

All credits to:

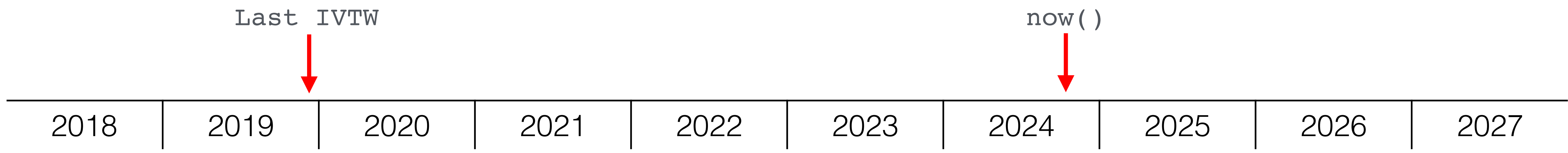
Aard Keimpema

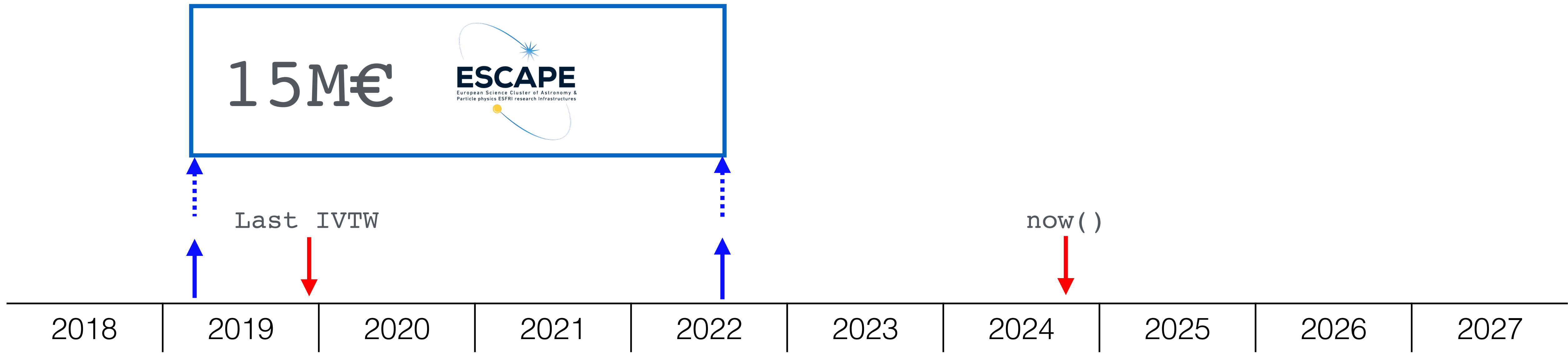
Des Small

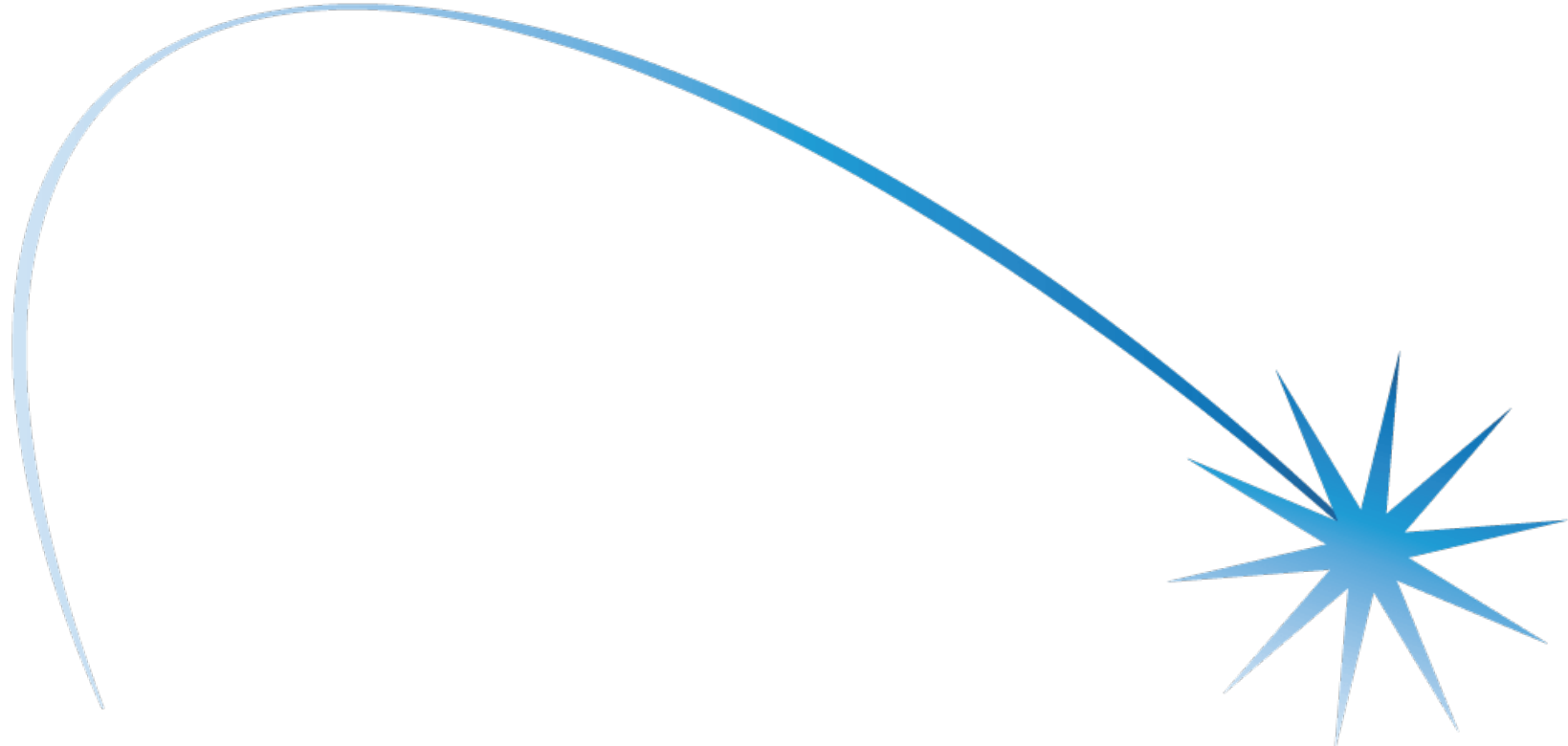
Mark Kettensis

Bob Eldering



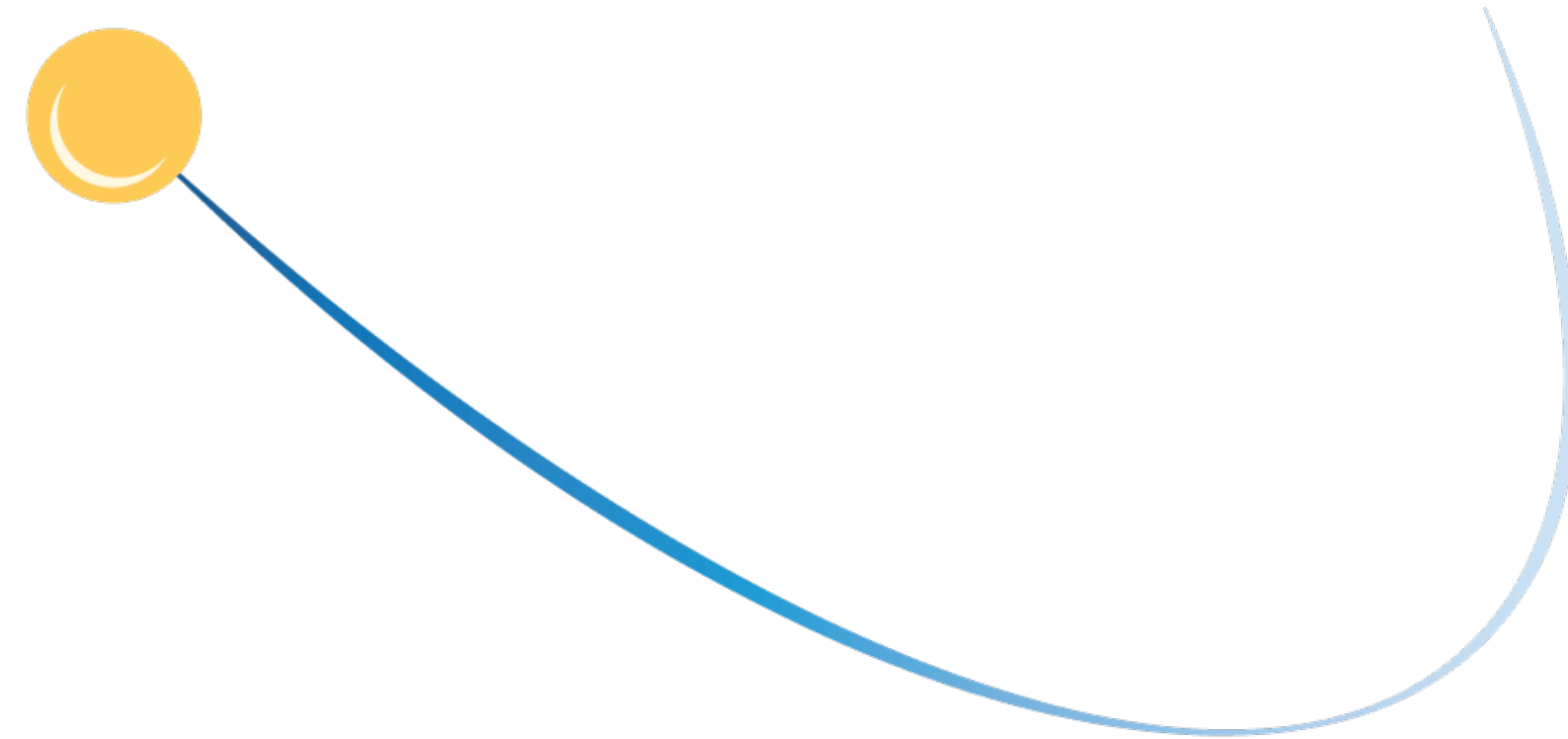






ESCAPE

European Science Cluster of Astronomy &
Particle physics ESFRI research Infrastructures



ESCAPE - The European Science Cluster of Astronomy & Particle Physics ESFRI Research Infrastructures has received funding from the European Union's Horizon 2020 research and innovation programme under the Grant Agreement n° 824064.

Make astronomy data, tools and
results FAIR

Findable

Accessible

Interoperable

Repeatable

WP1 : MIND

WP2 : DIOS

WP3 : OSSR

WP4 : CEVO

WP5 : ESAP

WP6 : ECO

WP3 : OSSR

WP4 : CEVO

WP5 : ESAP

WP3 : OSSR

Open-source scientific Software and Service Repository

“The aim of WP3 is to expose the tools of the ESCAPE ESFRI projects in a repository under the EOSC catalogue of services, ensuring compatibility with *FAIR* principles.”

WP4 : CEVO

Connection ESFRI projects to EOSC through VO framework

“Assess and implement the connection of the ESFRI and other astronomy Research Infrastructures to the EOSC through the Virtual Observatory framework, establish data stewardship practices for adding value to the scientific content of ESFRI data archives”

WP5 : ESAP **ESFRI Science Analysis Platform**

“ ... a high-performance, scalable science platform to provide users with an analysis environment close to the data collections, where they can run their own customized processing and share results and workflows with other researchers, ... ”

WP5 : ESAP ESFRI Science Analysis Platform

Goal is to provide a portal which links the tools from **WP3** with data found through **WP4** and deploy that on resources you have access to through e.g. EduGain or other scientific collaborations

Thu 24 Mar 2022



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Daily Mail Front Page 24...
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Metro Front Page 24th of ...
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Newspaper front pages f...
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March 24, 2022 | New York ...
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Financial Times on Twitte...
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UK Newspaper Front Pag...
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m.thepaperboy.com



Thursday, 24 March 2022
thesouthafrican.com



Daily Telegraph Front Pa...
tomorrowpapers.co.uk



UK Newspaper Front Pag...
m.thepaperboy.com



The Guardian on Twitter: "G...
twitter.com



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globalcirculate.com



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twitter.com



UK Newspaper Front Page...
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```

# Import Virtual Observatory support modules
import pyvo
from pyvo.registry import search as regsearch

# helper function to actually do the query once we have a service
def do_query(service):
    print("#####")
    print("Asking service '{0}'".format(service.res_title) )
    try:
        res = pyvo.dal.TAPService( service.access_url ).run_sync( """
            SELECT TOP 20 *
            FROM ivoa.obscore
            ORDER BY t_min ASC""" )
        if res:
            print( res )
        else:
            print("No results")
    except Exception as e:
        print(e)
    print()

#####
# Ask the Virtual Observatory for
# all radio catalog services
#####
for srv in regsearch(waveband='radio', datamodel='ObsCore'):
    do_query( srv )

```

```

<22/search_evn_archive.py [FORMAT=unix] [TYPE=PYTHON] [ASCII=035] [HEX=23] [POS=0001,0001][2%] [LEN=36]
"search_evn_archive.py" 36L, 878B

```

```
res = pyvo.dal.TAPService( service.access_url ).run_sync( """  
    SELECT TOP 20 *  
    FROM ivoa.obscore  
    ORDER BY t_min ASC""" )
```

```
#####  
# Ask the Virtual Observatory for  
#   all radio catalog services  
#####  
for srv in regsearch(waveband='radio', datamodel='ObsCore'):  
    do_query( srv )
```


#####

Asking service 'The VO @ ASTRON TAP service'

obs_id	dataprod	product_type	calib_level	t_min	target_name	s_ra	s_dec	obs_freq_mhz
				d		deg	deg	m
object		object	int16	float64	object	float64	float64	float64

sauron/ngc2549.Mom0.High.fits		image	2	53007.04466		124.7450	57.80194	--

#####

Asking service 'EVN Data Archive TAP service'

obs_id	dataprod	product_type	calib_level	t_min	target_name	s_ra	s_dec	obs_freq_mhz
				d		deg	deg	m
object		object	int16	float64	object	float64	float64	float64

ES023		visibility	1	51139.81287326384	0016+731	4.940776791666668	73.45833816666666	4998.948305764251

#####

Asking service 'Nobeyama Radio Telescope FITS Archive'

obs_id	dataprod	product_type	calib_level [1]	t_min [1]	target_name	s_ra [1]	s_dec [1]	obs_freq_MHz [1]
				d		deg	deg	
object		object	int32	float64	object	float64	float64	float64

FGN00000003		cube	3	0.0	01100+0000 (2x2)	272.45757353	-19.41493744	110314.4162485902

#####

Asking service 'The VO @ ASTRON TAP service'

obs_id	dataprod	uct_type	calib_level	t_min	target_name	s_ra	s_dec	obs_freq_mhz
				d		deg	deg	m
object	object		int16	float64	object	float64	float64	float64

sauron/ngc2549.Mom0.High.fits	image		2	53007.04466		124.7450	57.80194	--

#####

Asking service 'EVN Data Archive TAP service'

obs_id	dataprod	uct_type	calib_level	t_min	target_name	s_ra	s_dec	obs_freq_mhz
				d		deg	deg	m
object	object		int16	float64	object	float64	float64	float64

ES023	visibility		1	51139.81287326384	0016+731	4.940776791666668	73.45833816666666	4998.948305764251

#####

Asking service 'Nobeyama Radio Telescope FITS Archive'

obs_id	dataprod	uct_type	calib_level	[1]	t_min	[1]	target_name	s_ra	[1]	s_dec	[1]	obs_freq_MHz	[1]
					d				deg	deg			
object	object		int32		float64		object	float64		float64		float64	

FGN00000003	cube		3		0.0		01100+0000 (2x2)	272.45757353		-19.41493744		110314.4162485902	



EVN Data Archive

JIVE provides VO publication services on behalf of the European VLBI Network (EVN).

Please check out our [site help](#). Data [on this site](#) can also be queried through [TAP](#) and an [ADQL form](#).

Services available here

By Title

By Subject

E...

- [EVN Archive TAP service](#)

The EVN Archive's TAP end point. The Table Access Protocol (TAP) lets you execute queries against our database tables, inspect various metadata, and upload your own data. It is thus the VO's premier way to access public data holdings. Tables exposed through this endpoint include: main from the evn schema, obscure from the ivoa schema, columns, groups, key_columns, keys, schemas, tables from the tap_schema schema.

https://evn-vo.jive.eu

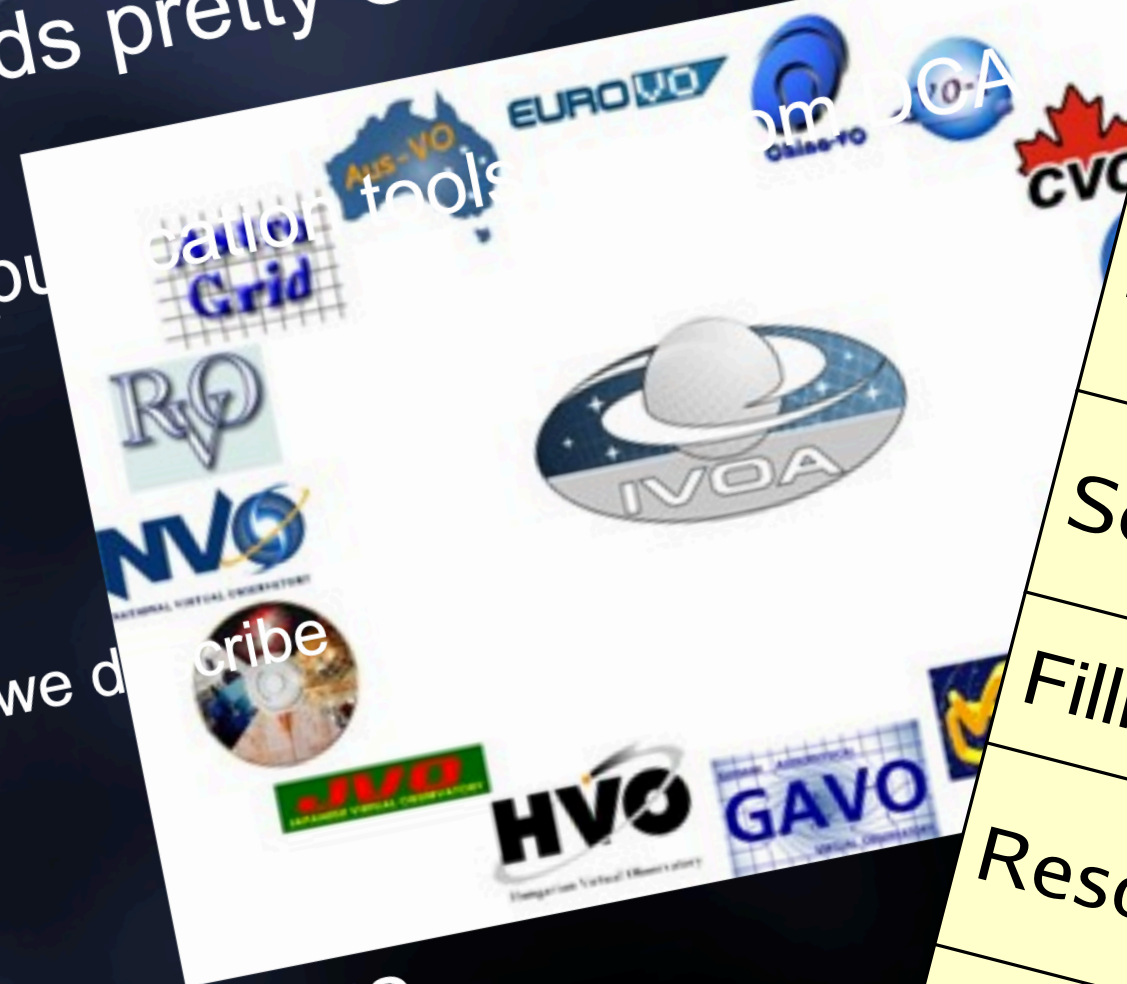
IVOA: Interoperability Standards

- VOExplorer exposes registry weaknesses
 - mostly in schemata/implimentation
- Images, spectra, time standards pretty OK
 - maybe some too complex
 - User-friendly metadata entry/ p
- Not yet covered:
 - Polarization
 - Visibility plane quantities
 - Do we need these or should we d

• Do we need a proposals standard?

Northstar/ALMA phase 1 to 2 proposal tools Dec 2007 Jive

SlidePlayer 10 / 10



A M S Richards

Visibility data for potential products

General	Spatial	Temporal	Spectral	Observable
frame/units	ICRF, deg	MJD	MHz	Jy/beam
Location	13.123456 +55.987654	50613.5	1658	0.001
Bounds	12.92, +55.58 13.32, +56.38	50613.0 50614.0	1650 1665	0.0002 0.5 (or function)
Support	13.123456 +55.987654	(on-source scan listing URL)	1650 1665	undef
Sensitivity	0.4	undef	(bandpass LUT URL)	1
Filling Factor	1	0.7	0.93	undef
Resolution	0".2 2".0 0".2 2".0	5 m	1000 kHz	50 100 μJy/beam
Sampling	0".04 0".0625 0".04 0".0625	16 s	1000 kHz	undef

From presentations by A.M. Richards, c. 2006, AVO+EuroVO workshops

Visibility data for potential

ral

Spatial

Temporal

- User-friendly metadata entry
- Not yet covered:
 - Polarization
 - Visibility plane quantities
- Do we need these or should we d



From presentations by A.M. Richards, c. 2006, AVO+EuroVO workshops



JIVE

Joint Institute for VLBI
ERIC

VLBI into the VO

thanks to the ESCAPE project

Private

Not Secure — archive.jive.nl/scripts/portal.php

Settings

Info

Share

+

Grid

iptables - Fai... - Server Fault SFXC Real-Time Fringe Plot INFRA-TECH P...oogle Drive ESCAPE plan... Documenten Overview - ...REDMINE OSU VLBI and PFB - Google Docs Joint R&D an...Group topics JIVE ZoomRoom#1 JIVE ZoomRoom#2 >>

EVN Data Archive at JIVE

The [EVN](#) Data Archive at [JIVE](#) contains correlated data associated with [EVN](#) observations processed at [JIVE](#). The archive includes a growing database of VLBI observations that have entered the public domain.

In addition, the archive makes available various correlator and pipeline products that give an impression of the data quality. In some cases, preliminary images of calibrators and target sources are also available. The archive allows these to be combined with external VO resources in a natural way.

Select EVN experiment

ES023

Access to EVN archive

- [Show experiment ES023](#)
- [Show catalogue of experiments](#)
- [Search archive by sourcename or position](#)
- [The Bologna archive of EVN observations.](#)

Info

- [Increase of data since 2000](#)
- [Web statistics](#) since June 2004

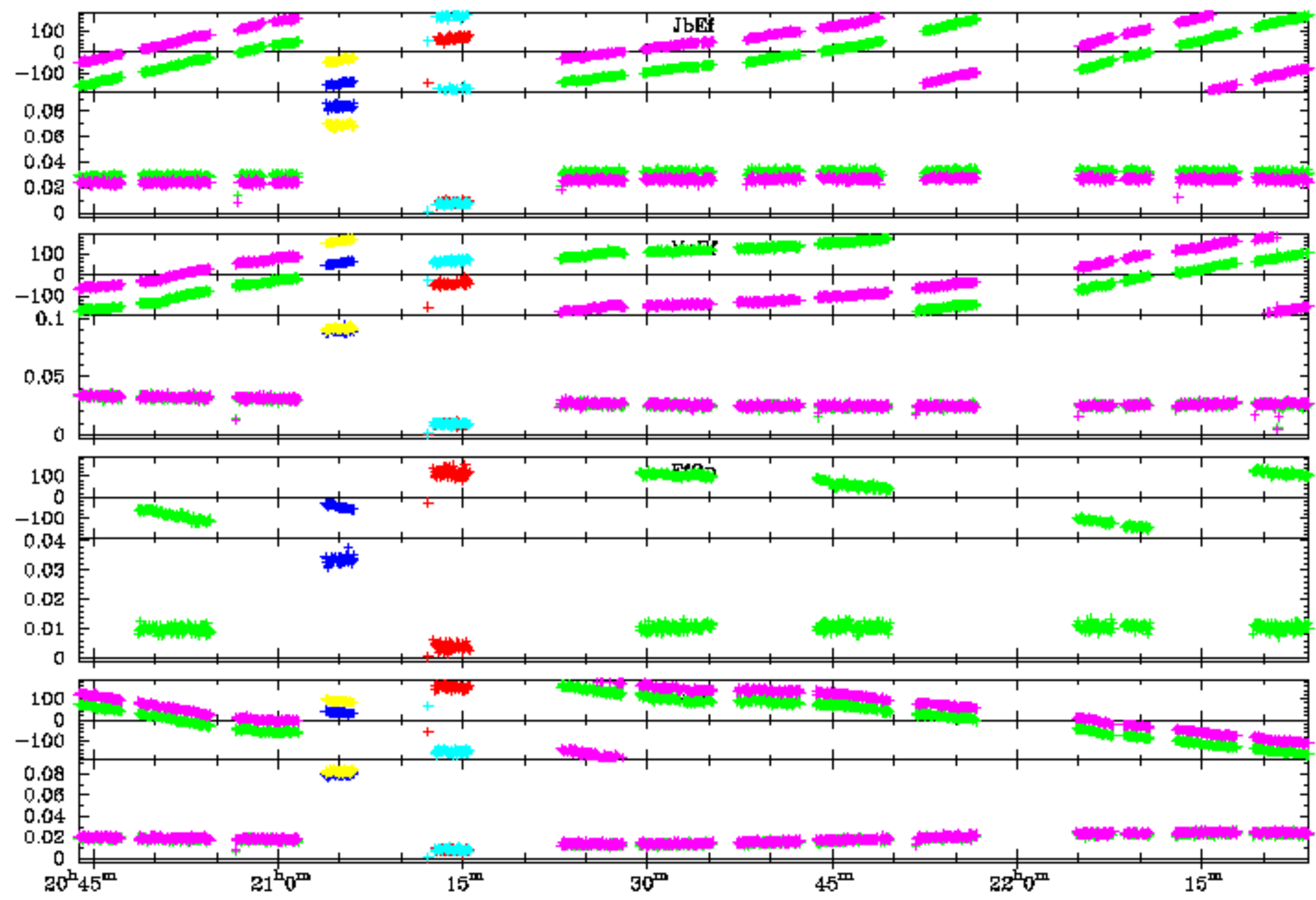
Select a sourceposition from EVN experiment ES023

Ra	Dec	Source	Image	Image
4.9408	73.4583	0016+731		
68.1521	41.6412	0429+415		
88.8784	39.8137	DA193		
202.7845	30.5092	1328+307	sdss	

Access to VO archives

- [Aladin Sky Atlas](#)
- [Sloan Digital Sky Survey](#)

phase
amplitude



→ time

#####

Asking service 'The VO @ ASTRON TAP service'

obs_id	dataprod	type	calib_level	t_min	target_name	s_ra	s_dec	obs_freq_mhz
				d		deg	deg	m
object	object		int16	float64	object	float64	float64	float64

sauron/ngc2549.Mom0.High.fits	image		2	53007.04466		124.7450	57.80194	--

#####

Asking service 'EVN Data Archive TAP service'

obs_id	dataprod	type	calib_level	t_min	target_name
				d	
object	object		int16	float64	object

ES023	visibility		1	51139.81287326384	0016+75

Level 0: Raw instrumental data
Level 1: Instrumental data in a standard format (FITS, VLA
Level 2: Calibrated, science ready data with the instrument
Level 3: Enhanced data products like mosaics, resample
Level 4: Analysis data products generated after some

#####

Asking service 'Nobeyama Radio Telescope FITS Archive'

obs_id	dataprod	type	calib_level	[1]	t_min	[1]	target_name	s_ra	[1]	s_dec	[1]	obs_freq_MHz	[1]
					d				deg	deg			
object	object		int32		float64		object	float64		float64		float64	

FGN00000003	cube		3		0.0		01100+0000 (2x2)	272.45757353		-19.41493744		110314.4162485902	

Private

Not Secure — archive.jive.nl/scripts/arch.php?exp=ES023

⚙️ ⓘ 📄 + 🗑️

iptables - Fai... - Server Fault

SFXC Real-Time Fringe Plot

INFRA-TECH P...oogle Drive

ESCAPE plan... Documenten

Overview - ...REDMINE OSU

VLBI and PFB - Google Docs

Joint R&D an...Group topics

JIVE ZoomRoom#1

JIVE ZoomRoom#2

>>

The EVN MkIV Data Processor at JIVE

archive.jive.nl/scripts/arch.php?exp=ES023

Archive Info

Station Feedback

Station Logfiles

Standard plots

Pipeline calibration

Fitsfiles

Abstract

EVN fitsfiles of experiment ES023

Access status: public

Download: Use right mousebutton -> Save target.

If the connection is slow, try [GNU wget. \(manual\)](#).
It can be obtained from the web, if not available.

A file selection can be made by filling in the wildcard after the -A option.
To get all fitsfiles of ES023 copy next line to your commandwindow:

