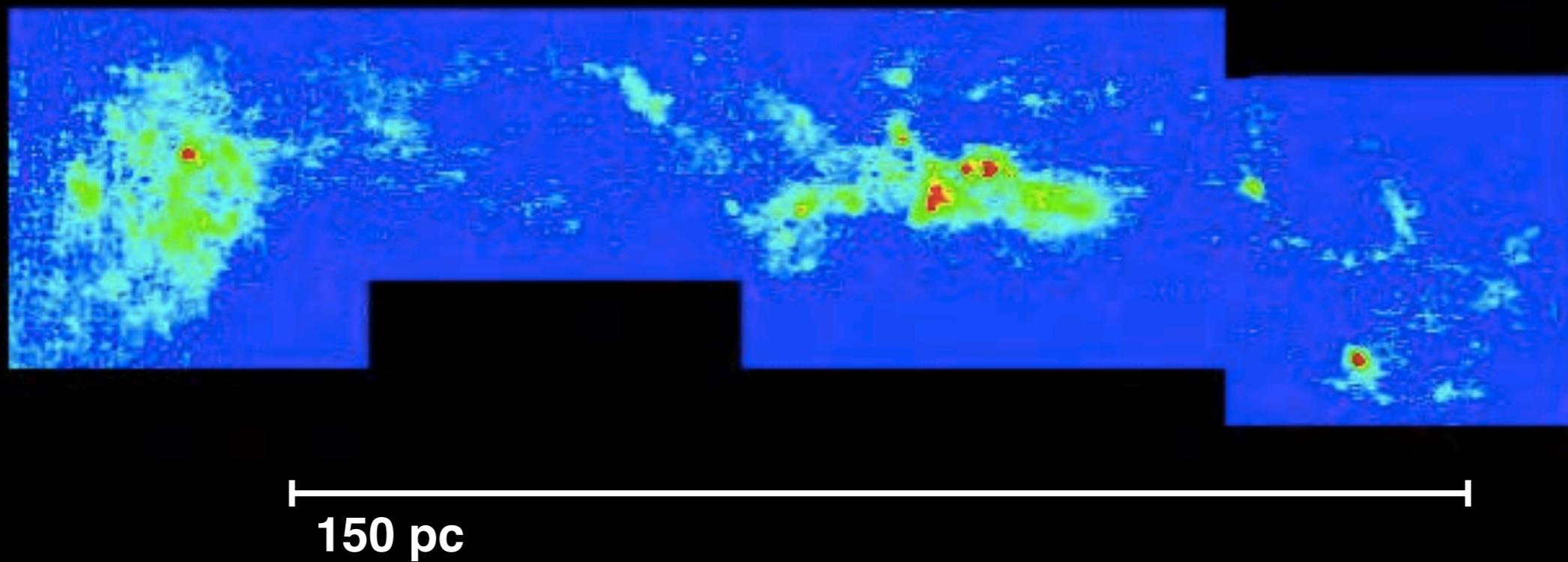


150 pc

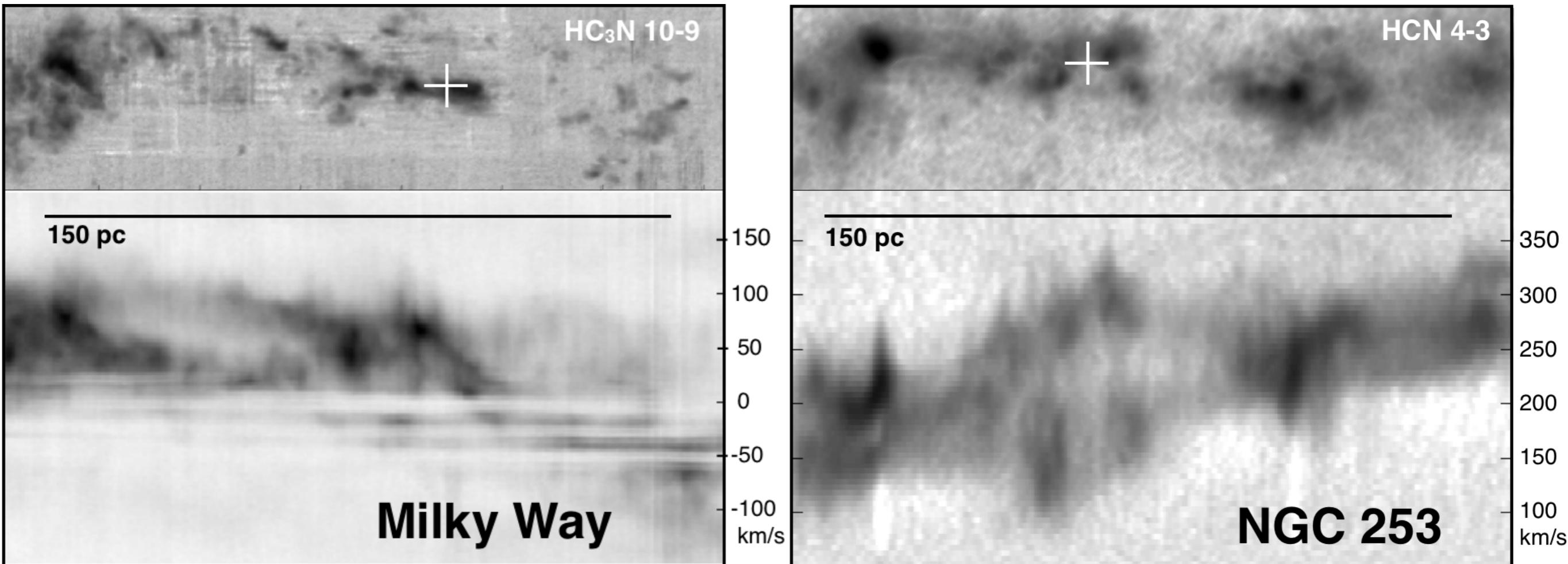
NGC253: An order of magnitude higher molecular mass



Galactic Center HCN 4-3 (ASTE, K. Tanaka et al.)



With ALMA, can now compare the centers of these galaxies at identical resolutions; working on building up a suite of identical tracers



So far:

- Appears to have a nearly identical orbital stream structure
- Broader line widths

Comparing NGC 253 and the Galactic center

Begin to isolate the effects of an active starburst (minus an AGN) on the gas reservoir

Contrast with ‘initial’ conditions of just gas accretion, without a starburst or AGN