```
wget -t45 -l1 -r -nd http://archive.jive.nl/exp/ES023_981124/fits -A "es023**"
```

The checksum file can be used to verify the checksum of all datafiles using:
md5sum -c es023.checksum (on unix systems).

Filename	Length x 10 ⁹ bytes
es023.checksum	0.000000043
es023_1_1.IDI1	0.513987840

FITS Interferometry Data Interchange (FITS-IDI) Convention
<https://fits.gsfc.nasa.gov/registry/fitsidi.html>

Pipeline products of experiment ES023	
Pipeline plots	
AIPS calibration. tables (FITS Format)	
AIPS history file.	
Short summary of CL/SN table contents.	
Input parameters for script.	
Associated EVN calibration.	
Associated VLBA / VLA / GBT file.	(Not available)
UVFLG flagged data.	(Not available)
UVFLG Band-edge Flagging.	
The pipeline logfile.	(Not available)
Pipeline-calibrated UV FITS files.	(Not available)

- Archive Info
- Station Feedback
- Station Logfiles
- Standard plots
- Pipeline calibration
- Fitsfiles
- Abstract

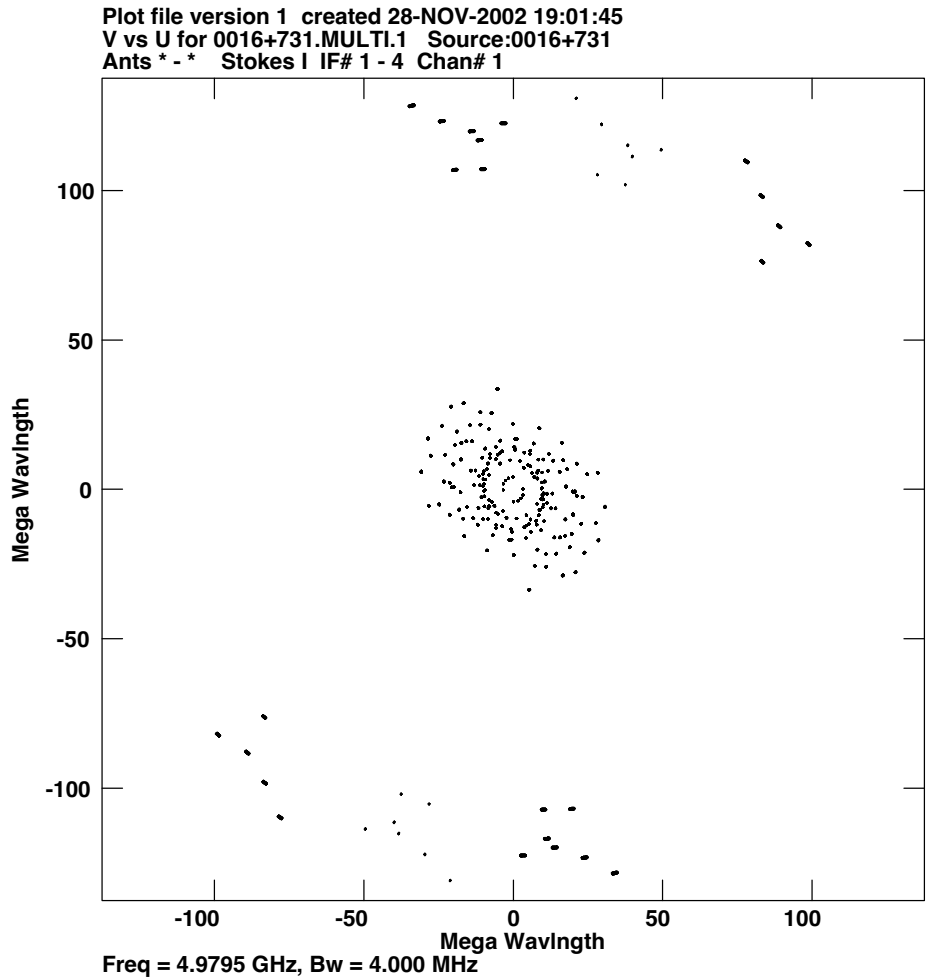
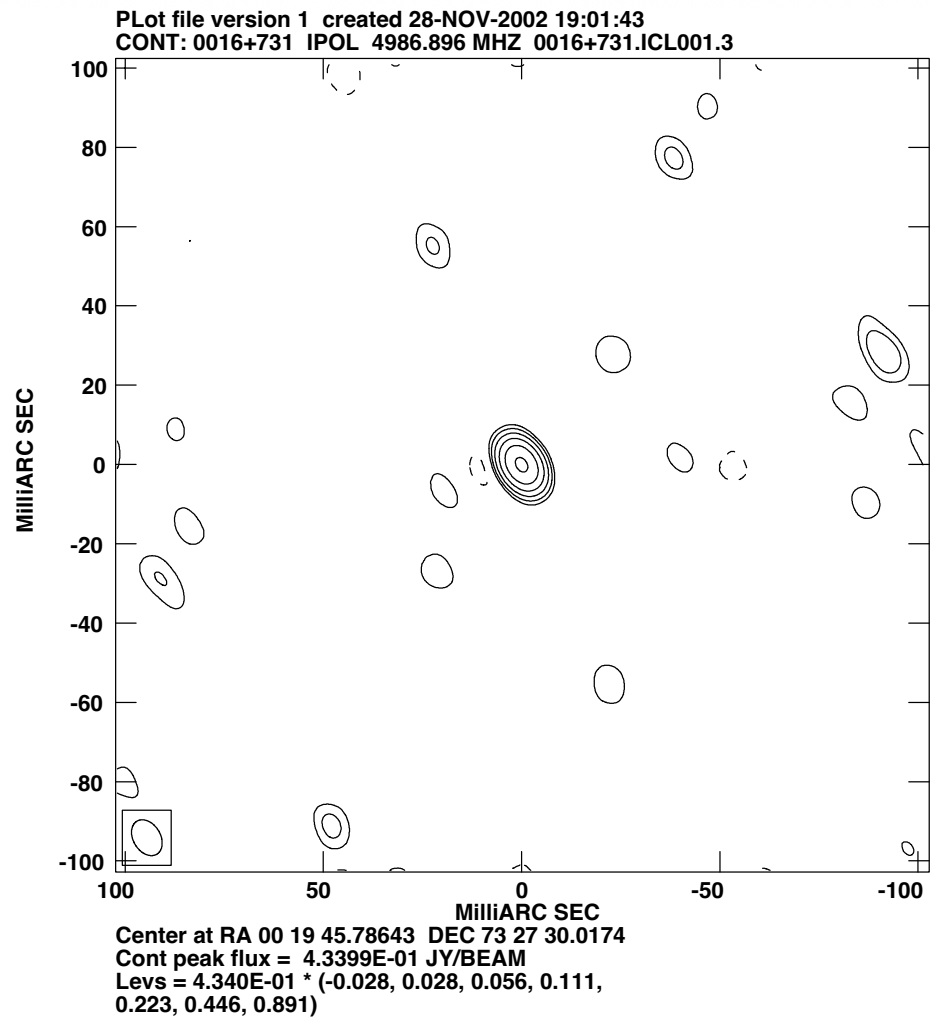
The target: Up to 250 Jy at 10M lambda. Detected up to 141M lambda.

- [u,v coverage for 0016+731.](#)
- [u,v coverage for 1328+307.](#)
- [u,v coverage for DA193.](#)
- Comments.

0016+731: A central condensation elongated in the NE-SW direction up to 35M lambda, and (u,v) data points in the north-south direction around 120M lambda.
1328+307: Sparse distribution up to 12M lambda.
DA193: A central condensation elongated in the north-south direction up to 40M lambda, and (u,v) data points in the NE-SW direction around 120M lambda.
The target: A central condensation elongated in the north-south direction up to 40M lambda, and (u,v) data points in the NE-SW direction at 110-130M lambda.

- [Crude map of 0016+731 from the pipeline.](#)
- [Crude map of 1328+307 from the pipeline.](#)
- [Crude map of DA193 from the pipeline.](#)
- Comments.

0016+731: An unresolved component.
1328+307: Out of the field of view.
DA193: An unresolved component. Bright side lobes still exist.
The target: A couple of components.



Why add VLBI
to the VO then?

Virtual Observatory use cases

- ▶ Historic data (“before picture”) for high-res follow-up of:
 - Gravitational Wave events
 - Gamma Ray Bursts
 - Fast Radio Bursts
- ▶ Field-of-view in data set is larger than primary science goal
 - possibility to reprocess Level 1 data for other science
 - apply new(er) calibration algorithms to improve
- ▶ Standardized access to archive data for science platform
 - JupyterLab environment
 - both human and machine accessible
 - not just the Level 1 data

Applicable protocols:

ObsCore

“defines the core components of the Observation data model that are necessary to perform data discovery when querying data centers for astronomical observations of interest.”

<https://ivoa.net/documents/ObsCore/20170509/index.html>

DataLink

“linking of data discovery metadata to access to the data itself, further detailed metadata, related resources, and to services that perform operations on the data.”

<https://ivoa.net/documents/DataLink/20150617/index.html>

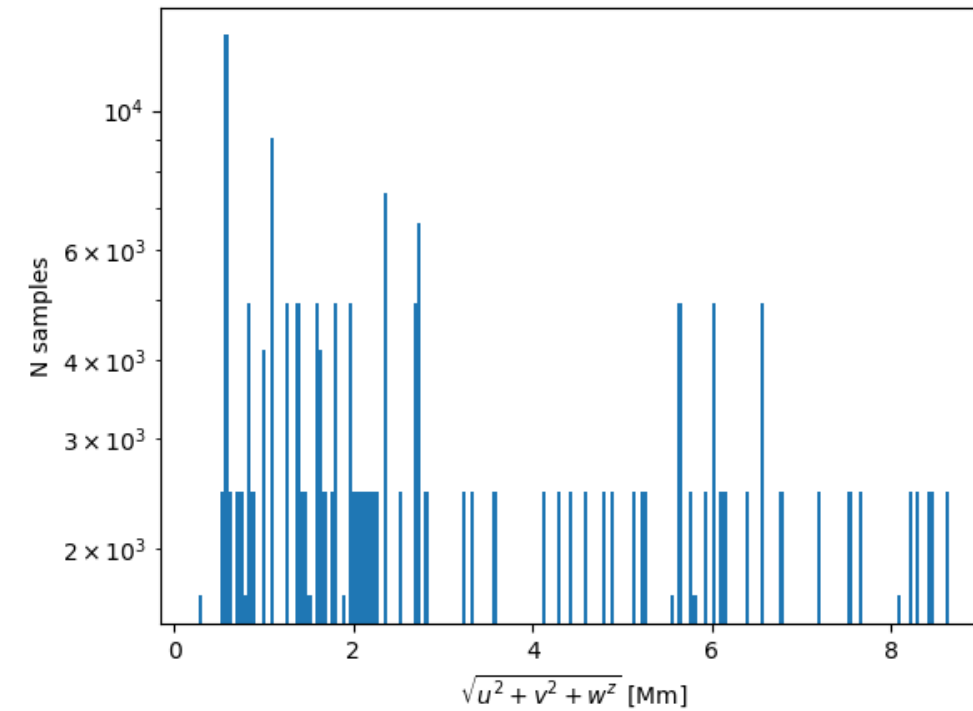
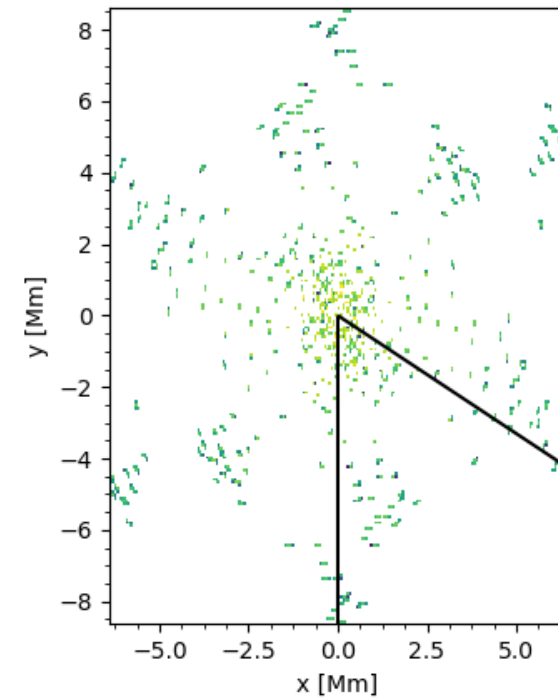
ESCAPE WP4 enabled:

Visibility support in ObsCore

- JIVE participation in IVOA Radio SIG
- extend ObsCore data model w/ u,v-plane values
- experimental u,v-plane characterisation

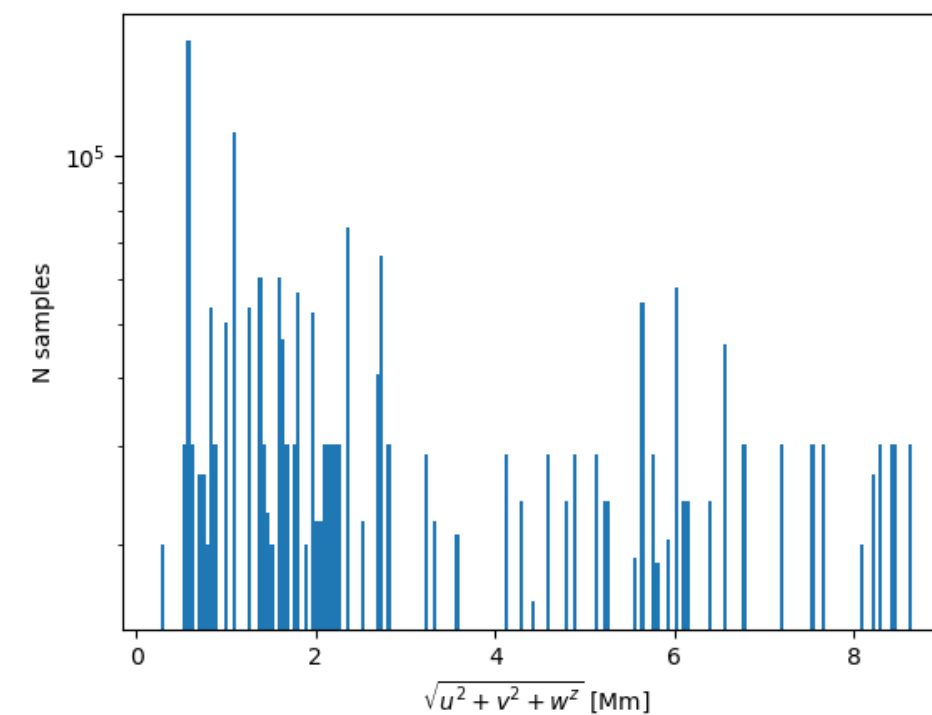
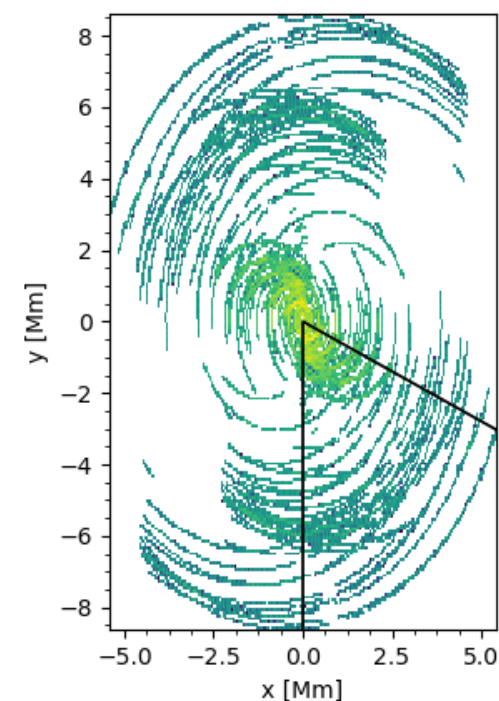
u,v-plane characterisation

Plots derived from software developed by Mattia Mancini (ASTRON)



$$e = 0.74$$
$$f = 0.03$$

calibrator source



$$e = 0.63$$
$$f = 0.25$$

target source

e = ellipticity
 f = filling factor

ESCAPE WP4 enabled:

Visibility support in ObsCore

- JIVE participation in IVOA Radio SIG
- extend data model
- experimental u,v-plane characterisation

ObsTAP service at JIVE:

- design mapping of FITS-ID1 content to ObsCore
- develop s/w to analyse 100 TB of FITS-ID1
- hardware to host <https://evn-vo.jive.eu>

Help

Service info

Related

[Tables available for ADQL](#)

Metadata

Identifier

ivo://jive.eu/___system_

Cite this

[Advice on citing this re](#)

Description

An endpoint for submit

Keywords

[Virtual observatories](#)

Creator

Mark Kettenis

Created

2008-09-20T12:00:00Z

Data updated

2020-12-15T16:41:59Z

Metadata updated

Parameters

- ADQL query: SELECT obs_id, t_min, target_name, s_ra, s_dec, (3e8/em_min)/1e6 as obs_freq_MHz FROM ivoa.obscore WHERE 1 = CONTAINS(POINT('ICRS', s_ra, s_dec), CIRCLE('ICRS', 4.9375, 73.4583, 0.09999999)) ORDER BY t_min

Result

Matched: 12

Send via SAMP

Quick Plot

Obs_id	T_min	Target_name	S_ra [deg]	S_dec [deg]	Obs_freq_mhz [m]
ES023	1998-11-22T19:30:32Z	0016+731	4.940776791666668	73.45833816666666	4998.948305764251
N05L4	2005-06-08T04:46:00Z	J0019+7327	4.940776733333333	73.45833819166667	1667.643687020306
EP049C	2005-06-13T04:20:27Z	J0019+7327	4.940776666666667	73.45833819444447	5057.989150614334
GB063A	2007-10-31T01:34:16Z	0019+732	4.940776779166668	73.45833818055554	6703.627614274413
EZ020A	2010-06-03T19:35:08Z	0019+732	4.940776779166668	73.45833818055554	6703.627614274413
EG061A	2012-02-29T06:37:46Z	J0019+73	4.940776612500001	73.45833819444447	22424.003074820543
...	2014-06-	...	...	...	...

Astronomical Data Query Language – <https://www.ivoa.net/documents/ADQL/>

IVTW @ MIT Haystack

verkouter@jive.eu

October 2024

Private

< >

A A

Not Secure — saada.unistra.fr/taphandle/?url=http://evn-vo.jive.eu/tap#

⚙ ⓘ ⬆ + ☰

iptables - Fai... - Server Fault SFXC Real-Time Fringe Plot INFRA-TECH P...oogle Drive ESCAPE plan... Documenten Overview - ...REDMINE OSU VLBI and PFB - Google Docs Joint R&D an...Group topics JIVE ZoomRoom#1 JIVE ZoomRoom#2 Private Zoom Zoom Rooms | ...ron intranet

AST... * F2... What... Slac... (-) E... rdtsc... Coun... Stud... Crea... NAEI... IVOA... EVN... ADQ... TapHan... The... archi... archi...

Preview of http://evn-vo.jive.eu/evn/q/dl/dlmeta?ID=ivo%3A//jive.eu/~%3FES023_981124_1_1_1328%252B307_4979.49MHz

Links for ivo://jive.eu/~?ES023_981124_1_1_1328%2B307_4979.49MHz

▼ the data itself

the primary (as opposed to related) data of the identified resource

FITS-IDI

▼ Applicable Calibration

Data products that can be used to remove instrumental signatures from #this. Note that the calibration steps such data products feed have not been applied to #this yet.

UVFLG

ANTAB

Tap Nodes

jiveeu-ta

tap_sch

ivoa

evn

Goodies (n

_dec	s_fc
0.509155	0.00
3.458338	0.00
9.813657	0.00
1.641231	0.00
9.810276	0.00
.69839262	0.00
9.810276	0.00
.69839262	0.00



jiveeu-tap **SJAU**


tap_schema

ivoa

evn

Goodies (not used yet)

```

SIMPLE = T /Yep! It's a FITS file...                                BITPIX = 8
0                                          EXTEND = T                                           GROUPS =
GCOUNT = 0                                          PCOUNT = 0
XTENSION= 'BINTABLE'                               BITPIX = 8
2 /A table is a matrix isn't it?                NAXIS1 = 64                               NAXIS2 =
PCOUNT = 0                                          GCOUNT = 1
7 /Number of columns in each row                  EXTNAME = 'ARRAY_GEOMETRY'           EXTVER =
TTYPE1 = 'ANNAME ' /The antennaname              TFORM1 = '8A '
'                                                  TFORM2 = '3D '                               TUNIT1 =
TTYPE3 = 'DERXYZ '                                  TFORM3 = '3E '                               TFORM3 =
'                                                  TTYPE4 = 'ORBPARM ' /Orbital parameters      TFORM4 =
TTYPE5 = 'NOSTA ' /Antenna number                 TFORM5 = '1J '                               TTYPE5 =
' /Mount                                            TFORM6 = '1J '                               TUNIT7 =
offset                                             TFORM7 = '3E '                               TUNIT7 = 'METERS
TABREV = 1                                          STK_1 = -1                                     NO_BAND =
4                                          REF_FREQ= 4.979490000000E+09
NO_CHAN = 16                                      REF_PIXL= 1.0000000E+00
2.500000000000E+05                              ARRAYX = 0.000000000000E+00
FRAME = 'GEOCENTRIC'                             ARRAYZ = 0.000000000000E+00
0.000000000000E+00                              FREQ = 4.979490000000E+09
DEGPDY = 3.609856422982E+02                      POLARX = 4.931838513442E+00
11-22'                                           IATUTC = 3.100000000000E+01
UT1UTC = -2.370471954346E-01                    ARRNAM = 'EVN '                               END
JB AM*[ ^s[ OÁ[ EQû~' QASg[ z¹õz[000000000000[0000>ê~ú[00000000MC AQ[I|üPHA,[Y¼Æ?[APùEÍ[W?000000000000[0000?éä÷00000000EF ANÆÍ¹ÇySA
³[000000000000[0000<I...ð[00000000ON AI·ë[ç3æA%¶Söz['ATHH[-Ûè[000000000000[000[e ¯š[00000000NT ARÒèÅ% A4(ñNQÑ·AM
Š@ÁTÊ[000000000000[0000?èQ[00000000TR AKÂ[=fİBA2¥QÀH[ðAS^[1£~([000000000000[000000000000SH ÁEš£~$Ý/AQÖ%j&€[AHý[ÔVÕÐ[000000000000
AMèð[K]ì@£Ü[û~' QAS Ã[[êK[000000000000[0000?ÀbN[00000000WB AM5nÑ[ @OA[,ž‡_oðASR6ehr°[000000000000 000[ežff[00000000
XTENSION= 'BINTABLE'                               BITPIX = 8
2 /A table is a matrix isn't it?                NAXIS1 = 84                               NAXIS2 =
PCOUNT = 0                                          GCOUNT = 1
5 /Number of columns in each row                  EXTNAME = 'FREQUENCY'           EXTVER =
TTYPE1 = 'FREQID '                                  TFORM1 = '1J '
'BANDFREQ'                                         TFORM2 = '4D '
TTYPE3 = 'CH_WIDTH'                                 TFORM3 = '4E '
'                                                  TTYPE4 = 'TOTAL_BANDWIDTH'       TFORM4 =
TUNIT4 = 'HZ '                                       TTYPE5 = 'SIDE BAND'            TUNIT4 =
'                                                  TABREV = 1                          OBSERVE =
NO_STKD = 4                                          STK_1 = -1
4                                          NO_CHAN = 16                            REF_FREQ= 4.979490000000E+09
CHAN_BW = 2.500000000000E+05                     REF_PIXL= 1.0000000E+00
000[00000000AN,,€0000A^,,€0000Afā`0000Ht$0Ht$0Ht$0Ht$0Jt$0Jt$0Jt$0Jt$0000[000[000[000[
XTENSION= 'BINTABLE'                               BITPIX = 8
2 /A table is a matrix isn't it?                NAXIS1 = 268                               NAXIS2 =
PCOUNT = 0                                          GCOUNT = 1
23 /Number of columns in each row                 EXTNAME = 'SOURCE '             EXTVER =

```

More ESCAPE impact!

ESCAPE WP3 (open source repository)

- improve VLBI/radio data reduction tools*
- provide Jupyter kernel w/ data reduction s/w

** that means CASA*

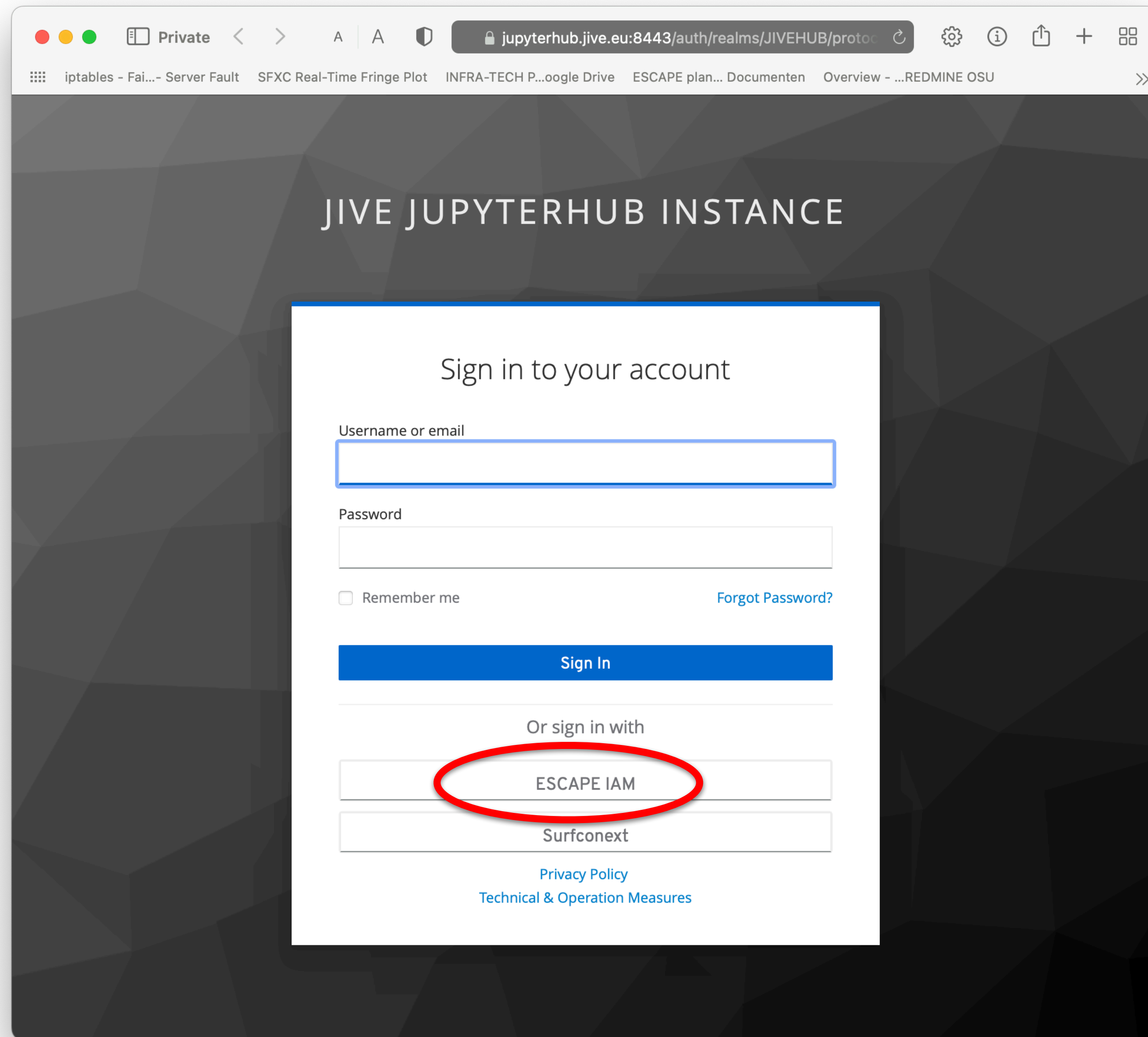
More ESCAPE impact!

ESCAPE WP3 (open source repository)

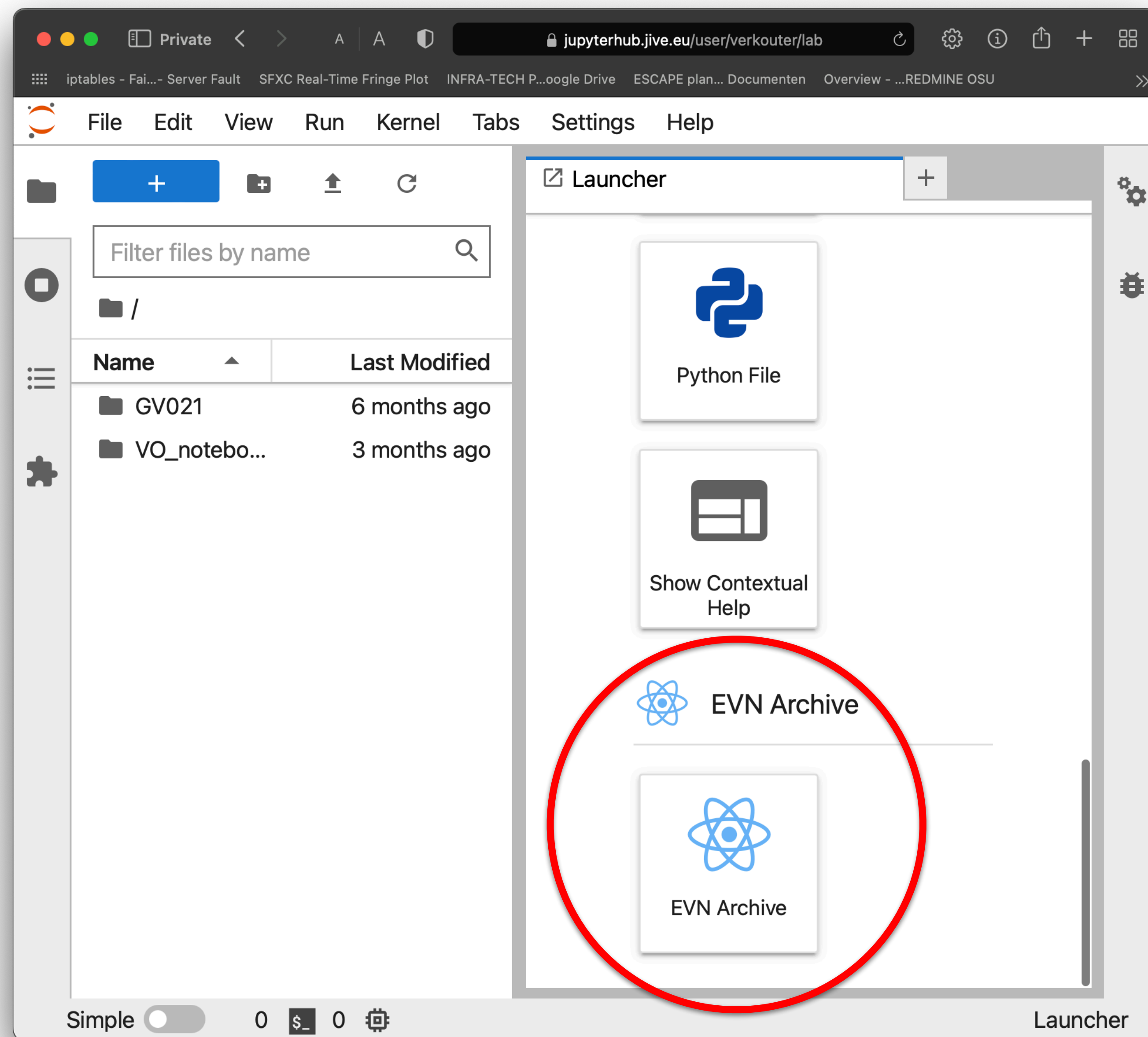
- improve CASA VLBI data reduction tools
- provide Jupyter kernel w/ data reduction s/w

ESCAPE WP5 (science analysis platform)

- provide Jupyterhub hardware (compute+storage)
- development of Jupyterhub plugins:
 - EVN archive search w/ notebook deployment
 - version controlled notebooks (provenance)



`https://jupyterhub.jive.eu`



Private

jupyterhub.jive.eu/user/verkouter/lab

iptables - Fai... - Server FaultSFXC Real-Time Fringe PlotINFRA-TECH P...oogle DriveESCAPE plan... DocumentenOverview - ...REDMINE OSUVLBI and PFB - Google DocsJoint R&D an...Group topicsJIVE ZoomRoom#1JIVE ZoomRoom#2Private Zoom

FileEditViewRunKernelTabsSettingsHelp

Filter files by name

ES0233 minutes ago

GV021a minut ago

VO_noteb...3 months ago

Launcher

Experiment selection

ExperimentSourceObserving Band

Experiment codeSource nameBand L, C, X, etc.

SEARCH

Right Ascension00:19:45.8Declination73:27:30.0Radius [arcseconds]36

SEARCH

Search results

Actions	Experiment	Source	Dec	Exp. time [s]	Distance [deg]	
↓	ES023	0016+731	00h19m45.786s	73d27'30.017"	908	0.000016808050437962152
Open notebook	EZ020A	0019+732	00h19m45.786s	73d27'30.017"	180	0.000016815457395087267
↓	GB063A	0019+732	00h19m45.786s	73d27'30.017"	88	0.000016815457395087267
↓	GP051D	0019+732	00h19m45.786s	73d27'30.017"	90	0.000016815457410190017
↓	N05L4	J0019+7327	00h19m45.786s	73d27'30.017"	1813.5	0.0000168311570963371

Simple0\$2

Experiment selection

Uses WP 4 EVN VO ObsTAP service to query!

IVTW @ MIT Haystack

verkouter@jive.eu

October 2024

Private

jupyterhub.jive.eu/user/verkouter/lab

iptables - Fai... - Server FaultSFXC Real-Time Fringe PlotINFRA-TECH P...oogle DriveESCAPE plan... DocumentenOverview - ...REDMINE OSUVLBI and PFB - Google DocsJoint R&D an...Group topicsJIVE ZoomRoom#1JIVE ZoomRoom#2Private Zoom

FileEditViewRunKernelTabsSettingsHelp

Filter files by name

/

Name	Last Modified
ES023	3 minutes ago
GV021	a minute ago
VO_noteb...	3 months ago

Launcher

Experiment selection

ExperimentSourceObserving Band

Experiment codeSource nameBand L, C, X, etc.

SEARCH

Right Ascension00:19:45.8Declination73:27:30.0Radius [arcseconds]36

SEARCH

Search results

Source	Ra	Dec	Exp. time [s]	Distance [deg]
0016+731	00h19m45.786s	73d27'30.017"	908	0.000016808050437962152
0019+732	00h19m45.786s	73d27'30.017"	180	0.000016815457395087267
0019+732	00h19m45.786s	73d27'30.017"	88	0.000016815457395087267
GP051D	00h19m45.786s	73d27'30.017"	90	0.000016815457410190017
N05L4	J0019+7327	00h19m45.786s	1813.5	0.0000168311570963371

Simple0\$2

Experiment selection



Private

jupyterhub.jive.eu/user/verkouter/lab/tree/ES023/evn_continuum_ES023.ipynb

iptables - Fai... - Server FaultSFXC Real-Time Fringe PlotINFRA-TECH P...oogle DriveESCAPE plan... DocumentenOverview - ...REDMINE OSUVLBI and PFB - Google DocsJoint R&D an...Group topicsJIVE ZoomRoom#1JIVE ZoomRoom#2Private Zoom

FileEditViewRunKernelTabsSettingsHelp

Filter files by name

/

Name	Last Modified
ES023	seconds ago
GV021	6 months ago
VO_notebo...	3 months ago

LauncherExperiment selectionevn_continuum_ES023.ipynk

Markdown

once!

```
[ ]: from evn_tools.evn_import_exp import evn_import_exp
      expname = "ES023"
      evn_import_exp(expname)
```

Experiment information

Note that the path setting is the path inside the virtual container. The default path is `/home/verkouter/evn`. When launching the container in a Docker environment you can set this with the `-v` option for Docker. If you choose a different path, make sure to adjust the entire `mypath` variable below.

```
[ ]: # If your data is in a non-standard location then point the mypath variable to
      mypath=''
      if mypath != '':
          sys.path.append(mypath)

      # Observation ID in the EVN archive. This is always in small caps, e.g. n14c3
      obsid = 'es023'
```

Simple0\$1Casa | Idle

Mode: CommandLn 1, Col 1 evn_continuum_ES023.ipynb

evn_import_exp(...) uses VO DataLink to download FITS-IDI and calibration data!

Available data → 31342

● in view ● out view

asu.cas.cz → 1
au.csiro → 3
bira-iasb → 1
byu.arvo → 2
cadn.nrc.ca → 2
cdpp → 3
cds.vizier → 1
chivo → 4
esavo → 10
eso.org → 2
fu-berlin.planet.hrsc → 1
gaia.aip.de → 1
geops.ipsl → 1
ia2.inaf.it → 5
iap → 2
idoc → 1
irsa.ipac → 1
jao.alma → 3
jive.eu → 1
EVN Data Archive TAP ser
jvo → 7
koa.ipac → 1
konkoly.hu → 2
lam.cesam → 1
latmos.ipsl → 1
lmd.jussieu → 1
musewide.aip.de → 1
nasa.heasarc → 35
nci.org.au → 3
ned.ipac → 1
oca → 1
org.gavo.dc → 47
osug-vo.osug → 2
ov-gso → 1
padc.obspm.astro → 3
padc.obspm.planeto → 2
padc.obspm.helio → 1
pds-ppi → 1
purx → 4
pvoll → 1
spectrum.iaa → 2
src.pas → 2
ssdc → 1
tohoku.univ.jp → 5
uni-heidelberg.de → 1
vopdc.obspm → 3
voxastro.org → 2
wfau.roe.ac.uk → 4
www.plate-archive.org → 1
xaovo → 2
xcatdb → 2

select

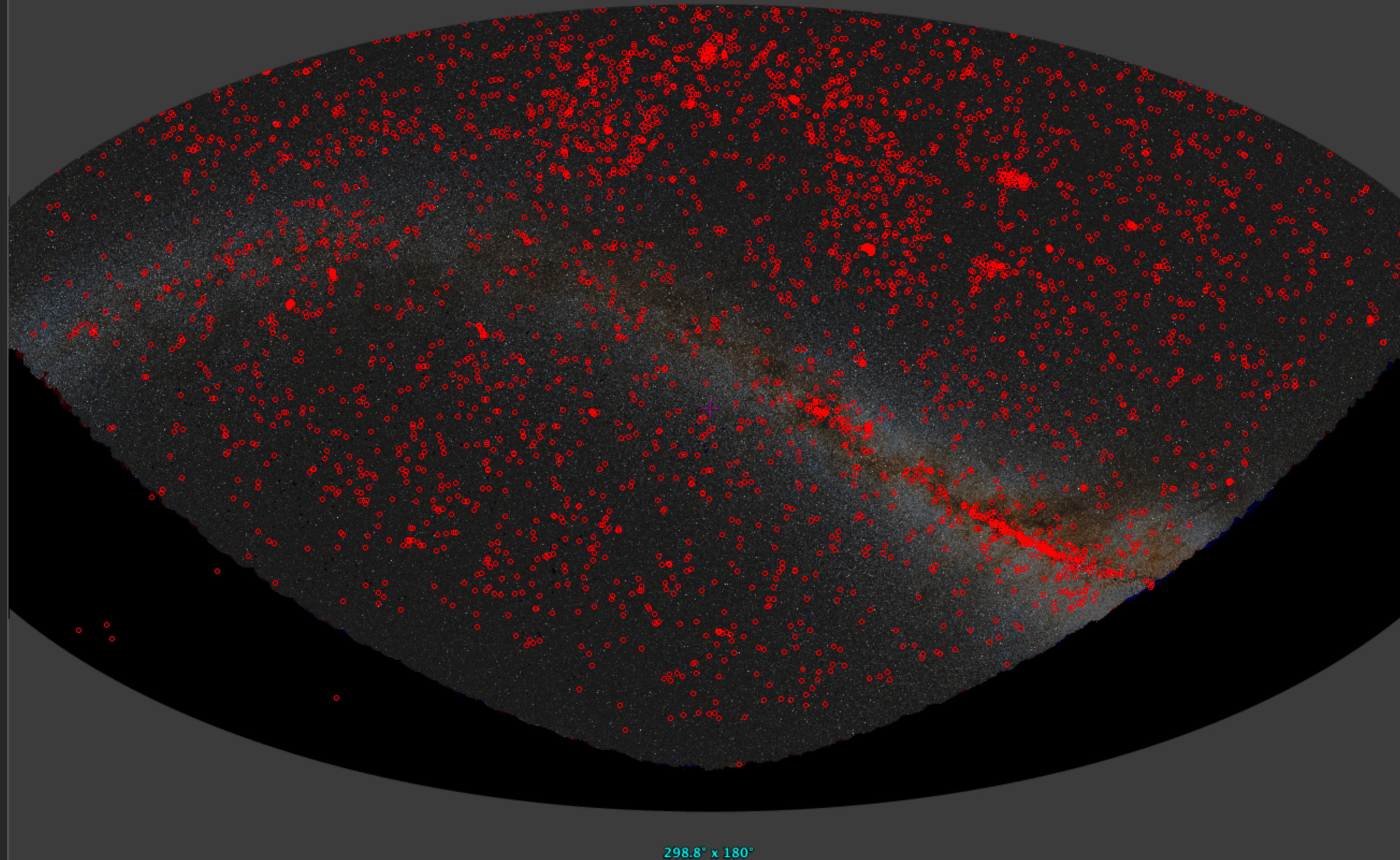
from -- all collections --

coll. sort view scan filter

Command f

DSS PanSTARRS SDSS 2MASS GALEX Gaia Simbad NED +

PanSTARRS DR1 color-z-zg-g



298.8° x 180°

grid studywink redonorthhdr multiview match

Search

ALADIN

select
pan

dist

phot

draw

tag

moc

spect

filter

cross

x-y

rgb

assoc

crop

cont

pixel

prop

del

Imagine your
eye
looking through a
stack of planes
(below).

Each plane
contains its own
data set: image,
catalog, graphical
overlays...

You see the
combination of
them in the main
panel.

For accessing to
other data, use
the discovery tree
in the left panel,
or clic & drag
your own local
files.

jive.eu / tap

CDS / P / Par

epoch -

size -

dens. -

opac. -

zoom -

22:36:02.09

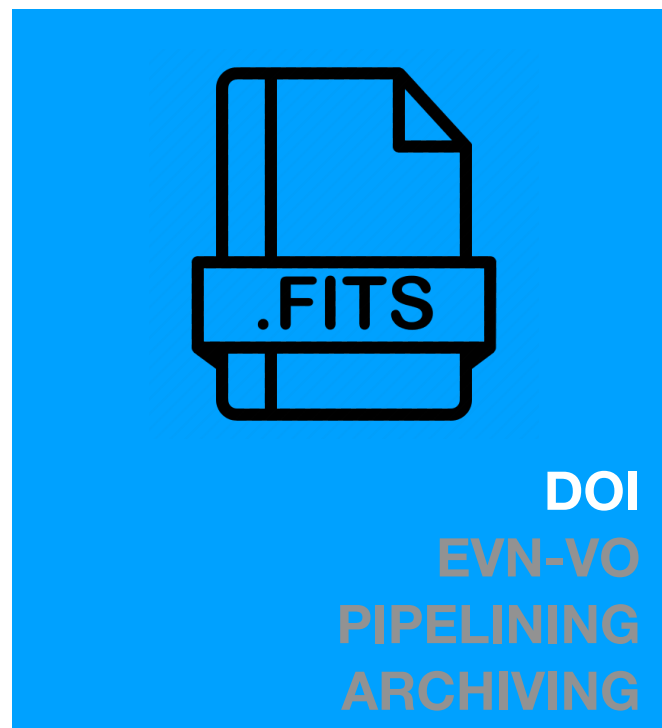
3.8° x 180°

-90

skv

1998-11-22 ...

0 sel / 19224 src 26fps / 1398Mb



<https://doi.org/10.48717/21qn-4a40>

Landing page for ES023

DOI

<https://doi.org/10.48717/21qn-4a40>

Project

ESo23

Title

The CCS Quasar 3C119

► Details

Data

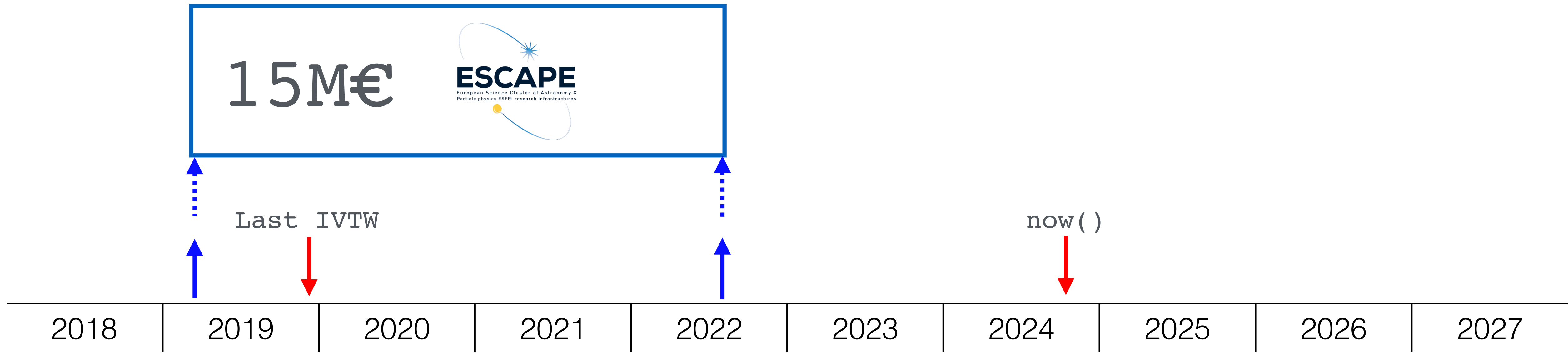
Pages for individual experiments for ESo23 at the EVN archive:

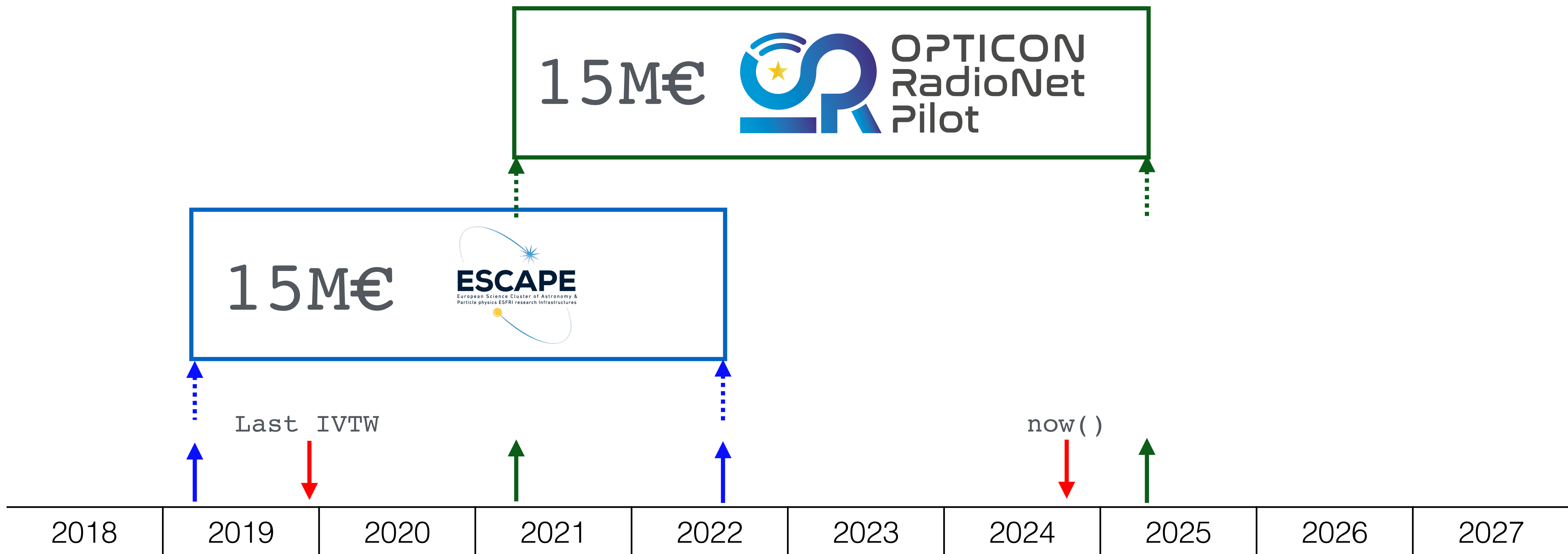
ESo23 (<https://doi.org/10.48717/xfbj-m707>)

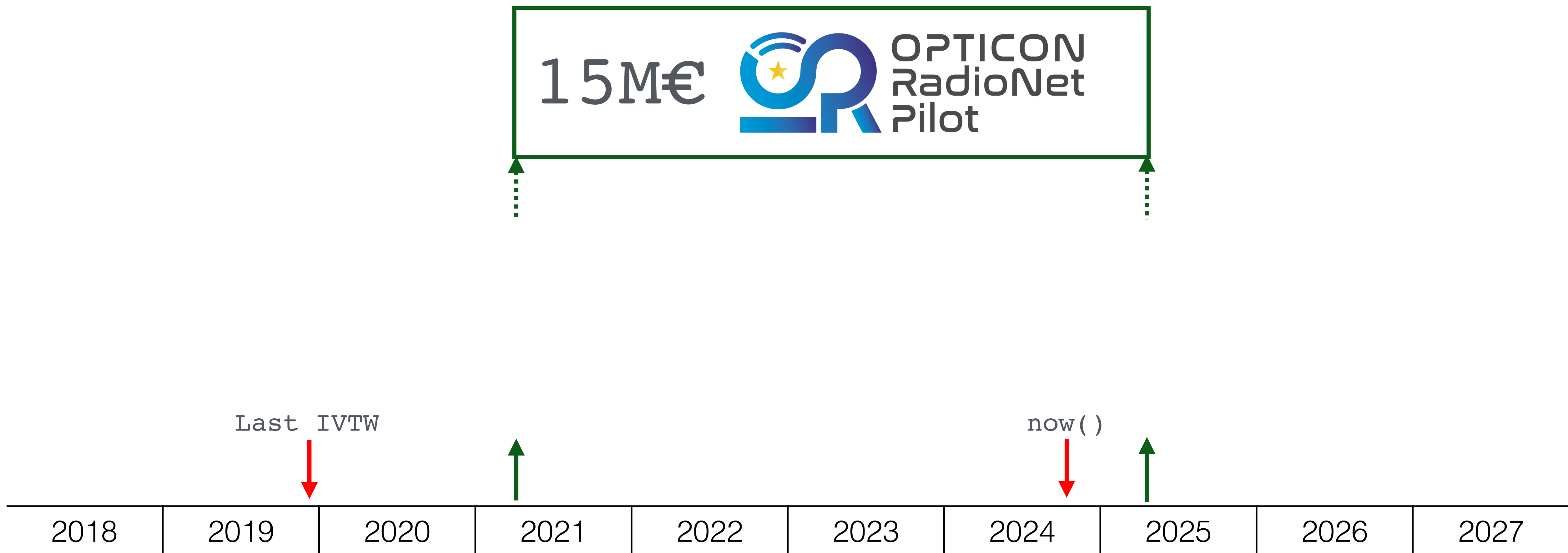
ESo23B (<https://doi.org/10.48717/cs68-tk75>)

Context

This data is part of the archive of VLBI data maintained by [JIVE](#) on behalf of the [EVN](#), a network of radio telescopes located primarily in Europe and Asia. The EVN archive itself has the DOI <https://doi.org/10.17616/R3Z197>









OPTICON RadioNet Pilot



Horizon2020
European Union Funding
for Research & Innovation

This project has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No 101004719





OPTical
Infrared
COordination
Network for
Astronomy

2001 – 2021

<https://arxiv.org/pdf/astro-ph/0105562>



RadioNet
Infrastructure
Cooperation
Network in
Radio Astronomy

1999 – 2021

<https://www.radionet-eu.org/index.html>

1998–2002	RadioNet	VLBI schools EVN symposium Coordination meetings (TOG, CBD)
-----------	-----------------	---

2002–2006	RadioNet	ParselTongue
	EXPreS	Production e-VLBI Distributed SFXC

2007–2013	RadioNet	Basic CASA VLBI improvements
	NEXPreS	UniBoard FPGA hardware FGPA VLBI correlator FlexBuff

2014–2020	RadioNet	White Rabbit T&F transfer CASA fringeFit e-transfer pySCHED
-----------	-----------------	--



OPTical
Infrared
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2001 – 2021

<https://arxiv.org/pdf/astro-ph/0105562>



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Opticon
RadioNet
PILOT

JA2 - *Seamless Performance*

- 2.1 Common-Access to Research Infrastructures
- 2.2 Time-Domain, Multi-Facility & Multi-Frequency access to Research Infrastructures
- 2.3 Common frameworks for Data access and processing procedures
- 2.4 Synergies between emerging and established interferometric communities
- 2.5 Preserving the sky for future generations

JA2 - *Seamless Performance*

- 2.1 Common-Access to Research Infrastructures
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- 2.3 Common frameworks for Data access and processing procedures
- 2.4 Synergies between emerging and established interferometric communities
- 2.5 Preserving the sky for future generations

JA2.2 - Access *to*:

multi-wavelength

multi-facility

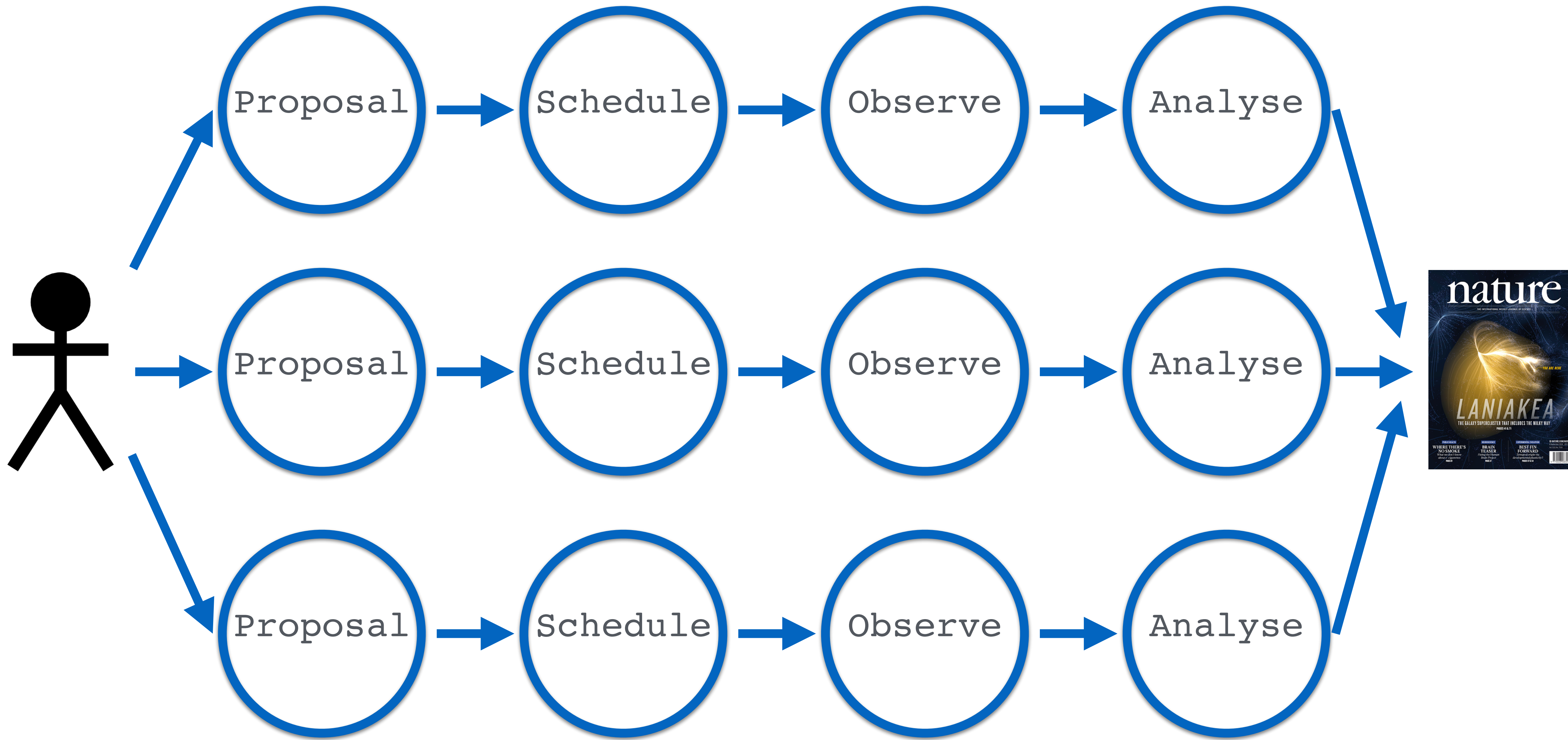
time domain

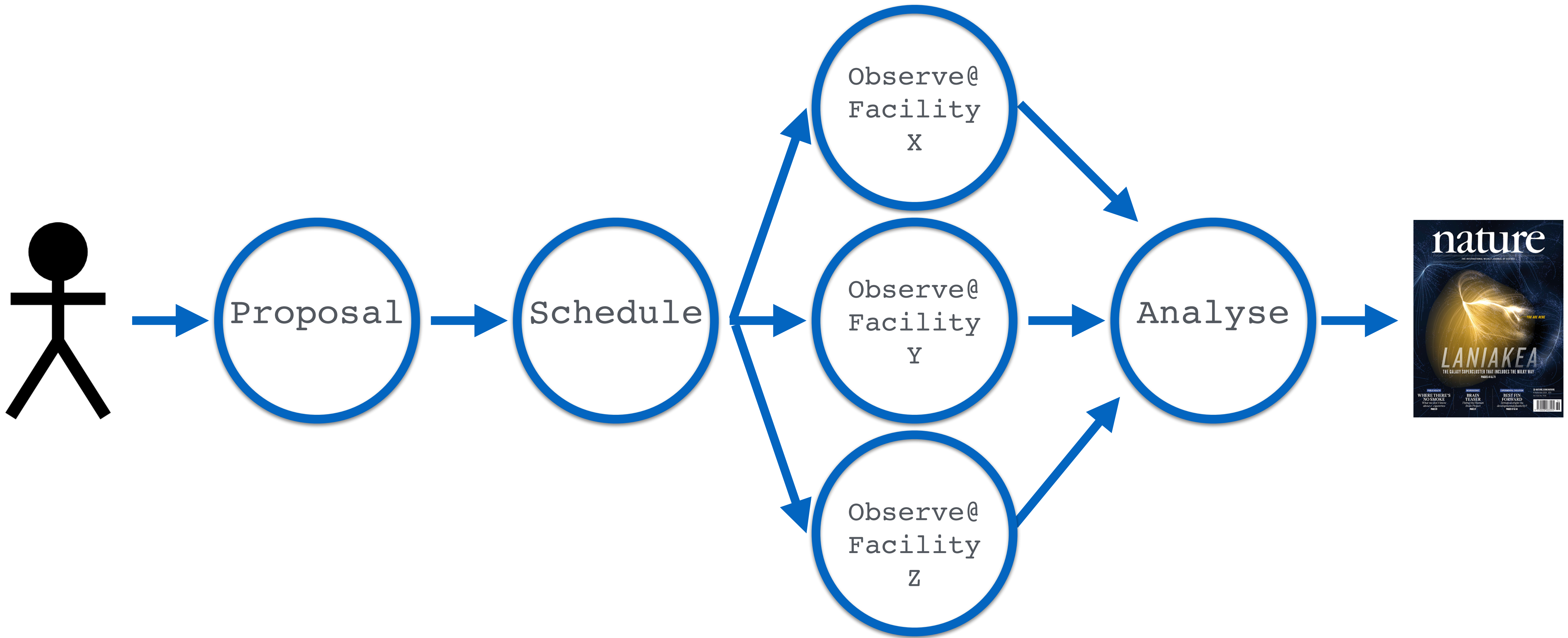
astronomy

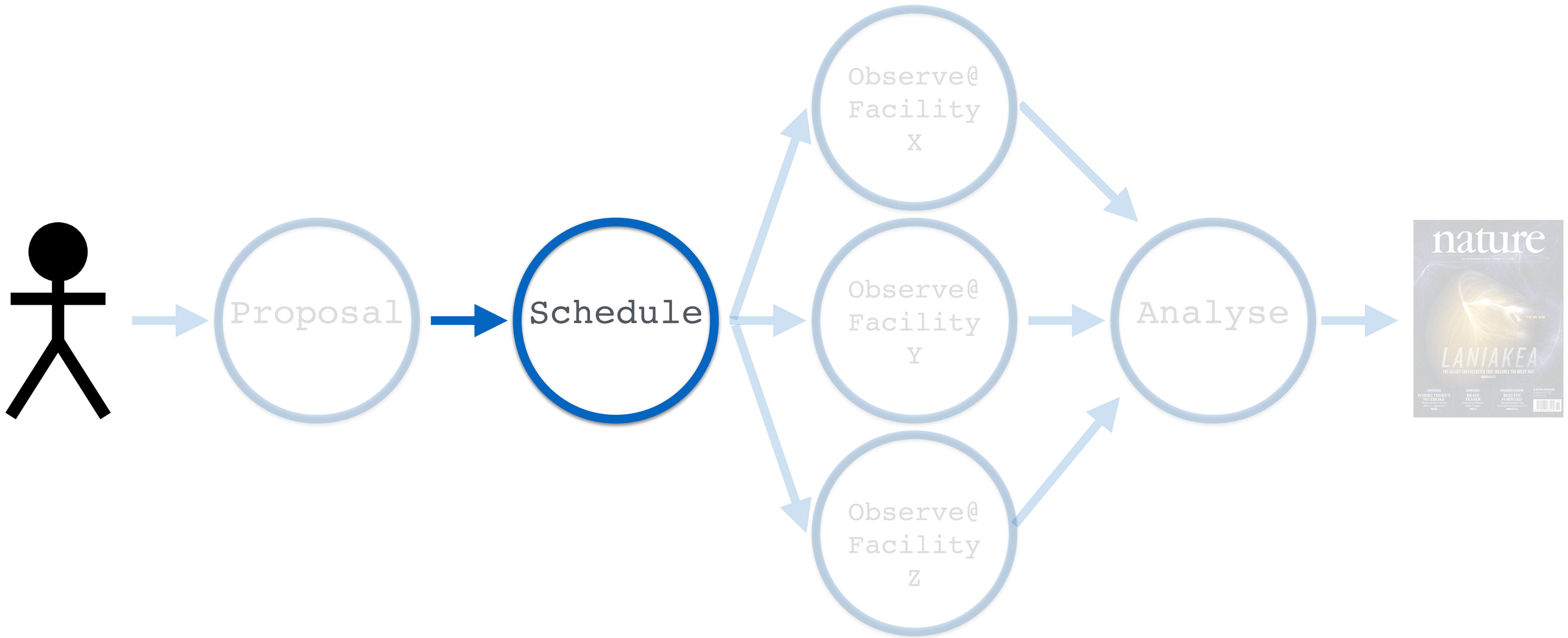
JA2.2 - Access to:

multi-facility*
time domain
astronomy

() multi-wavelength can be implicit through a judicious choice of facilities*

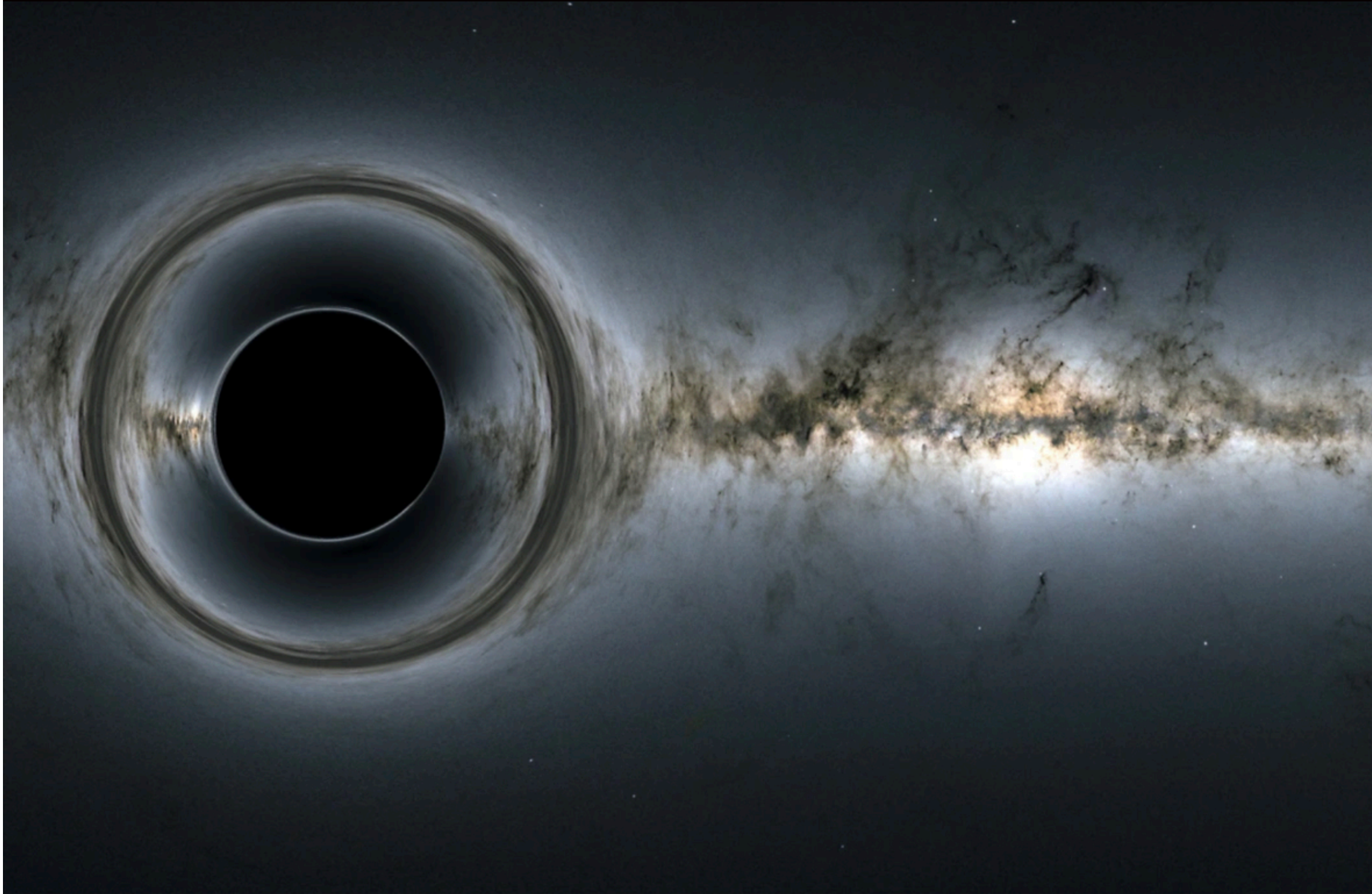






Targets ▾Target GroupingData ▾Observatory

Black Hole TOM 2.0



(Image courtesy of NASA/ESA/Gaia/DPAC)

Latest Comments

No comments yet.

Latest Targets

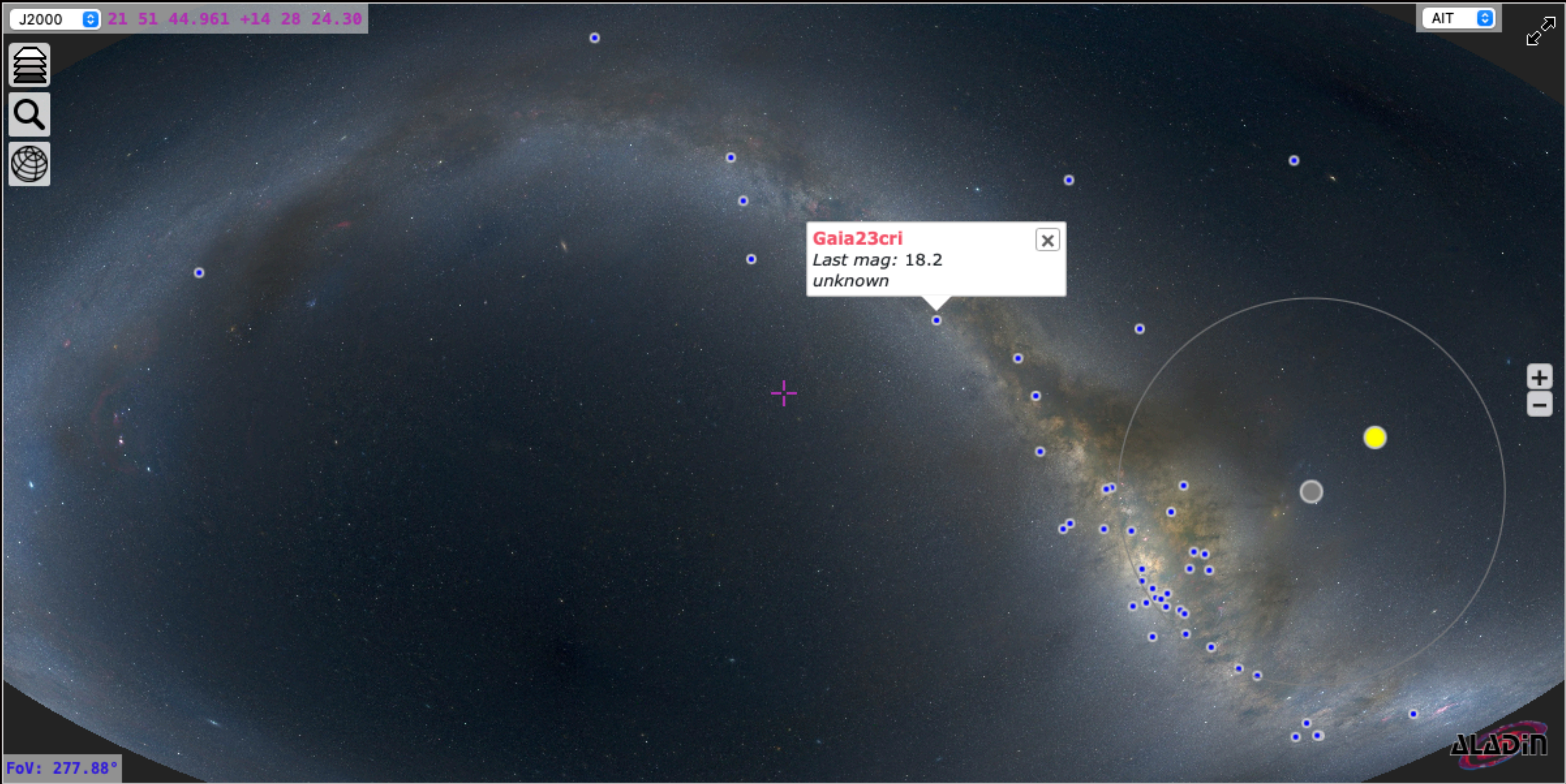
ID	Created
No targets. Create a target.	

`https://bh-tom2.astrolabs.pl/`

51 Targets

Create Targets

Export Filtered Targets



Add/Remove from grouping

old targets not for observing

Add

Move

Remove

Show 10 entries

	Names	RA	Dec	Nobs	Last Gmag	Last Filter	Importance	Created	Priority	Sun	Class
	Gaia21fkl	07:46:28.378	-21:52:32.016	4757	14.5	I(GaiaSP)	1.0	2023-10-03 02:10:39	0.0	102	Microlensing Event
	Gaia23ckg	17:43:04.003	-35:15:32.400	377	17.9	~G	1.0	2023-10-25 17:10:19	7.5	38	Microlensing Event
	B1721+343	17:23:20.809	+34:17:58.480	1520	15.4	~G	1.0	2023-11-13 15:11:28	95.5	61	Quasar(QSO)

51 Targets

Create Targets

Export Filtered Targets

AIT

+

-

8.2

X

move from grouping

old targets not for observing

Add

Move

Remove

	Importance	Created	Priority	Sun	Class
		2023-10-03 02:10:39	0.0	102	Microlensing Event
		2023-10-25 17:10:19	7.5	38	Microlensing Event
		2023-11-13 15:11:28	95.5	61	Quasar(QSO)

Names

SN2016ezh

Gaia24azi

OJ287

Gaia23dfy

Gaia19cuu

Gaia23dss

Gaia24bfr

Gaia24amk

Gaia23dnm

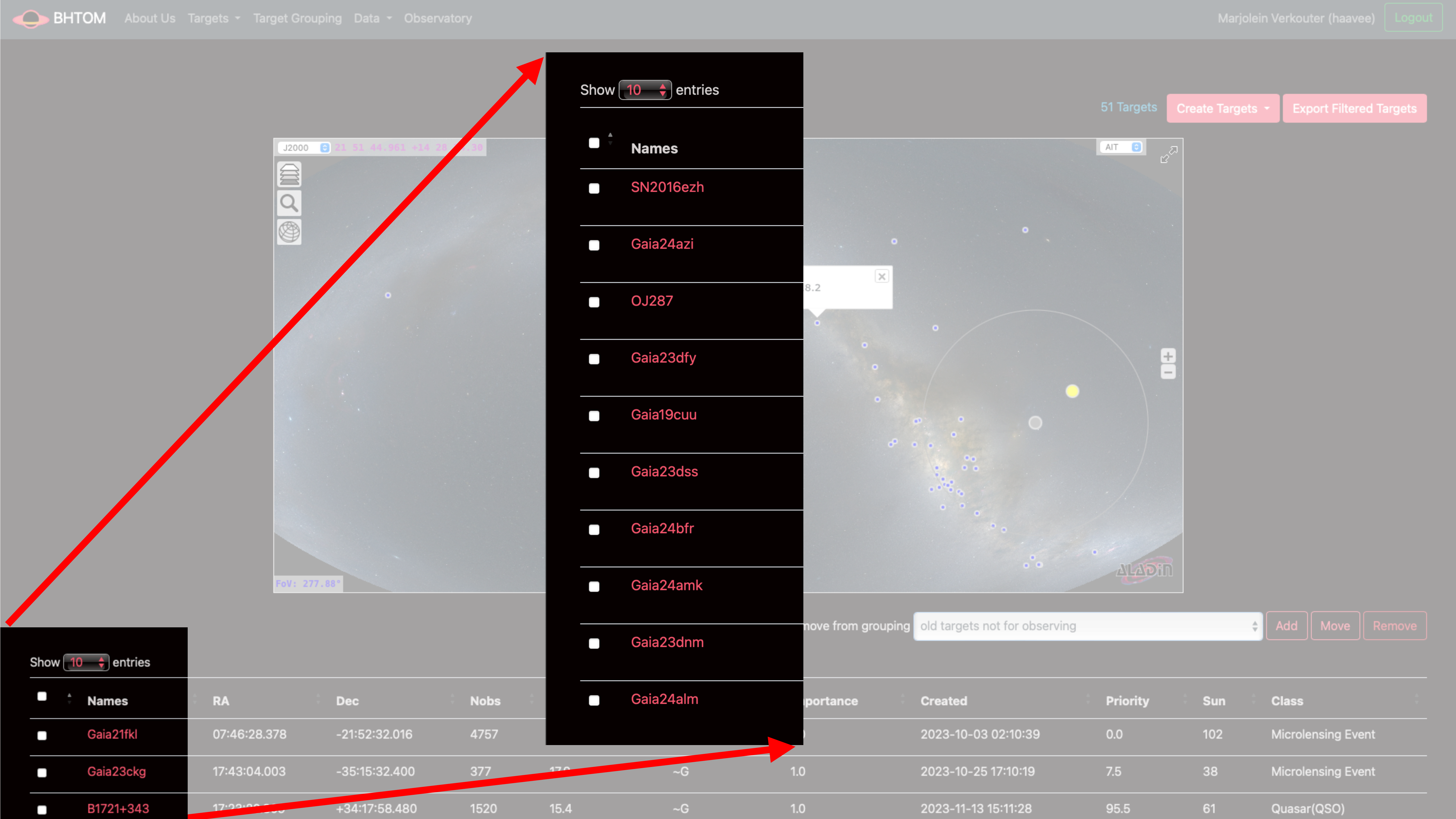
Gaia24alm

Names

Gaia21fkl

Gaia23ckg

B1721+343



Gaia24

Name

Ra,Dec

Galactic (l,b)

Constellation

Discovered

Class

Description

Phot.Class

Last MJD

Last Mag/Filter

Target importance (0-10)

Cadence requested (d)

Observing priority

Sun Separation (deg)

Other names:

ANTARES

ZTF

Aquila

2024-03-14 05:15:40

Microensing Event

candidate binary microlensing
event, now gone

YSO 47.8%

60591.39382756

21.3 mag [~G]

1.0

5.0

0.3

92.0

Photometry

Models

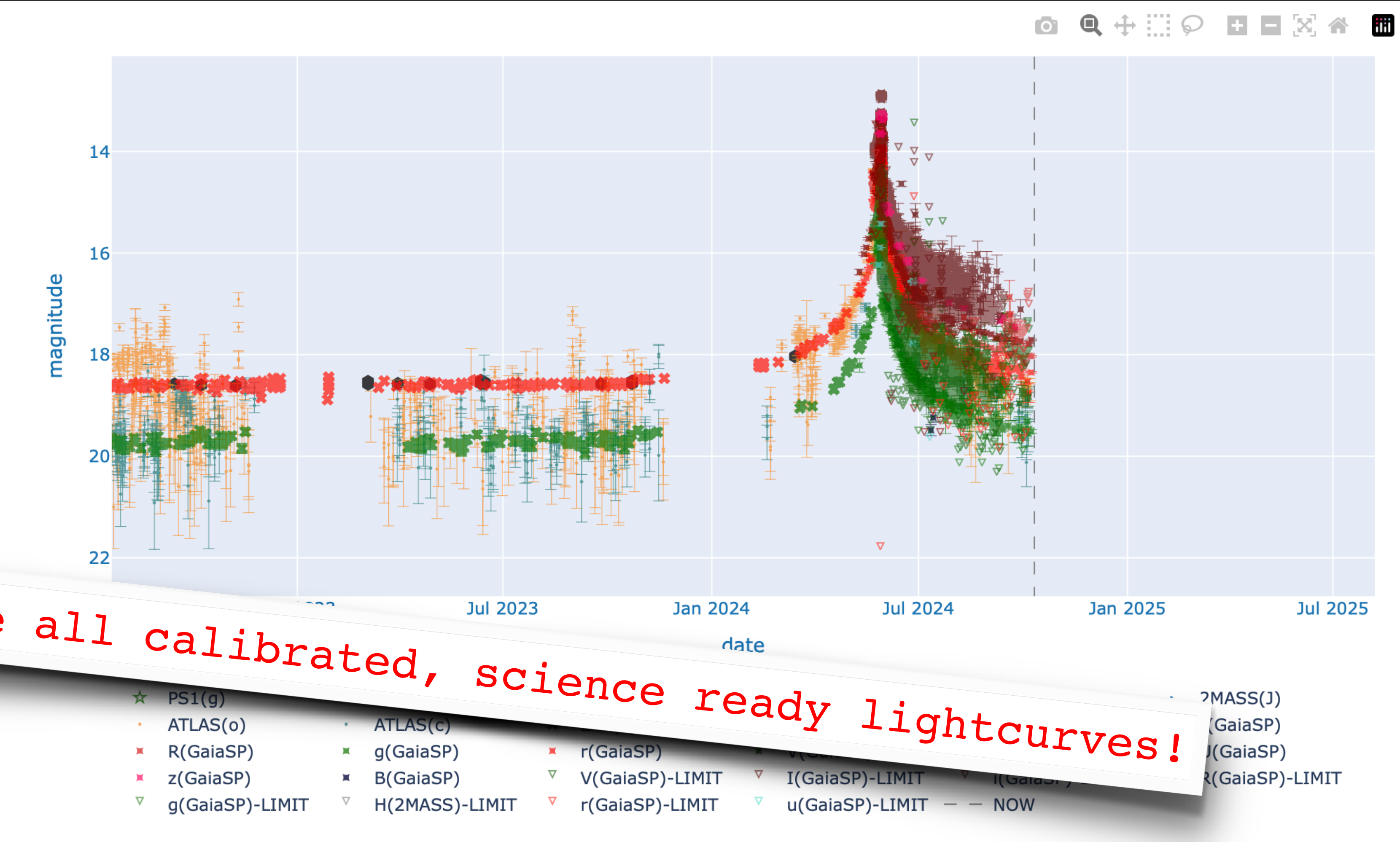
Spectroscopy

Publication

Manage Data

Manage Groups

Photometry



Download photometry data

Download radio data

Photometry (per facility)



Las Cumbres Observatory

Private [iptables - Fail2ban](#) [SFXC Fringes](#) [JIVE ZoomRoom#1](#) [Private Zoom](#) [SharePoint](#) [JIVE](#) [SharePoints](#) [Zooms](#) [Slacks](#)

 **Explore LCO**  **Education & Outreach**  **Science**  **For Observers** 

TOM Toolkit



TOM TOOLKIT 

**Target and Observation Manager
software for powerful, programmable
control of astronomical observing
programs**

Astronomical surveys are producing ever increasing catalogs of new discoveries at ever faster rates and astronomers have found it necessary to build database-driven systems to handle the flood of information on their targets, observations, and data products - Target and Observation Managers or TOMs.

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<https://lco.global/tomtoolkit/>

#PB - C4KTY82

RequestGroup # 1127853

State

Updated

Submitted

Proposal

Submitter

IPP

Observation Type

Options

✓ COMPLETED

2021-01-13 14:09:39

2021-01-12 22:57:05

FTPEPO2014A-004

paul_breitenstein1

1.050000

NORMAL

Sub-requests #2355340

2355340

⌚ Duration: 672 seconds

📷 Instrument: 2.0 meter Spectral

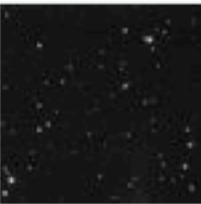
✓ COMPLETED

📋 Acceptability Threshold: 60.0%

📅 2021-01-13 14:09:39

🔗 View in API

📄 Download



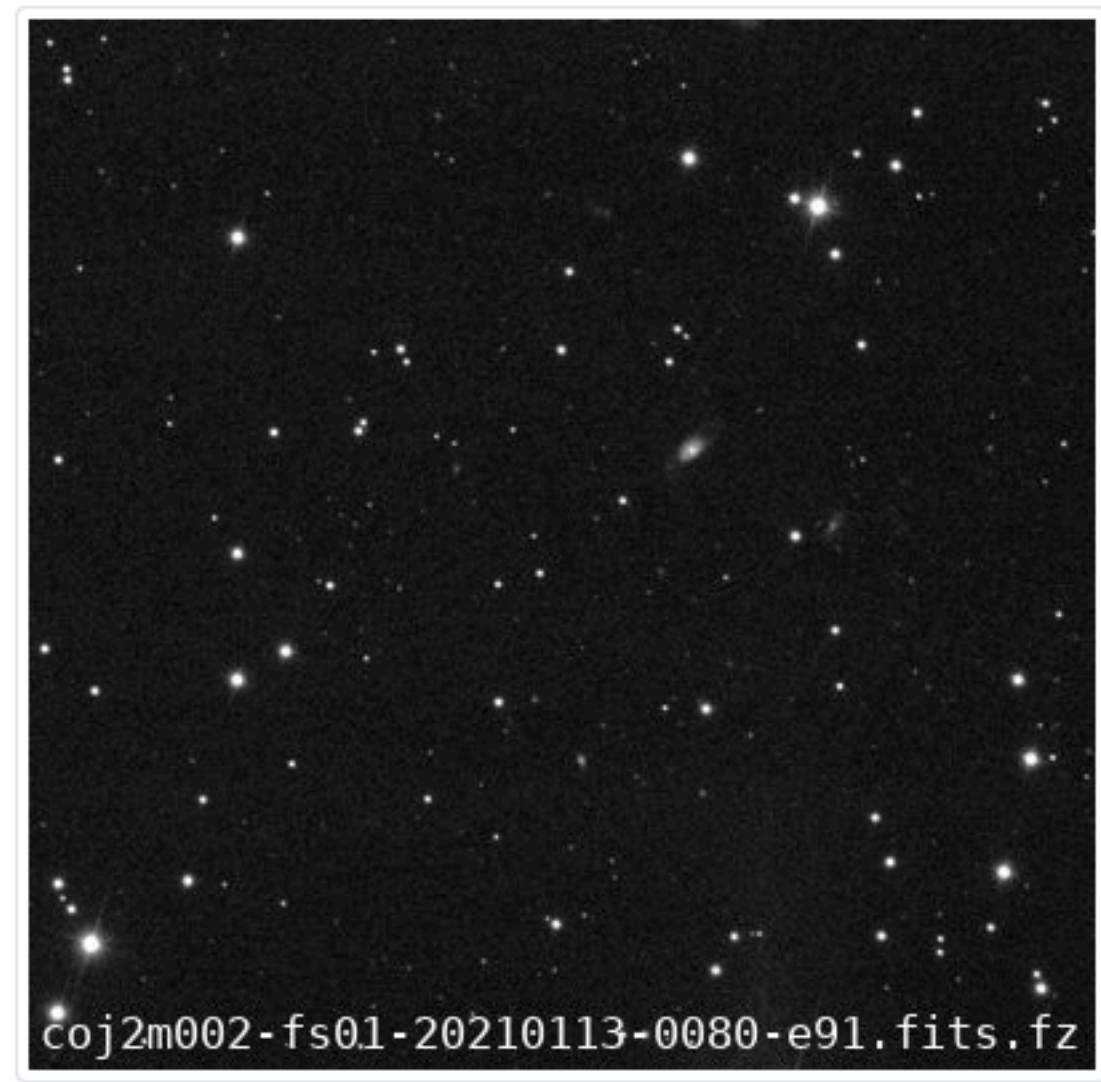
Details

Scheduling

Visibility

Data

Click a row in the data table to preview the file below. Click preview for a larger version.

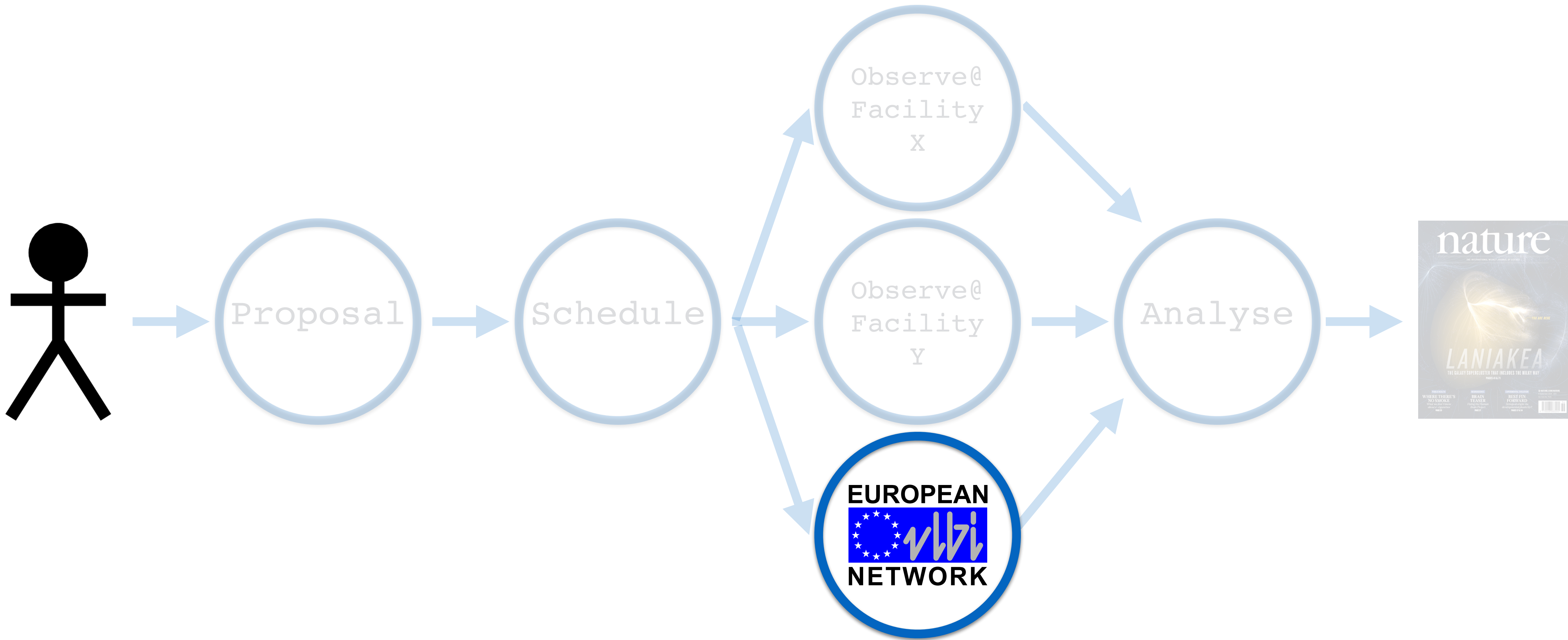


✓ Download Selected

📄 Download All

➔ View on Archive

☐	filename	DATE_OBS	filter	obstype	Reduction
☐	coj2m002-fs01-20210113-0075-e00.fits.fz	2021-01-13 14:01:57	R	EXPOSE	raw
☐	coj2m002-fs01-20210113-0075-e91.fits.fz	2021-01-13 14:01:57	R	EXPOSE	reduced
☐	coj2m002-fs01-20210113-0076-e00.fits.fz	2021-01-13 14:03:14	R	EXPOSE	raw
☐	coj2m002-fs01-20210113-0076-e91.fits.fz	2021-01-13 14:03:14	R	EXPOSE	reduced
☐	coj2m002-fs01-20210113-0077-e00.fits.fz	2021-01-13 14:04:30	R	EXPOSE	raw
☐	coj2m002-fs01-20210113-0077-e91.fits.fz	2021-01-13 14:04:30	R	EXPOSE	reduced
☐	coj2m002-fs01-20210113-0078-e00.fits.fz	2021-01-13 14:05:47	R	EXPOSE	raw
☐	coj2m002-fs01-20210113-0078-e91.fits.fz	2021-01-13 14:05:47	R	EXPOSE	reduced
☐	coj2m002-fs01-20210113-0079-e00.fits.fz	2021-01-13 14:07:05	R	EXPOSE	raw
☐	coj2m002-fs01-20210113-0079-e91.fits.fz	2021-01-13 14:07:05	R	EXPOSE	reduced



```
$> cat SIMPLE_CONT.KEY
```

```
setini = {SETINI SETUP NAME} /  
  band    = '{BAND}'  
  nchan   = {NUMBER CHANNELS}  
  bits    = {NUMBER BITS}  
  bbfilter = {BANDWIDTH}  
  pol     = {POLARIZATION PRODUCTS}  
  pcal    = 'off'  
  freqref = {4966.49}  
  freqoff = {0,0,0,0, 16,16,16,16, 32,32,32,32, 48,48,48,48}  
  netside = {L,L,U,U,  L, L, U, U,  L, L, U, U,  L, L, U, U}  
  /  
endset  
/
```

```
setup = {SETINI SETUP NAME}  
{SCANS}
```

```
!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!  
! Simple continuum observation
```

```
source={FRINGE-FINDER} gap=0:00 dur={5:00} /
```

```
source={TARGET} gap={3:00} dur={3:00} /  
group 1 rep {X}  
  source={TARGET} gap={0:30} dur={14:30} /
```

```
source={FRINGE-FINDER} gap={3:00} dur={5:00} /
```

```
$> pySCHED -f SIMPLE_CONT.KEY
```

```
...
```

```
$> cat SIMPLE_CONT.KEY
```

```
setini = {SETINI SETUP NAME} /  
  band   = '{BAND}'  
  nchan  = {NUMBER CHANNELS}  
  bits   = {NUMBER BITS}  
  bbfilter = {BANDWIDTH}  
  pol    = {POLARIZATION PRODUCTS}  
  pcal   = 'off'  
  freqref = {4966.49}  
  freqoff = {0,0,0,0, 16,16,16,16, 32,32,32,32, 48,48,48,48}  
  netside = {L,L,U,U,  L, L, U, U,  L, L, U, U,  L, L, U, U}  
  /  
endset  
/
```

```
setup = {SETINI SETUP NAME}  
{SCANS}
```

```
!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!!  
! Simple continuum observation
```

```
source={FRINGE-FINDER} gap=0:00 dur={5:00} /
```

```
source={TARGET} gap={3:00} dur={3:00} /  
group 1 rep {X}  
  source={TARGET} gap={0:30} dur={14:30} /
```

```
source={FRINGE-FINDER} gap={3:00} dur={5:00} /
```

```
$> pySCHED -f SIMPLE_CONT.KEY
```

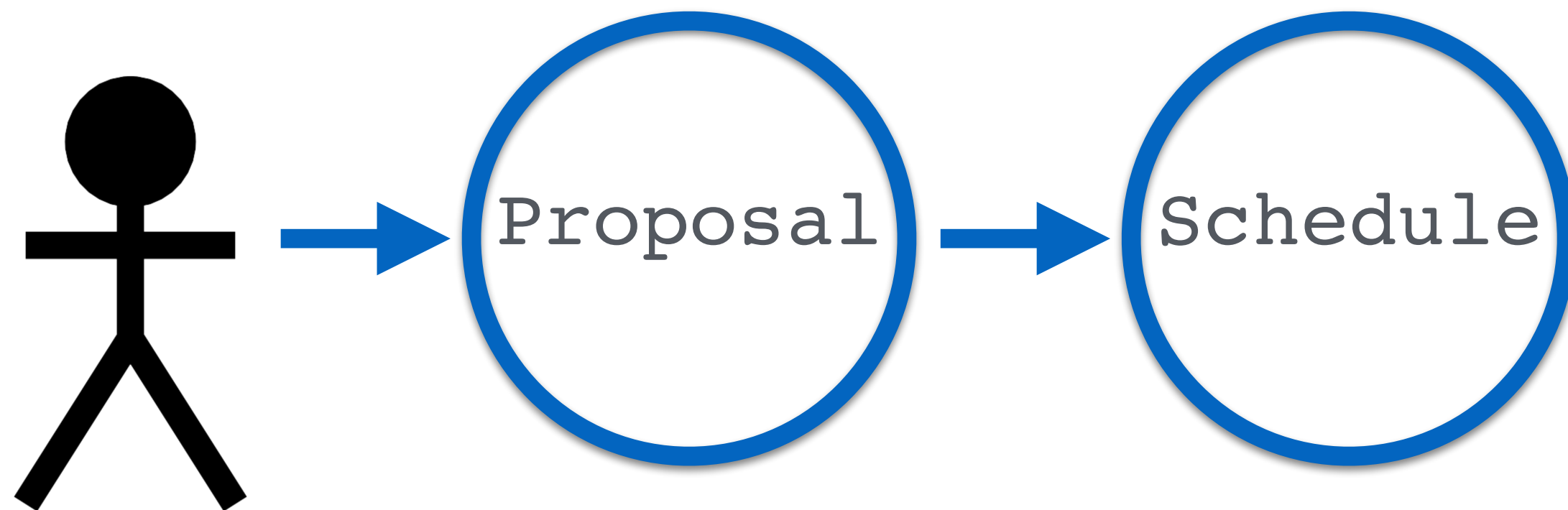
```
...
```

`{this}` means it is a
"Template variable"



JIVE

Joint Institute for VLBI
ERIC

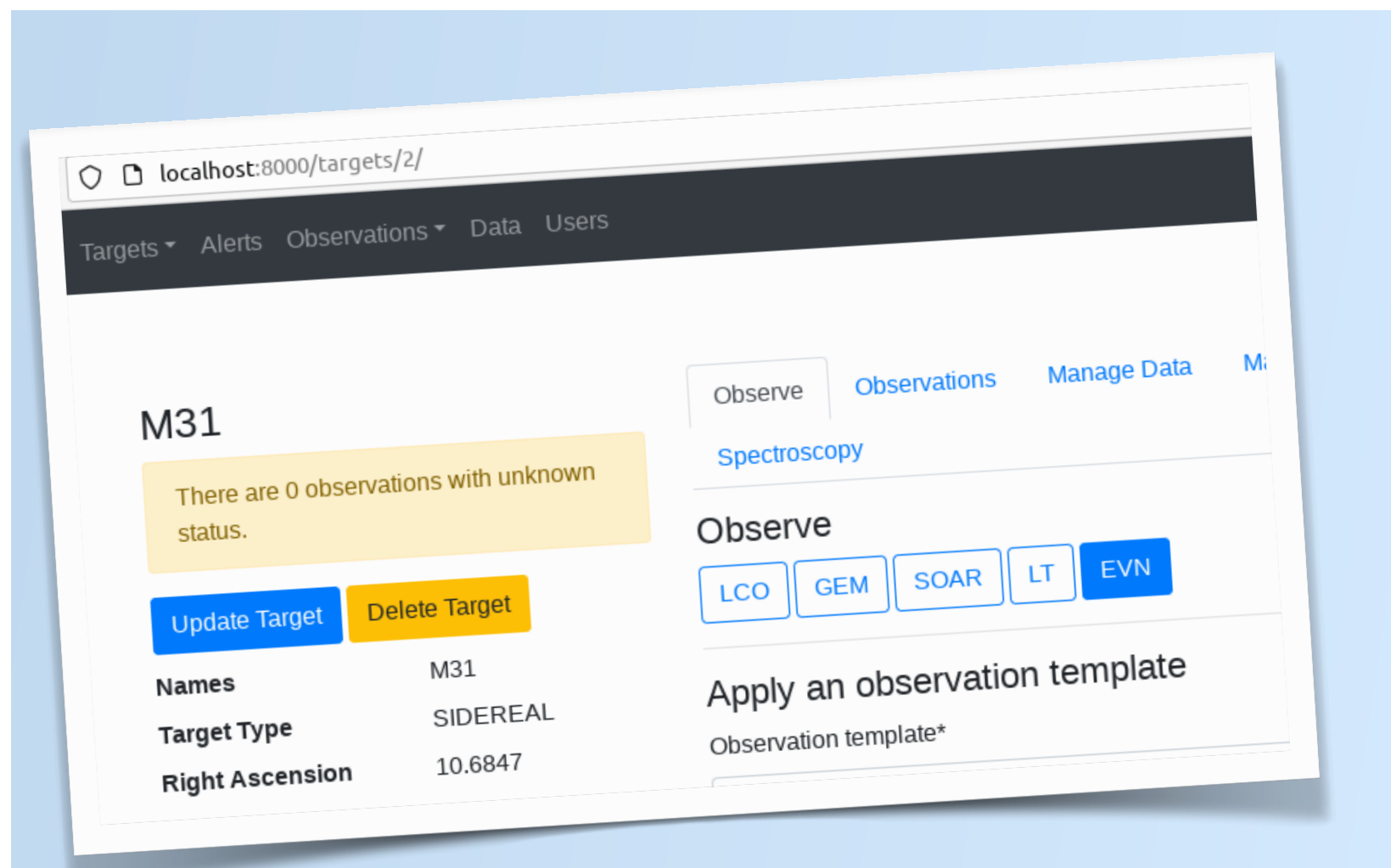


EVN-TOM

```
$> cat INPUT.KEY
setini = {SETINI SETUP NAME} /
band   = '{BAND}'
nchan  = {NUMBER CHANNELS}
bits   = {NUMBER BITS}
bbfilter = {BANDWIDTH}
pol    = {POLARIZATION
PRODUCTS}
pcal    = 'off'
freqref = {4966.49}
freqoff = {0,0,0,0,
16,16,16,16, 32,32,32,32,
48,48,48,48}
netside = {L,L,U,U,  L, L, U,
U,  L, L, U, U,  L, L, U, U}
/
endset
/

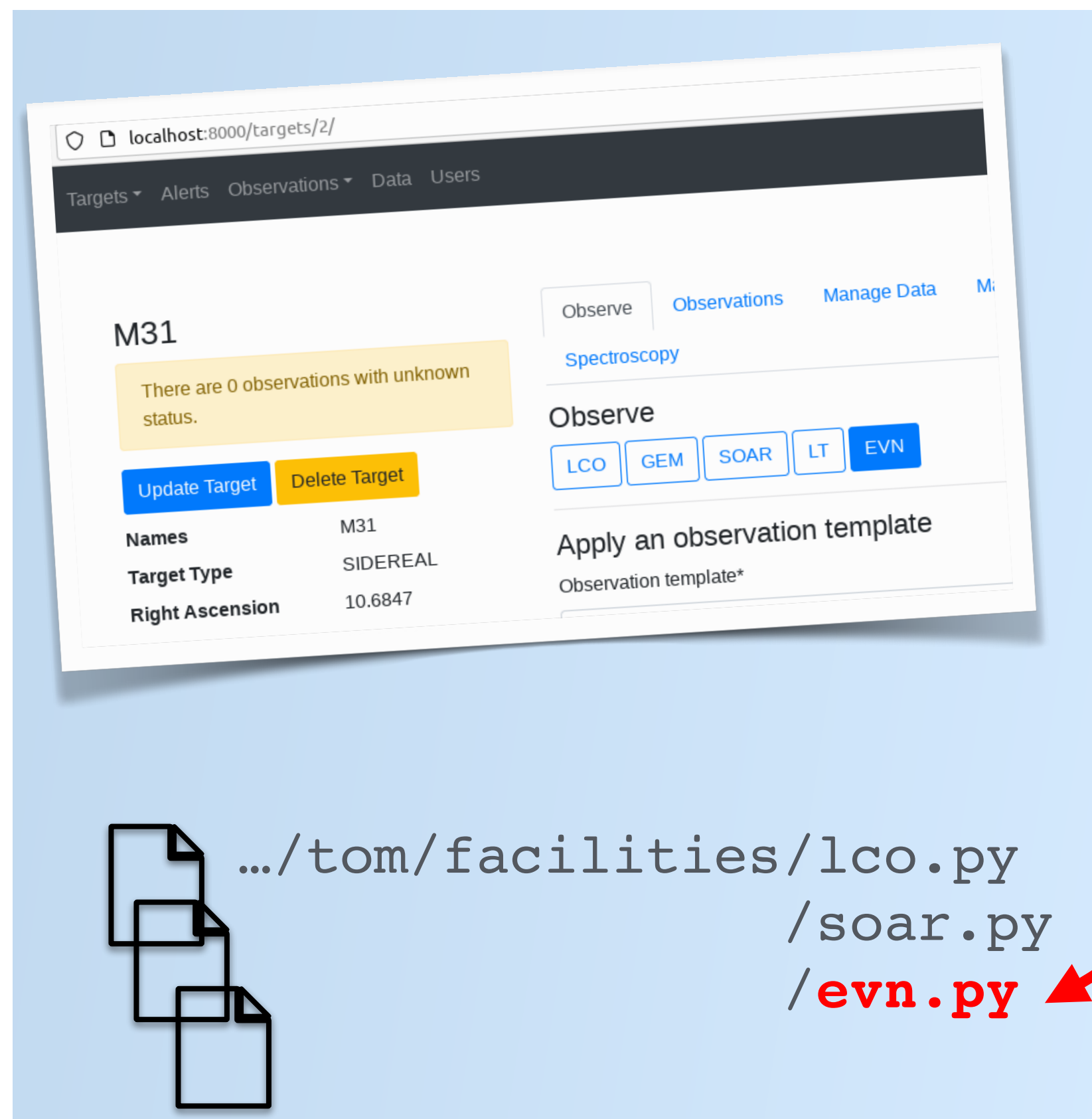
setup = {SETINI SETUP NAME}
{SCANS}

$> sched -f INPUT.KEY
```



.../tom/facilities/lco.py
/soar.py
/**evn.py**

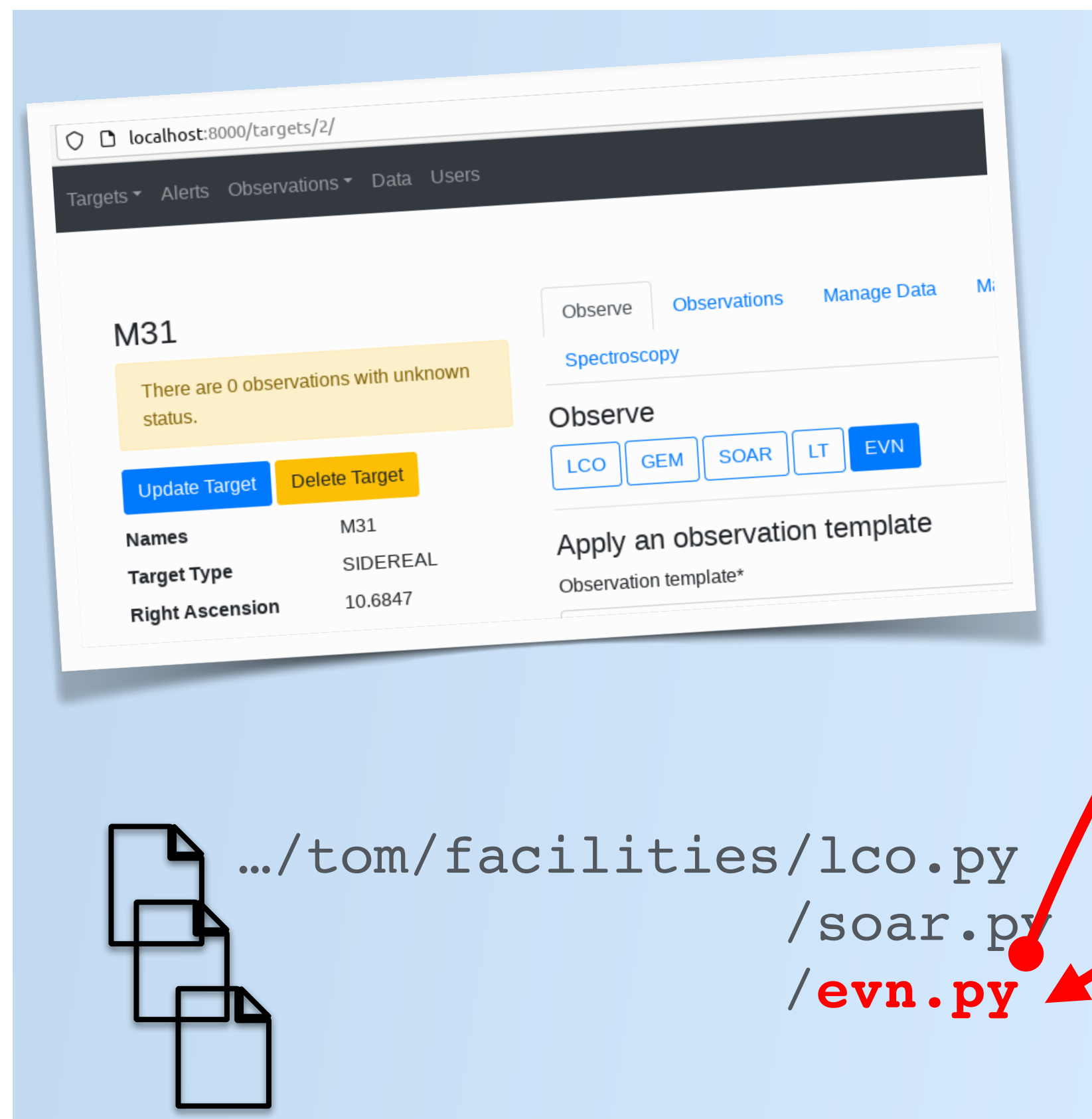
webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>



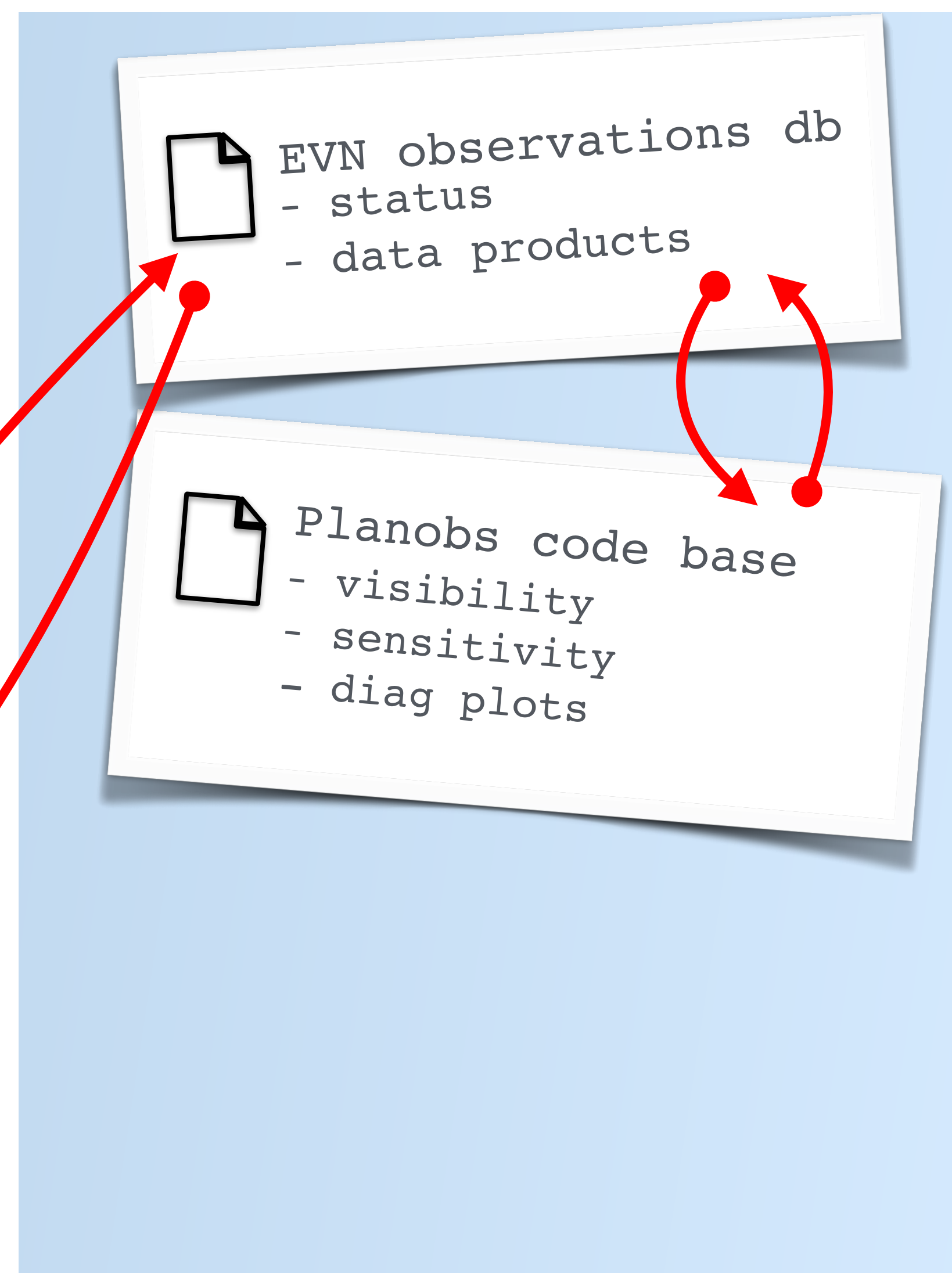
webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>



server running EVN backend
<https://tom-backend.jive.eu/>



webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>



server running EVN backend
<https://tom-backend.jive.eu/>

Private

les - Fai... - Server Fault

SFXC Real-Time Fringe Plot

INFRA-TECH P...oogle Drive

ESCAPE plan... Documenten

Overview - ...REDMINE OSU

VLBI and PFB - Google D

Antenna Selection

Summary

Elevations

UV Coverage

Documentation

COMPUTE OBSERVATION

Observing Band?

cm or 6 GHz

Real-time correlation??

e-EVN mode

Source and Epoch

Do not specify source nor epoch

Find epoch for given source

Define source and epoch

Both the target source and the observing epoch must be specified.

Source (name or coordinates)?

m87

12h30m49.4233823s

+12d23m28.0438581s

% of on-target time?

20 30 40 50 60 70 80 90 100

70% of the total observation.

Datarate per station?

2 Gbps

When is your source visible?

Interactive plots

A single click on one station in the legend will hide/show it. Double-click will hide/show all other antennas. You can also save the plot as png.

The following plot shows the source elevation for the different antennas during the proposed observation. The horizontal solid and dashed lines represent the elevation of 20 and 10 degrees, respectively.

Source elevation during the observation

Elevation (degrees)

Time (GST)

6 7 8 9 10 11 12 13 14 15 16 17 18

80

60

40

20

Effelsberg

Hartebeesthoek

Irbene

Jodrell Bank

Medicina

Noto

Onsala

Sardinia

Tianma

Torun

Westerbork

Resulting (u,v) coverage

150M

100M

50M

0

-50M

-100M

-150M

-200M

-200M

-100M

0

100M

200M

<https://planobs.jive.eu/>

Upload

Saved Data Products

▼

Add to group

Manage groups

Filename	Type	Created	Delete
No saved data for this observation.			

Unsaved data products

Save Selected

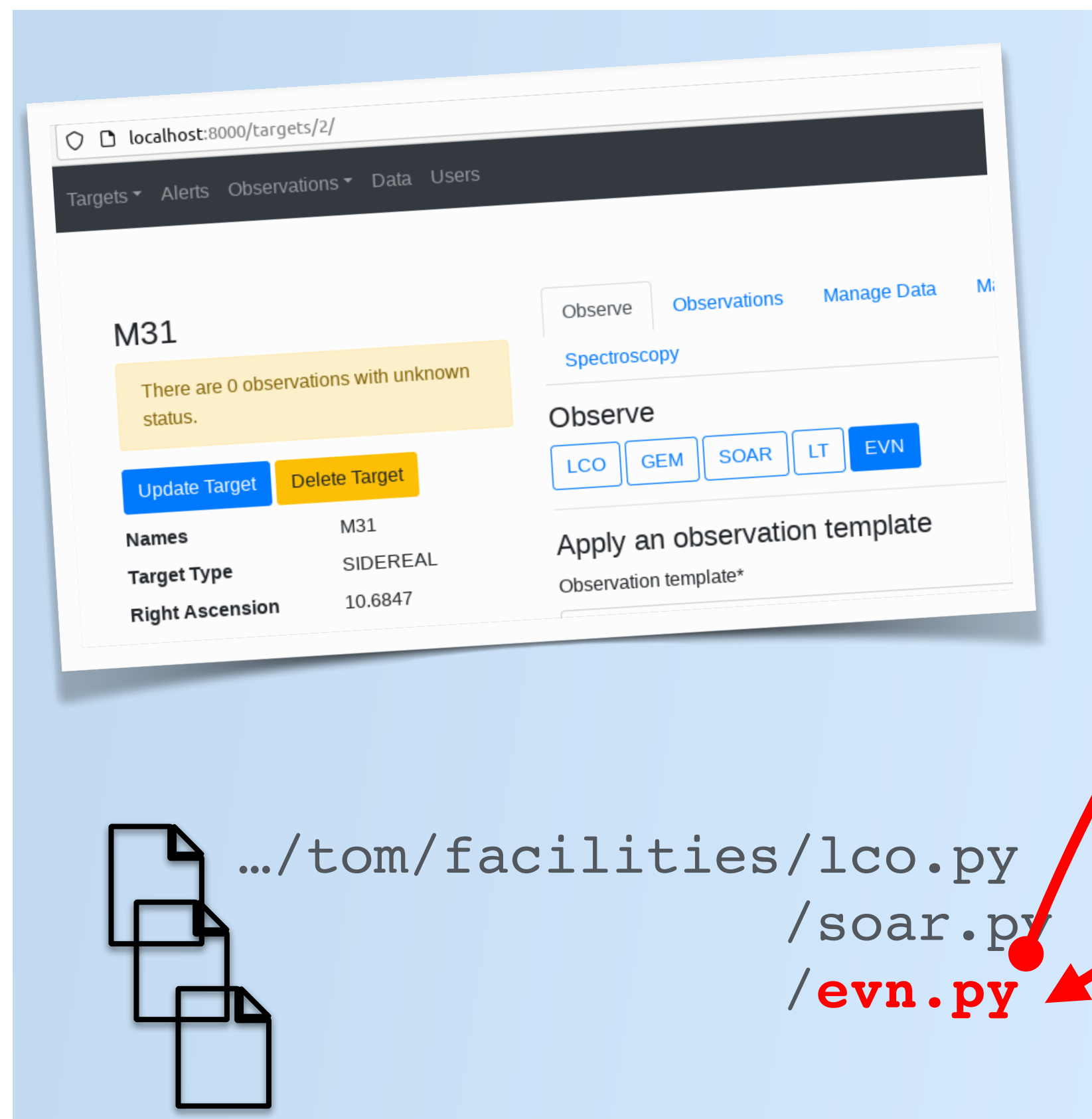
Save All

Created	Filename	Save
	planobs_33.pdf	☐

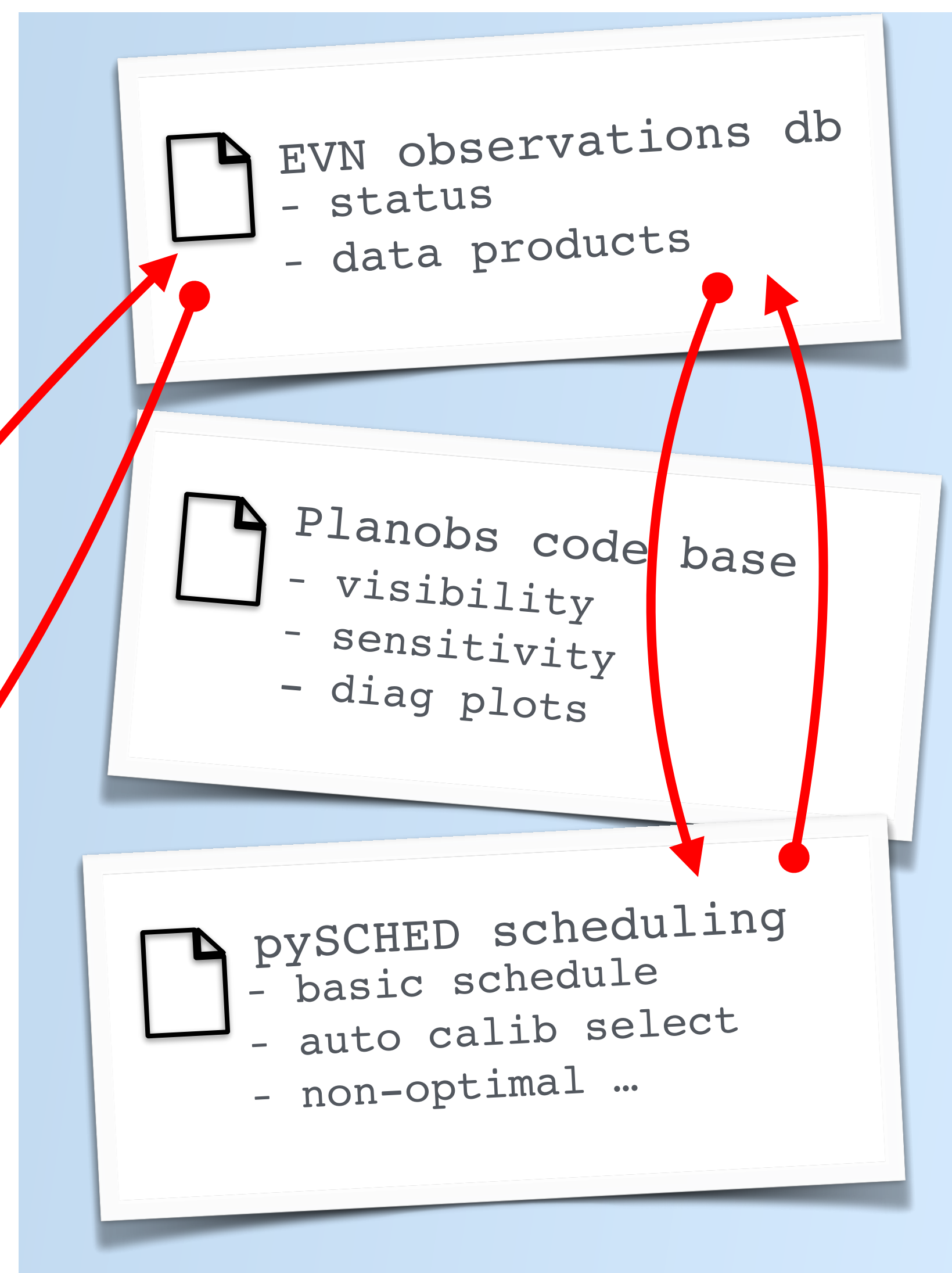
Request Parameters

Create template from request

cadence_strategy	
facility	EVN
observation_type	IMAGING
stations	['Ef', 'Hh', 'Ir', 'Jb', 'Mc', 'Nt', 'On', 'Tr', 'Ys']
band	21cm
start_time	2022-05-26T12:00
end_time	2022-05-26T16:00



webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>



server running EVN backend
<https://tom-backend.jive.eu/>

The EVN array fingerprint #1

“interferometric array”

- implies >> 1 telescope *required*

every telescope *needs* to observe

- same position
- at the same time
- at the same frequency

Log out

Select station

Configure availability

Add availability

Scheduled observations

Overview

Station: Effelsberg

Submit

Select station to manage

Add time range

Log out

Select station

Configure availability

Add availability

Scheduled observations

Overview

Start time: mm / dd / yyyy , -- : -- --

End time: mm / dd / yyyy , -- : -- --

Band: 21 cm or 1.4 GHz

Add

Add available time range to Ef

Configure availability

Log out

Select station

Configure availability

Add availability

Scheduled observations

Overview

Start time

End time

Band

11 / 04 / 2022 , 12 : 00 AM

11 / 11 / 2022 , 12 : 00 AM

21 cm or 1.4 GHz

Remove

11 / 15 / 2022 , 08 : 00 AM

11 / 18 / 2022 , 05 : 00 PM

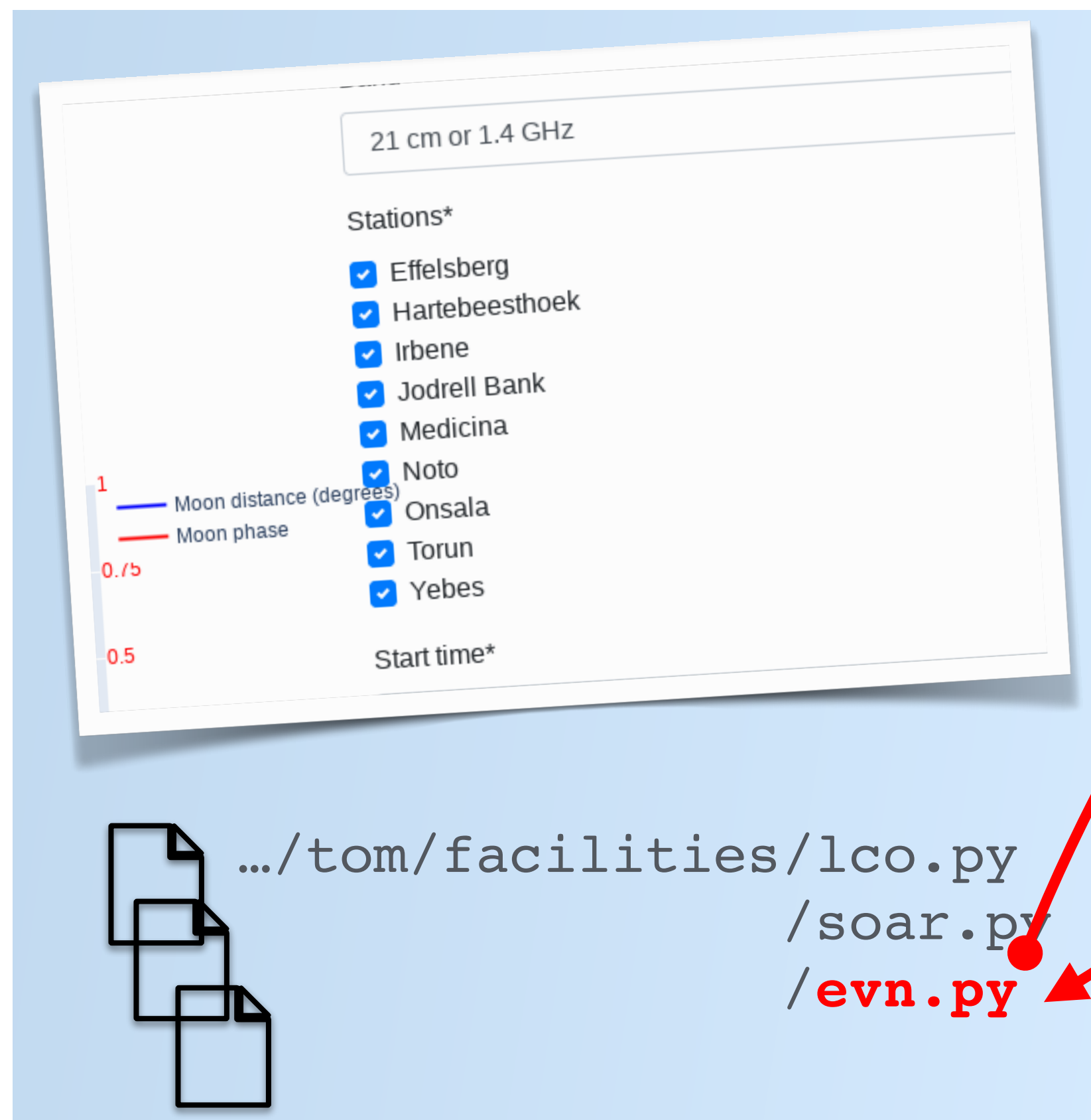
21 cm or 1.4 GHz

Remove

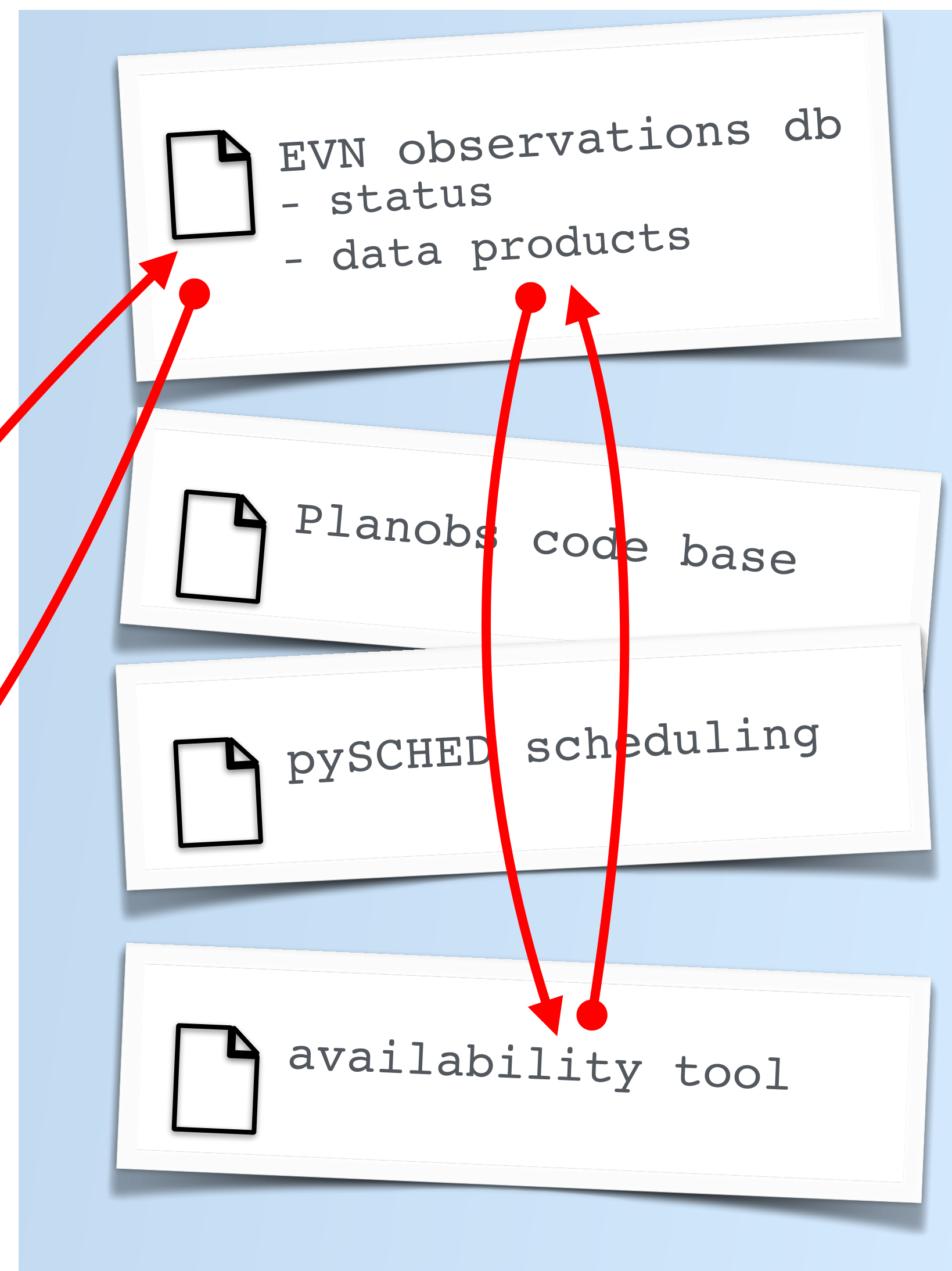
Update

Add availability

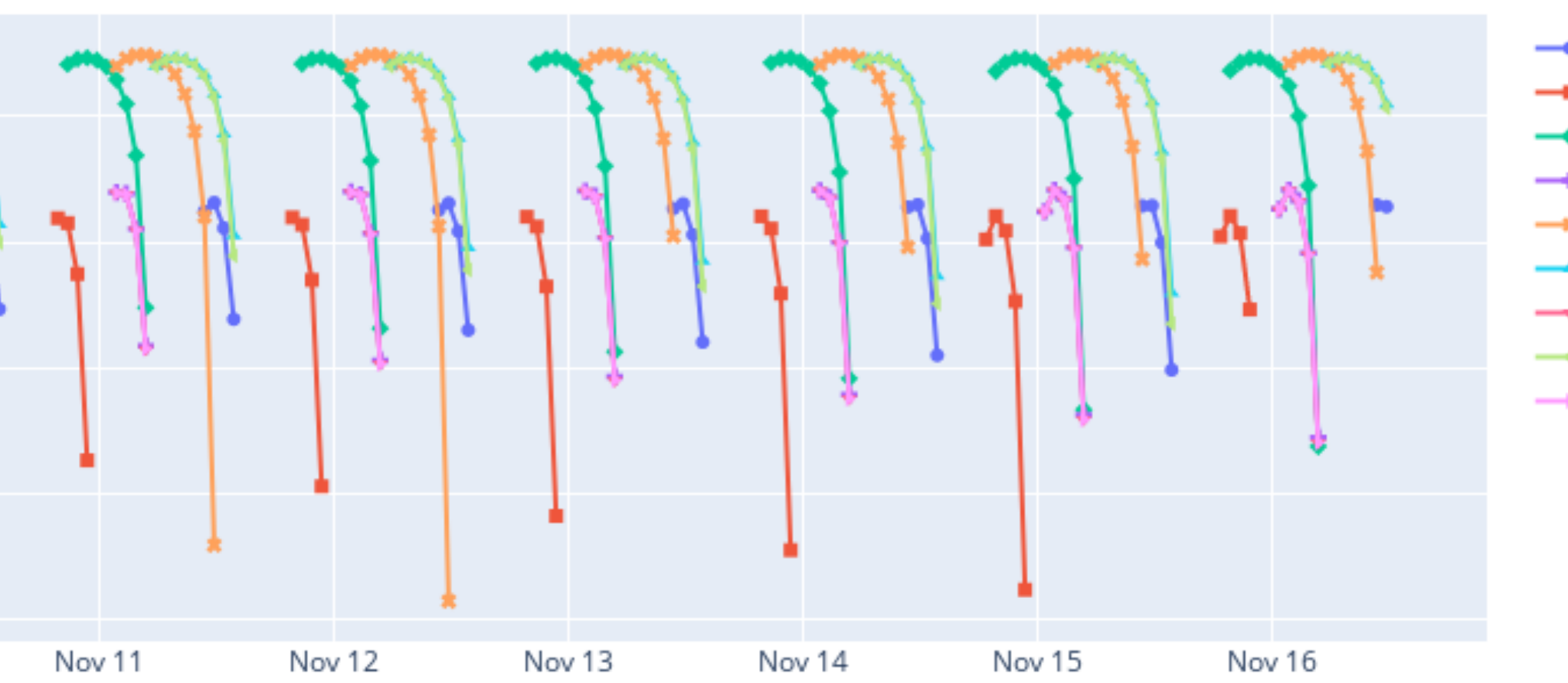
Configure available time ranges for Ef



webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>



server running EVN backend
<https://tom-backend.jive.eu/>



M31
SIDEREAL
10.6847
00:42:44.330
41.2687
+41:16:7.500

Date

Imaging

Band*
21 cm or 1.4 GHz

- Stations*
- ☒ Effelsberg
 - ☒ Medicina
 - ☐ Hartebeesthoek
 - ☐ Onsala

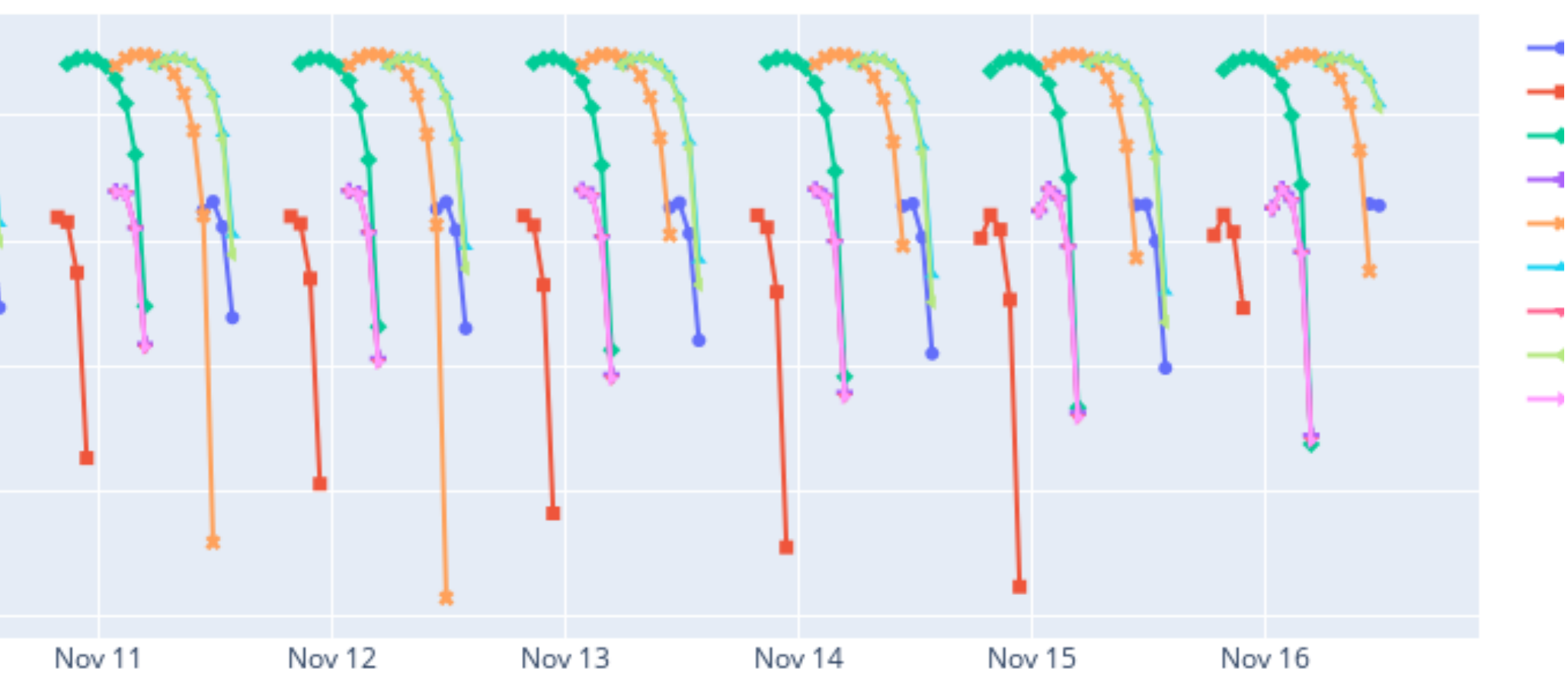
Start time*
11 / 10 / 2022 , 12:00 PM

End time*
11 / 10 / 2022 , 04:00 PM

Check station availability

Submit Back





Date

M31

SIDEREAL

10.6847

00:42:44.330

41.2687

+41:16:7.500

Imaging

Band*

21 cm or 1.4 GHz

Stations*

- ☒ Effelsberg
- ☒ Medicina
- ☐ Hartebeesthoek
- ☐ Onsala

Start time*

11 / 10 / 2022, 12:00 PM

End time*

11 / 10 / 2022, 04:00 PM

Check station availability

Submit

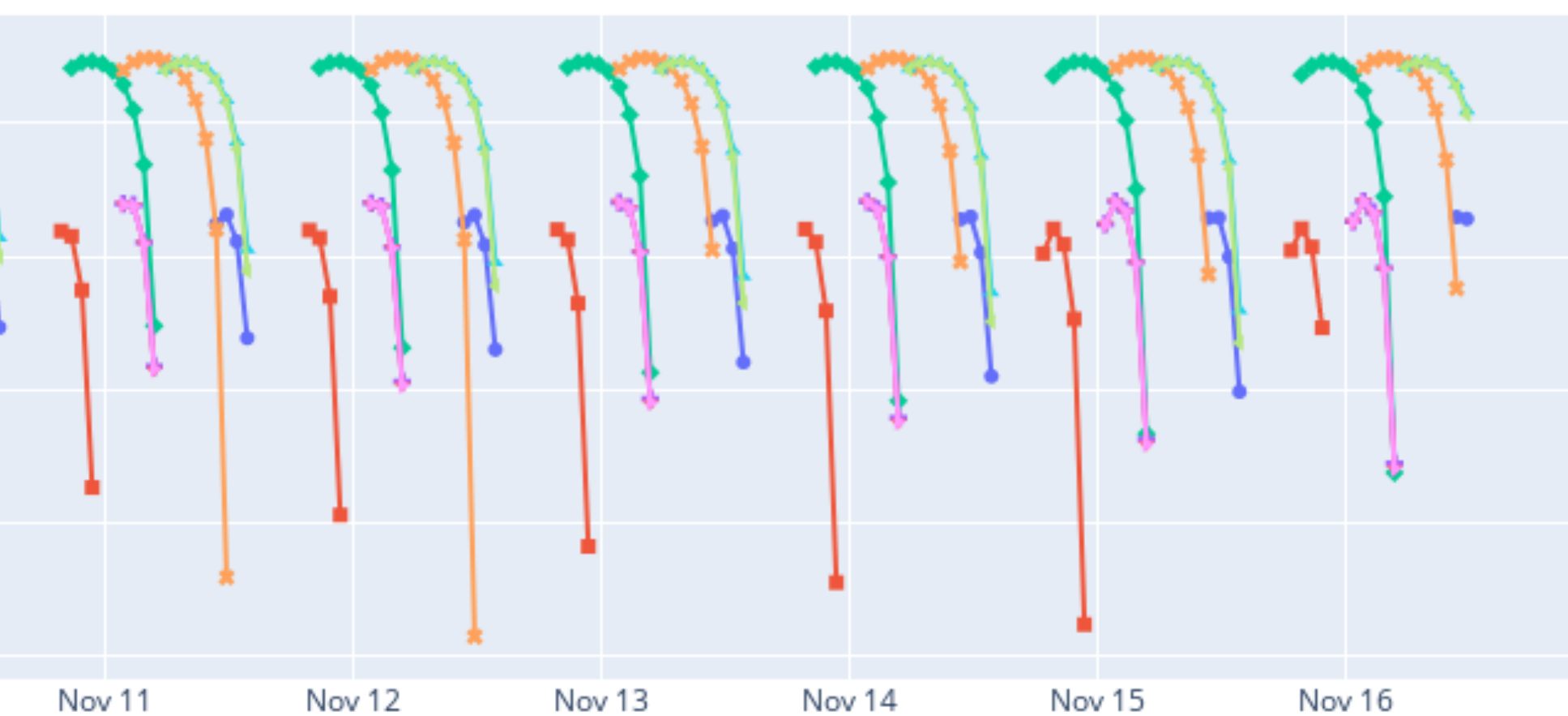
Back



Check station availability

Submit

Back



Date

M31
SIDEREAL
10.6847
00:42:44.330
41.2687
+41:16:7.500

Imaging

Band*

21 cm or 1.4 GHz

Stations*

- ☒ Effelsberg
- ☒ Medicina
- ☐ Hartebeesthoek
- ☐ Onsala

Start time*

11 / 10 / 2022, 12:00 PM

End time*

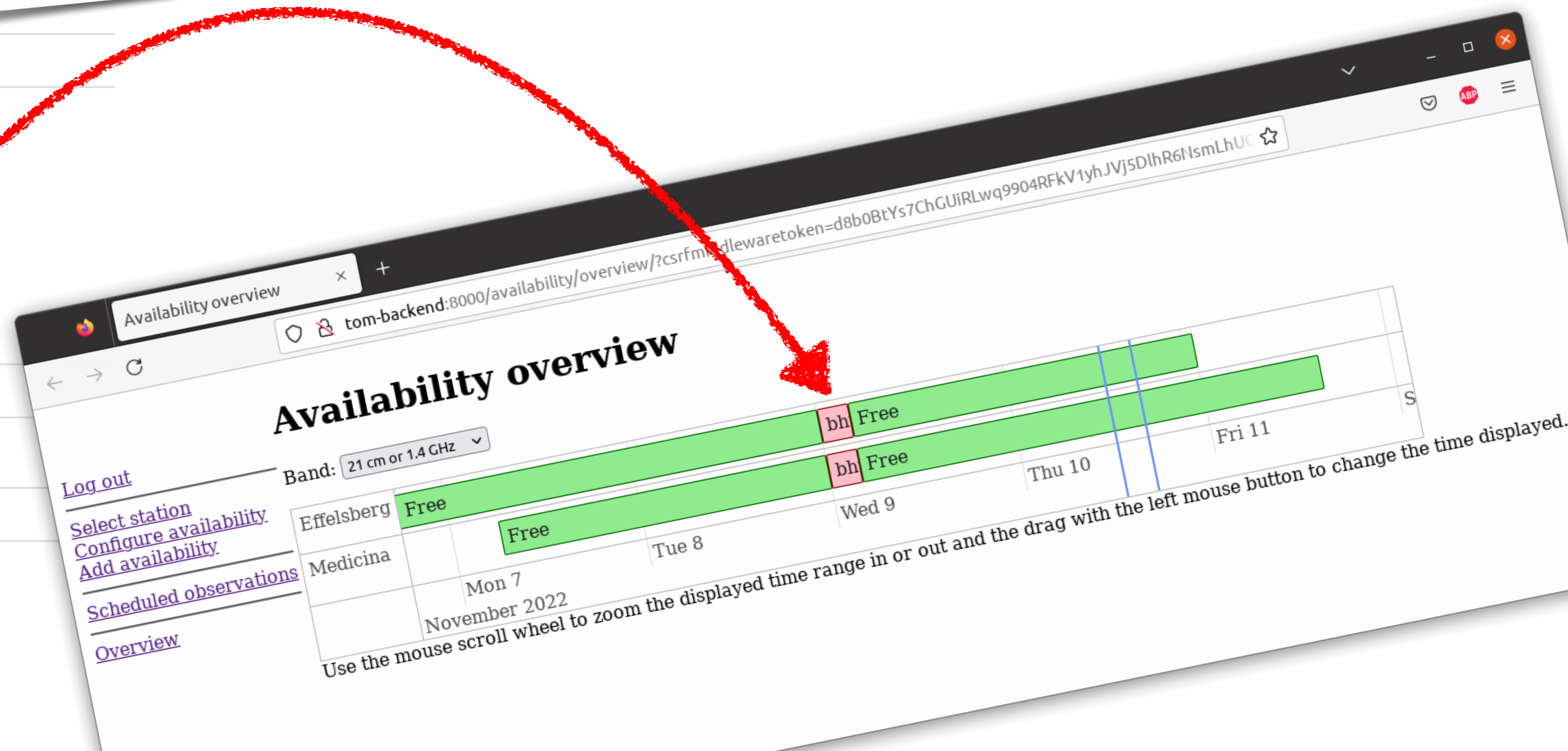
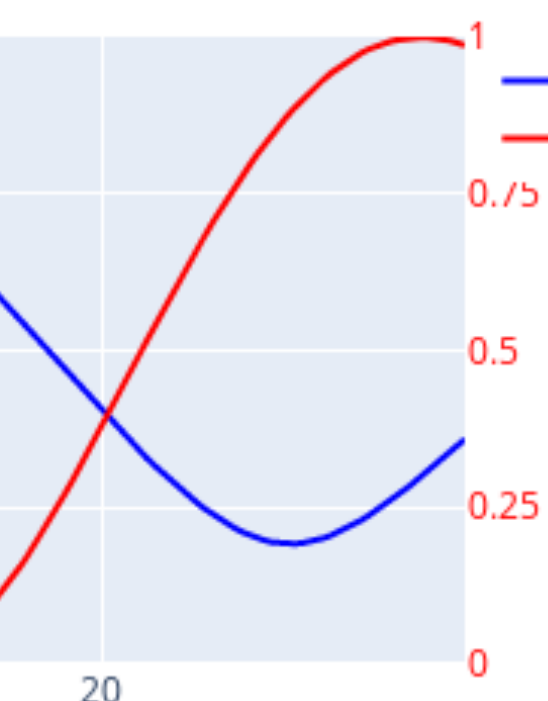
11 / 10 / 2022, 04:00 PM

Check station availability

Submit

Back

this planned observation!



Scheduled observation

tom-backend:8000/scheduler/scheduled/Ef/

Log out

Select station

Configure availability

Add availability

Scheduled observations

Overview

Start time

End time

Band

VEX

2022-11-09 0:00

2022-11-09 4:00

21cm

[bhtom13.vex](#)

Download scheduled observation files for Ef

The EVN array fingerprint #2

The Program Committee (PC)

- time allocation to approved proposals
- approves source lists in proposal

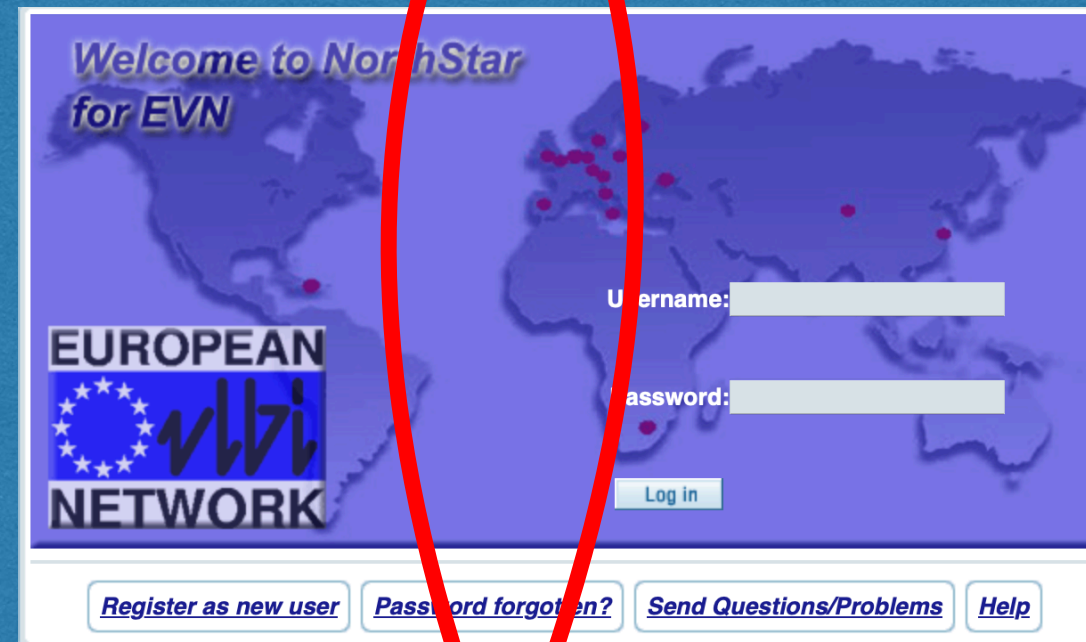
The EVN array fingerprint #2

The Program Committee (PC)

- time allocation to approved proposals
- approves source lists in proposal

OAuth2 Webservice

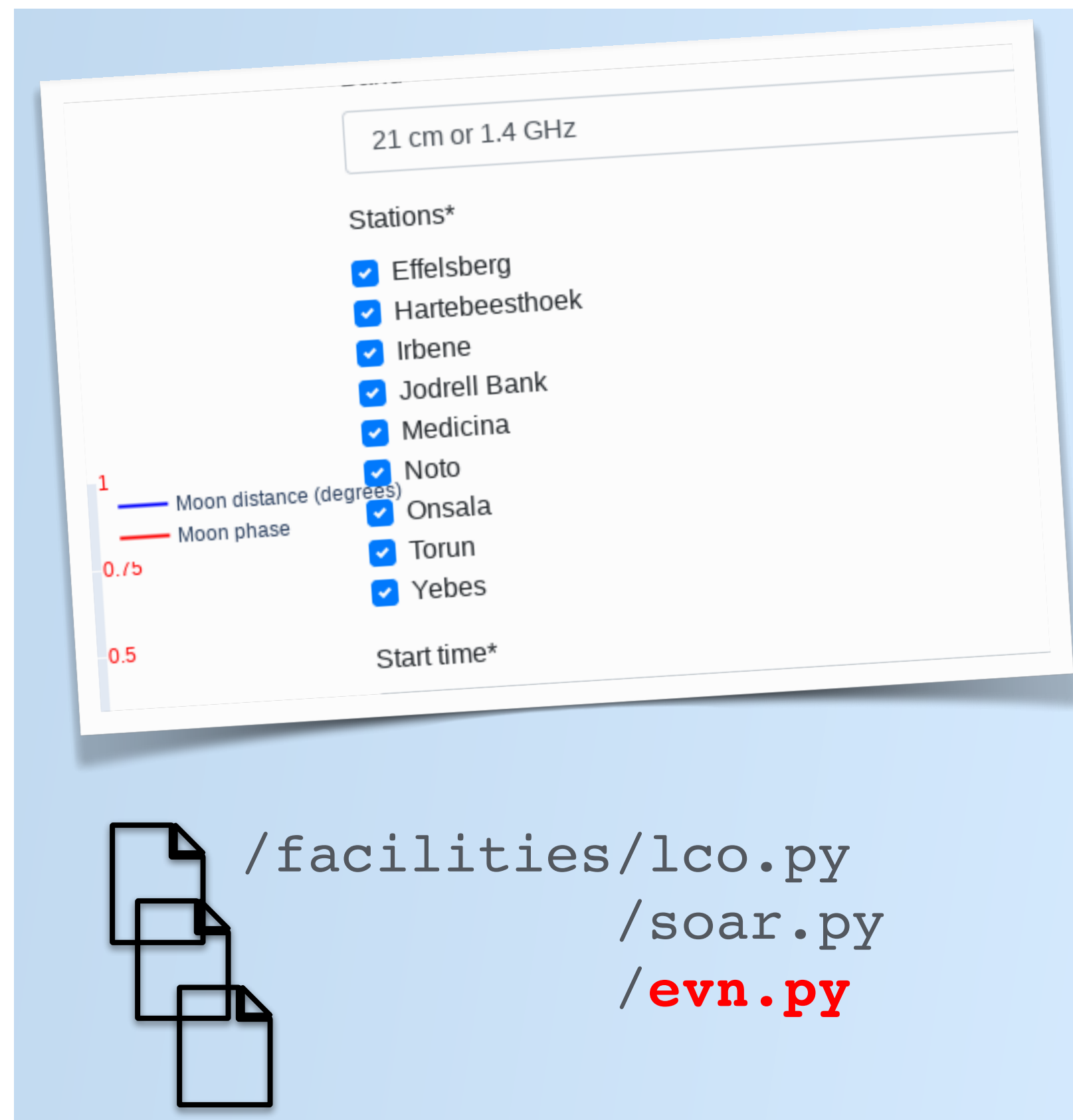
- authorization from **TOM** to **Northstar**
- RESTful API for plugin:
check proposal details



EVN Proposal Submission Tool

- own user administration
- proposal database:
 - sources
 - flux densities
 - phase-reference sources
 - ...

server running EVN PST
<https://proposal.jive.eu/>



webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>

EVN o
- stat
- data

Plan

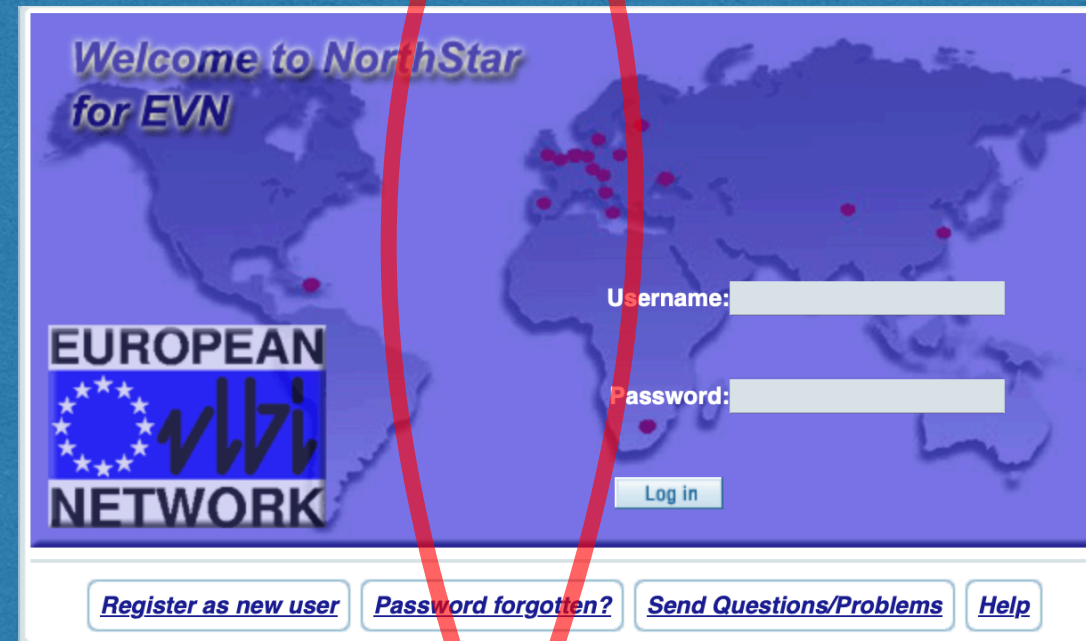
pySCH

avail

server running
<https://tom->

OAuth2 Webservice

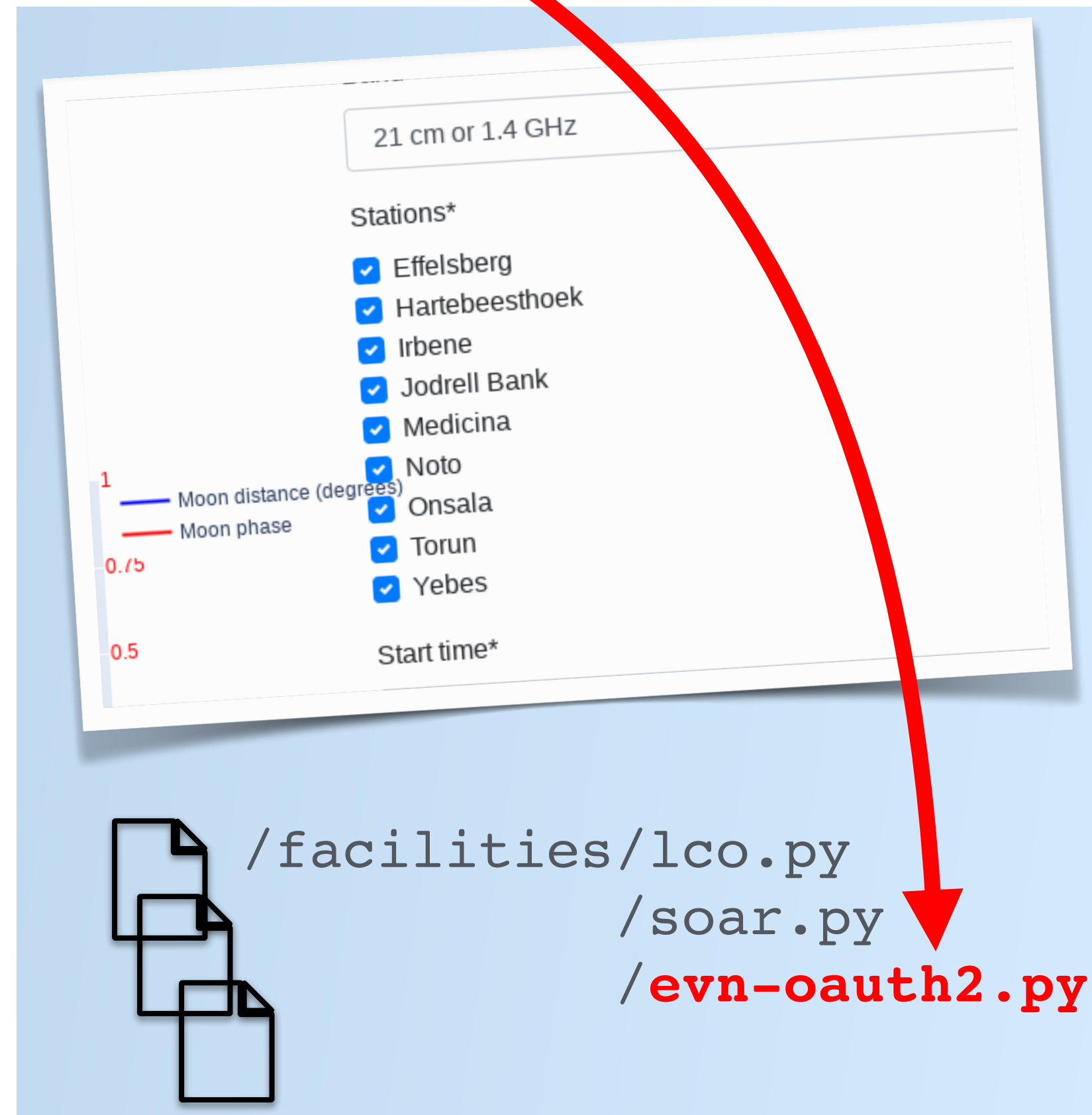
- authorization from **TOM** to **Northstar**
- RESTful API for plugin:
check proposal details



EVN Proposal Submission Tool

- own user administration
- **proposal database:**
 - sources
 - flux densities
 - phase-reference sources
 - ...

server running EVN PST
<https://proposal.jive.eu/>



webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>

EVN o
- stat
- data

Plan

pySCH

avail

server running
<https://tom->

The EVN array fingerprint #3

Automatically generated schedules?

The EVN array fingerprint #3

Automatically generated schedules?

Not really ...

The EVN array fingerprint #3

Calibration scheme is **critical** for succes

- flux calibrator source?
- band-pass calibrator?
- phase-reference source cycle pattern?

The EVN array fingerprint #3

Calibration scheme is critical for succes

- flux calibrator source?
- band-pass calibrator?
- phase-reference source cycle pattern?

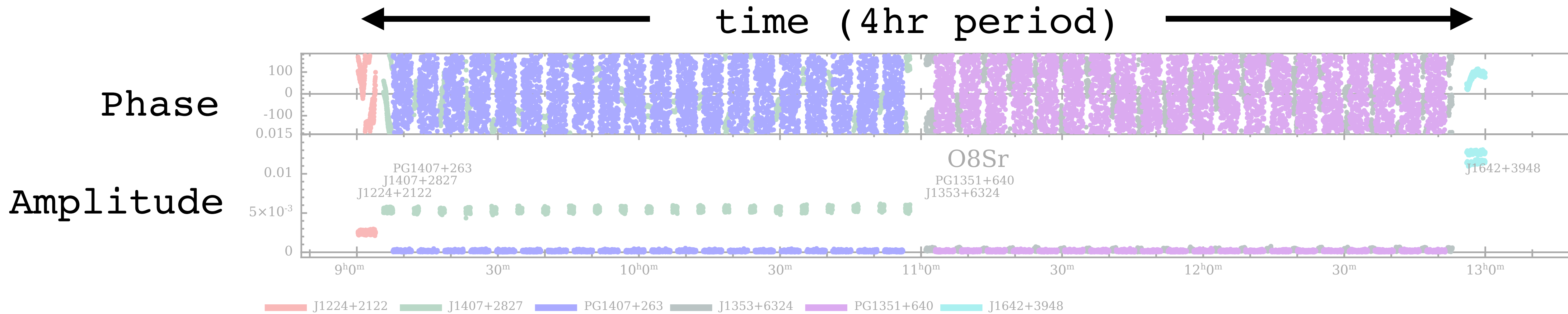
The EVN array fingerprint #3

Calibration scheme is **critical** for succes

- flux calibrator source?
- band-pass calibrator?
- phase-reference source cycle pattern?



Preferably close to the target source:
the big (radio)telescopes move slowly ...





All is lost then?



Deadline – 9th April-2024 at 1400UTC/1500BST/1600CEST

Proposal submission tool: <https://mfp.jive.eu>

As part of the European Commission’s Horizon 2020 OPTICON-RadioNet PILOT (GRP, Grant Agreement Number 101004719) we invite proposals for observing programs requiring multiple astronomical observing facilities to deliver unique science programs.

The [GRP](#) brings together expertise and access to a wide range of world leading ground based astronomical facilities covering a multitude of capabilities and wavelengths.

This call is an opportunity to request observing time on **multiple radio and/or optical facilities** via a **single science proposal and single review process**. **Access is provided to over eighteen individual facilities**.

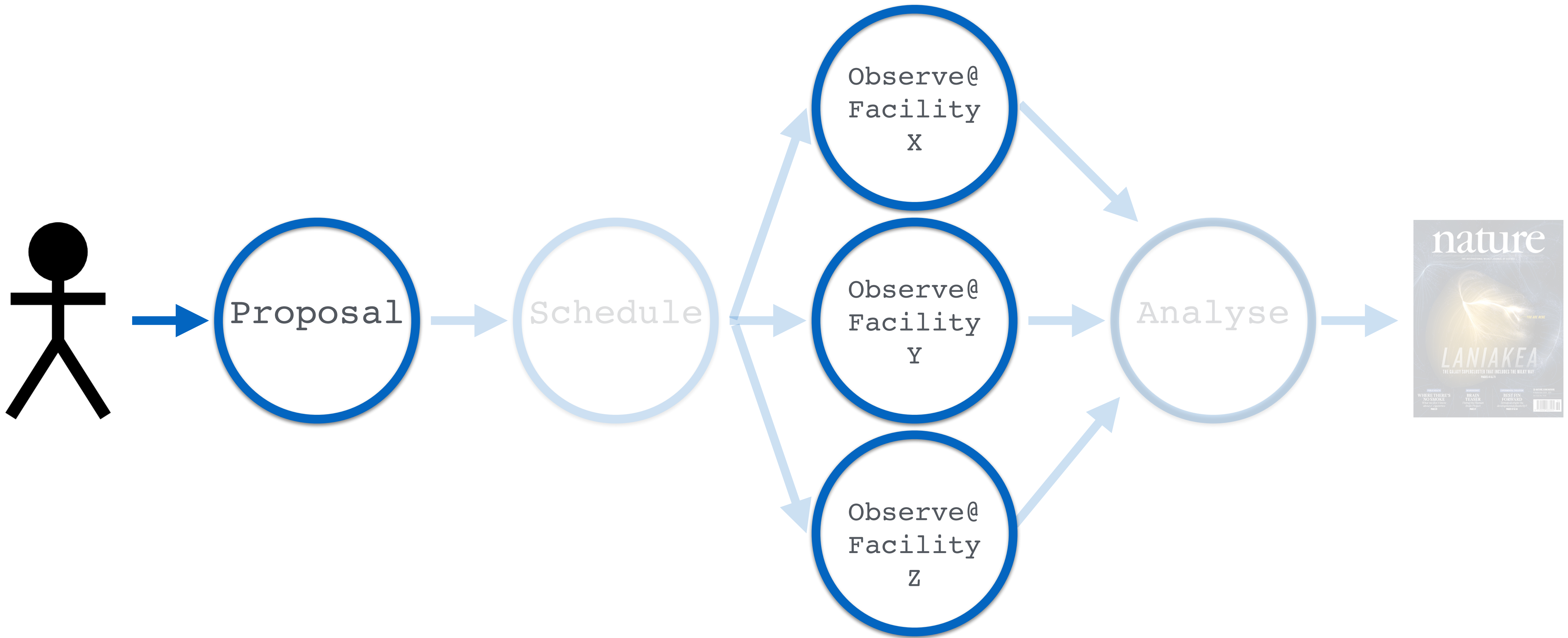
Technical and Science Opportunities

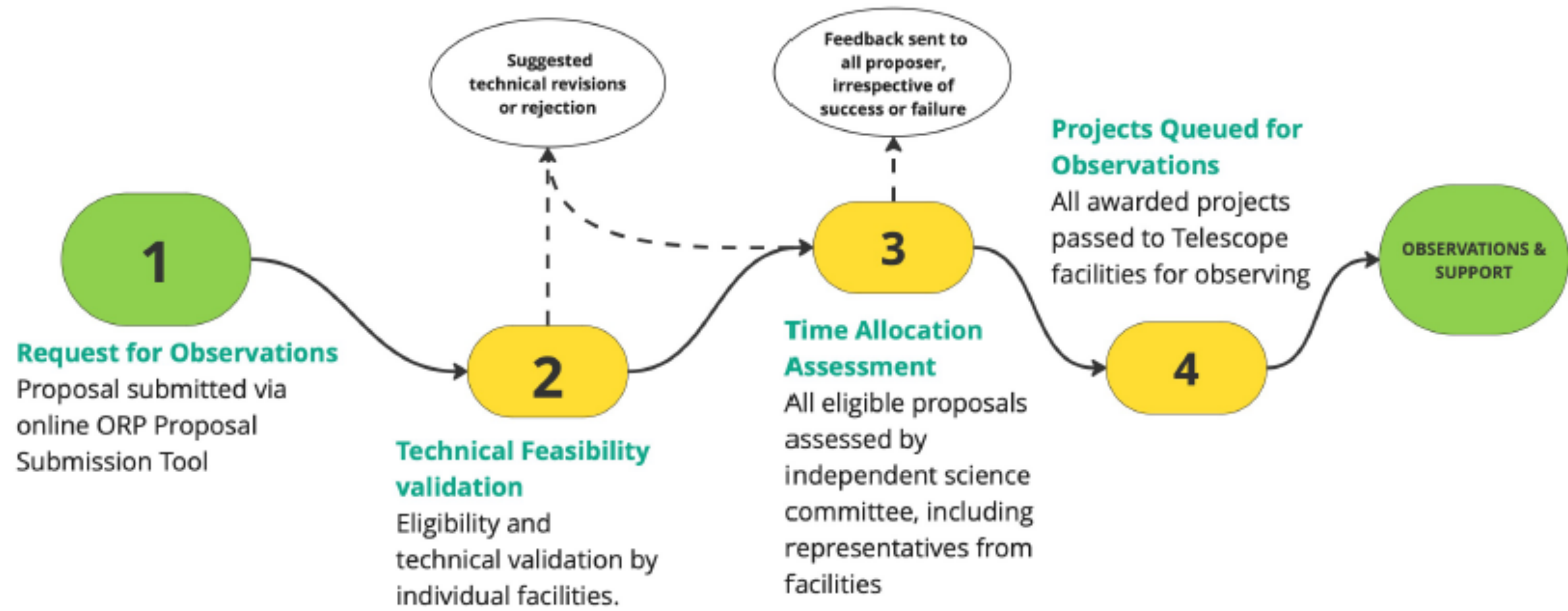
- **There are no constraints on the scientific areas that may be proposed for.** All requests must provide scientific and technical justification for the observing resources needed. All proposals are subject to eligibility requirements (see below).
- Observing programs must involve **two or more GRP** facilities. Observing proposals that request only one facility should be directed to regular calls for proposals of those facilities. Advice and feedback will be provided in this case.
- Requests for both non-simultaneous (nontime-critical) or time-critical/sensitive observations (e.g. contemporaneous or with requested cadence) will be considered. Proposers should carefully consider telescope scheduling capabilities and technical feasibilities (see links below). A technical review will be carried out prior to science assessment which may trigger follow-up requests for technical clarifications if required.

We advise prospective proposers to contact [Rob Beswick](#), [Marjolein Verkouter](#) or ta-call@orp-h2020.eu for advice, and any technical or science queries. Where necessary these will be passed on to experts at individual facilities.

Observing semester:

The observing semester is 1st August 2024 to 28th February 2025, but the semester dates for each telescope may vary from July 2024 though to March 2025 to accommodate their normal semesters and enable flexible scheduling (see table notes).





Facility	RI code	Allocation Unit
Anglo-Australian Telescope (Siding Spring)	AAT	15 nights
Canada France Hawaii Telescope	CFHT	4 nights
CAHA 2.2m (Calar Alto)	CAHA2.2	10 nights
CAHA 3.5m (Calar Alto)	CAHA3.5	5 nights
Carlos Sanchez Telescope (Observatorio del Teide, Tenerife)	TCS	14 nights
Effelsberg 100-m Radio Telescope	EF	100 hours
e-MERLIN	e-MERLIN	100 hours
European VLBI Network	EVN	100 hours
Las Cumbres Observatory	LCO	400 hours

Facility	RI code	Allocation Unit
Liverpool Telescope (ORM, La Palma)	LT	50 hours
LOFAR	LOFAR	50 hours
Nordic Optical Telescope (ORM, La Palma)	NOT	10 nights
Observatoire de Haut Provence 1.93m	OHP	10 nights
REM (La Silla, Chile)	REM	300 hours
South-African Large Telescope	SALT	50 hours
Sardinia Radio Telescope	SRT	100 hours
Telescope Bernard Lyot (Pic du Midi)	TBL	80 hours 10 nights
TNG Galileo (ORM, La Palma)	TNG	10 nights

Code	Title	Facilities requested
M1	Exploring Venus: a first "subsurface-to-space" snapshot	LOFAR, SRT, Effelsberg, TNG, LT, LCO, REM
M2	<i>Determining the progenitor types of interacting supernovae</i>	<i>NOT, LT, CAHA3.5</i>
M3	<i>Developing our understanding of extreme coronal line emitters</i>	<i>NOT, LCO</i>
M4	<i>A search to discover the first RV variable white dwarf with a planet</i>	<i>SALT, CAHA3.5</i>
M5	<i>Precession of the orbital planes of hot Jupiters: studying new candidates</i>	<i>NOT, TBL, OHP, TNG, REM, CAHA3.5</i>
M6	<i>Gamma-ray burst afterglows with robotic facilities available within the ORP call</i>	<i>LT, LCO</i>
M7	Probing the immediate aftermath of core-collapse supernova explosions	e-MERLIN, LT
M8	Simultaneous Multi-telescope, Multi-band SETI (SMM-SETI)	LOFAR, Effelsberg, SRT, e-MERLIN, EVN
M10	Investigating the second bump in the light curve of the closest recurrent nova	LOFAR, e-MERLIN, EVN, NOT
M11	Exploring a new population of Fast X-Ray Transients (FXRTs)	e-MERLIN, SALT, LCO, CAHA2.2
M12	Star formation in dwarf galaxy NGC 4163 with peculiar emission line ratios	e-MERLIN, NOT,
M13	<i>Luminous flares from the supermassive black holes in galactic nuclei</i>	<i>NOT, LT</i>
M14	An e-MERLIN-LCO Survey of VLBI-Gaia Offset Sources	e-MERLIN, LCO

Code	Title	Facilities requested
M1	Exploring Venus: a first "subsurface-to-space" snapshot	LOFAR, SRT, Effelsberg, TNG, LT, LCO, REM
M3	<i>Developing our understanding of extreme coronal line emitters</i>	<i>NOT, LCO</i>
M5	<i>Precession of the orbital planes of hot Jupiters: studying new candidates</i>	<i>NOT, TBL, OHP, TNG, REM, CAHA3.5</i>
M7	Probing the immediate aftermath of core-collapse supernova explosions	e-MERLIN, LT
M8	Simultaneous Multi-telescope, Multi-band SETI (SMM-SETI)	LOFAR, Effelsberg, SRT, e-MERLIN, EVN
M10	Investigating the second bump in the light curve of the closest recurrent nova	LOFAR, e-MERLIN, EVN, NOT
M11	Exploring a new population of Fast X-Ray Transients (FXRTs)	e-MERLIN, SALT, LCO, CAHA2.2
M12	Star formation in dwarf galaxy NGC 4163 with peculiar emission line ratios	e-MERLIN, NOT,
M14	An e-MERLIN-LCO Survey of VLBI-Gaia Offset Sources	e-MERLIN, LCO

Analysis: EVN, e-Merlin

Analysis: EVN, e-Merlin

Feasible to
automate or
benefit from TOM?

Basically, MF-CfP proposals seem to be of the form:

“Point network at target source until
we have a sensitivity of $X \text{ W/m}^2/\text{Hz}$ ”

The reasoning being:

“The calculator for the facility says
that sensitivity X is reached in Y hours
of observing, so we request Y hours.”

Expected target source flux density could
be a required parameter in the proposal?

Few of the proposals are specific about the calibration strategy:

“The mentioned observation strategy is a simple 10 minute cycle with a calibrator and the source”

Moderately specific - scheduling parameters might be taken from proposal

“Including standard calibration overheads, we expect this will result in a total of 25 hours of EVN time”

Seems to imply a “standard observation strategy”?

“30 min each on a flux and pass-band/gain calibrator source, then repeating a cycle of 5 minutes on the target source and 1 minute on a phase calibrator”

Quite specific - scheduling parameters might be taken from proposal

Requested response time?

“2 out 4 e-Merlin proposals require a quick reaction time of the network”

response times requested:

- on the order of a day
- within a week

Requested response time?

“2 out 4 e-Merlin proposals require a quick reaction time of the network”

response times requested:

- on the order of a day
- within a week

A TOM-based approach with some automation might be beneficial?

Future?



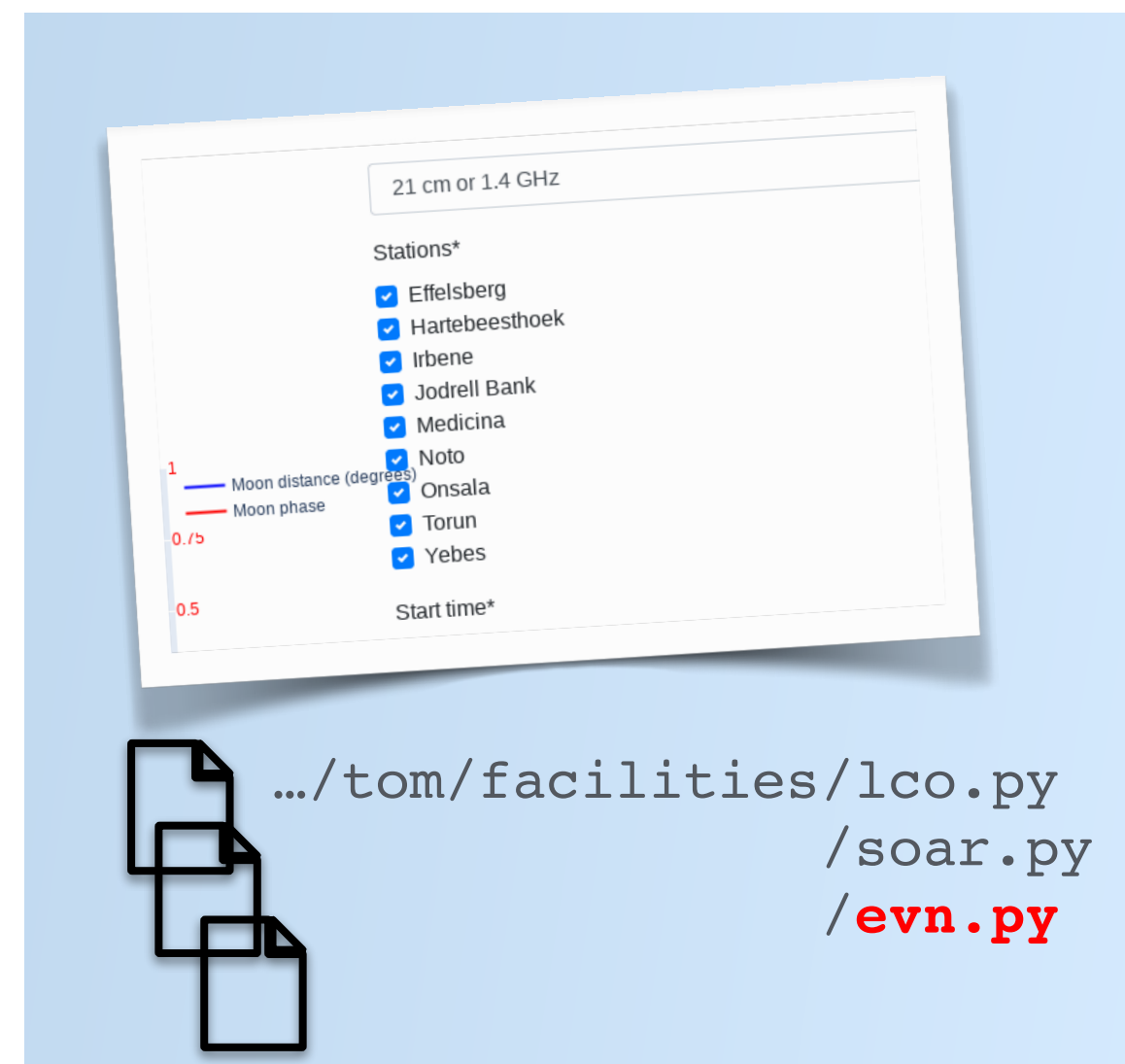
webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>

Future?

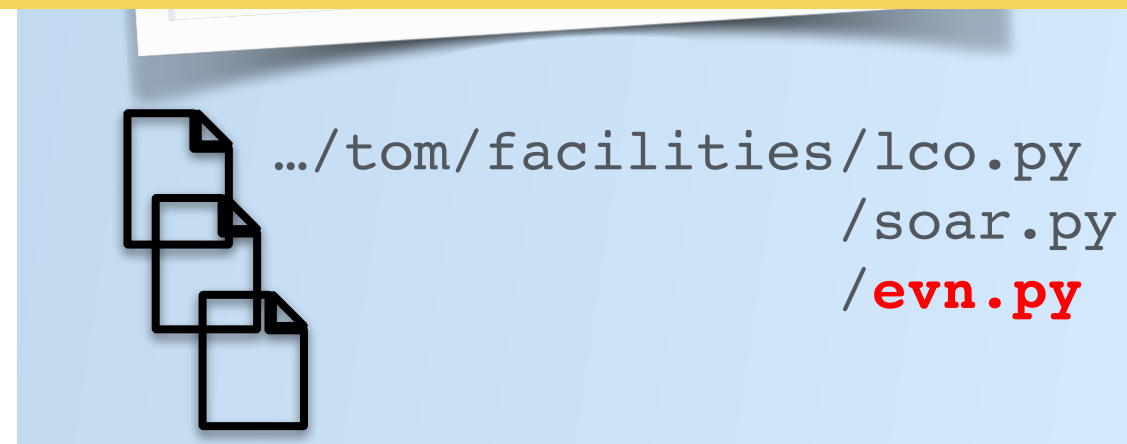


webserver running TOM e.g.
<https://bh-tom.astrolabs.pl/>

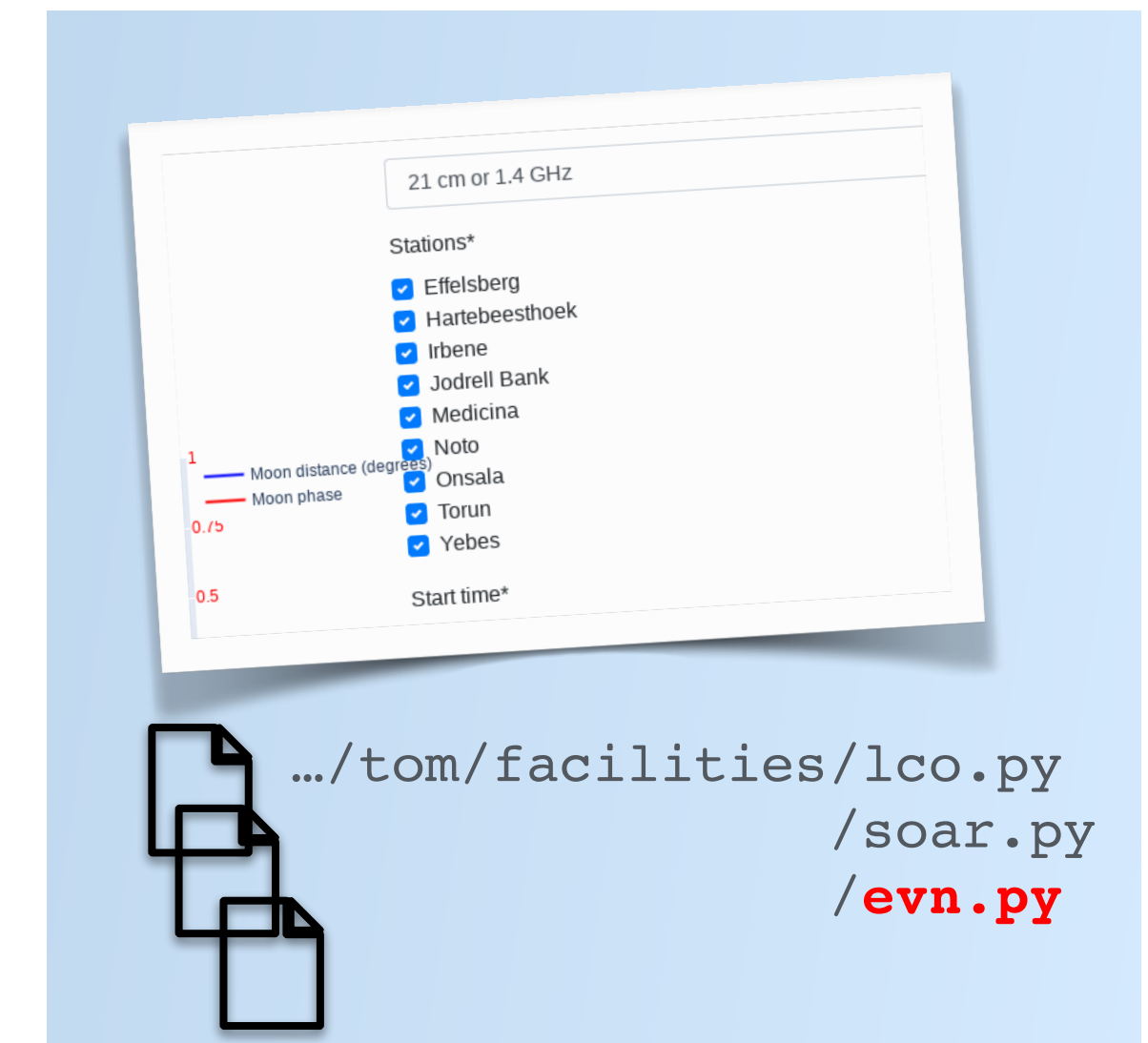
Future?



webserver running TOM



webserver running TOM



webserver running TOM

EVN observations db
- status
- data products

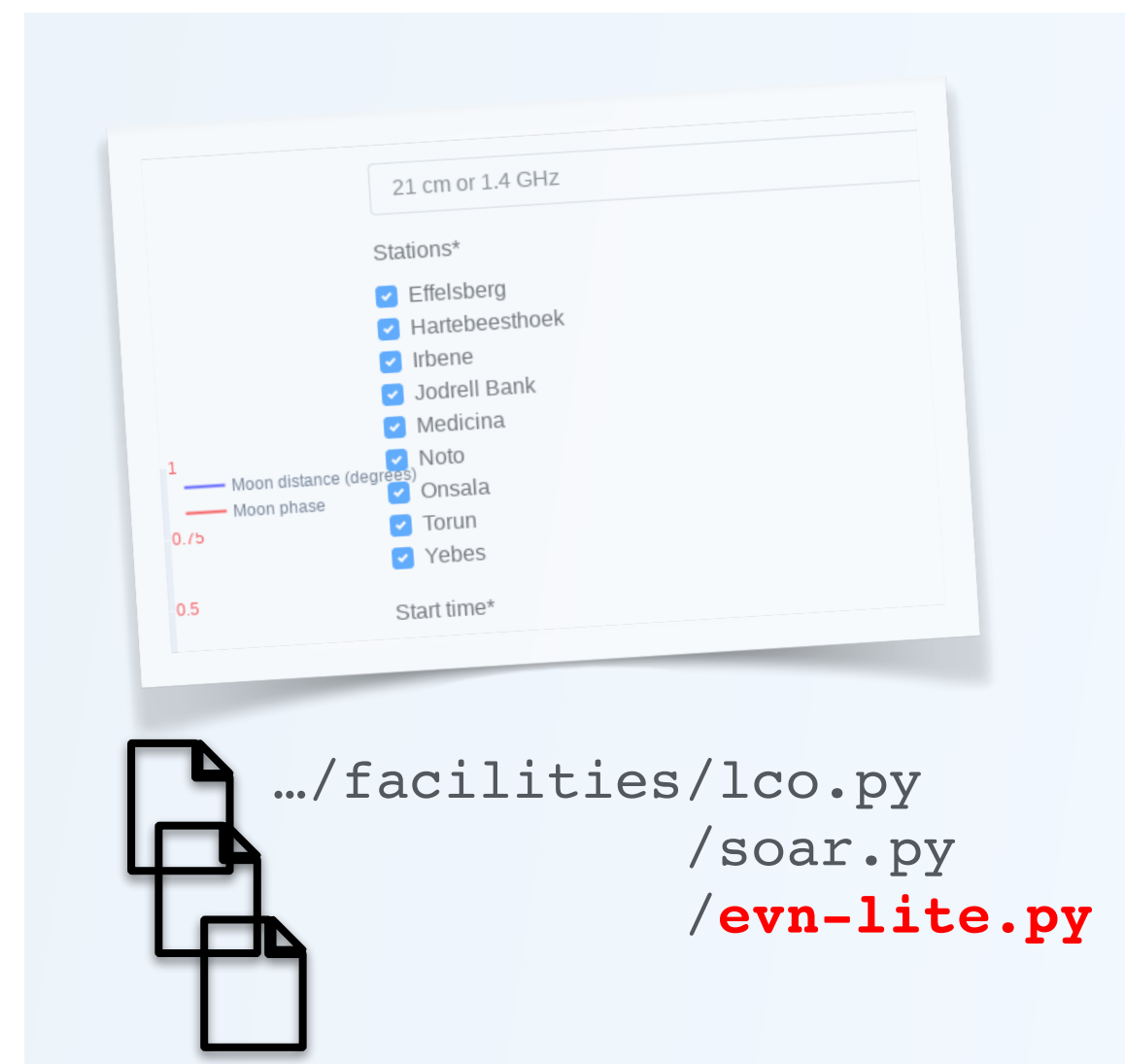
Planobs code base

pySCHED scheduling

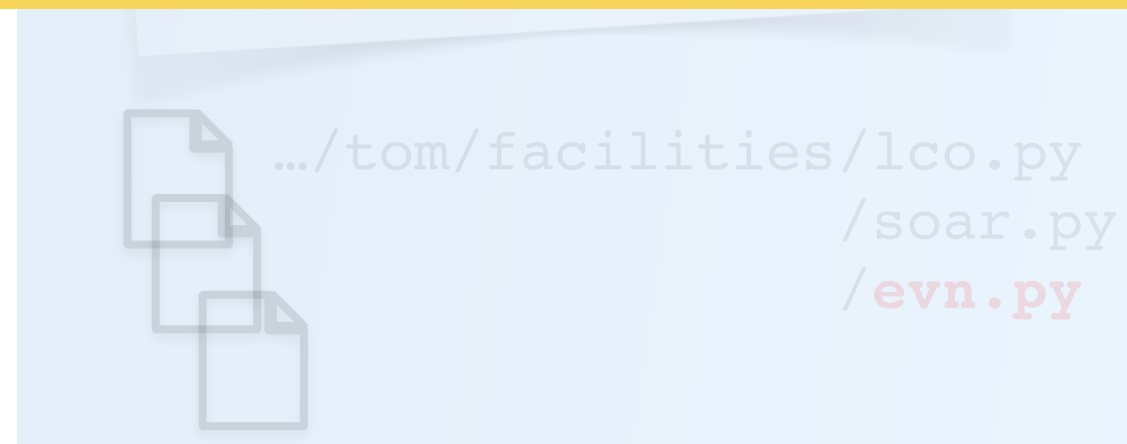
availability tool

server running EVN backend
<https://tom-backend.jive.eu/>

Future? EVN-lite subarray?



webserver running TOM



webserver running TOM

EVN observations db
- status
- data products

Planobs code base

pySCHED scheduling

availability tool

server running EVN backend
<https://tom-backend.jive.eu/>

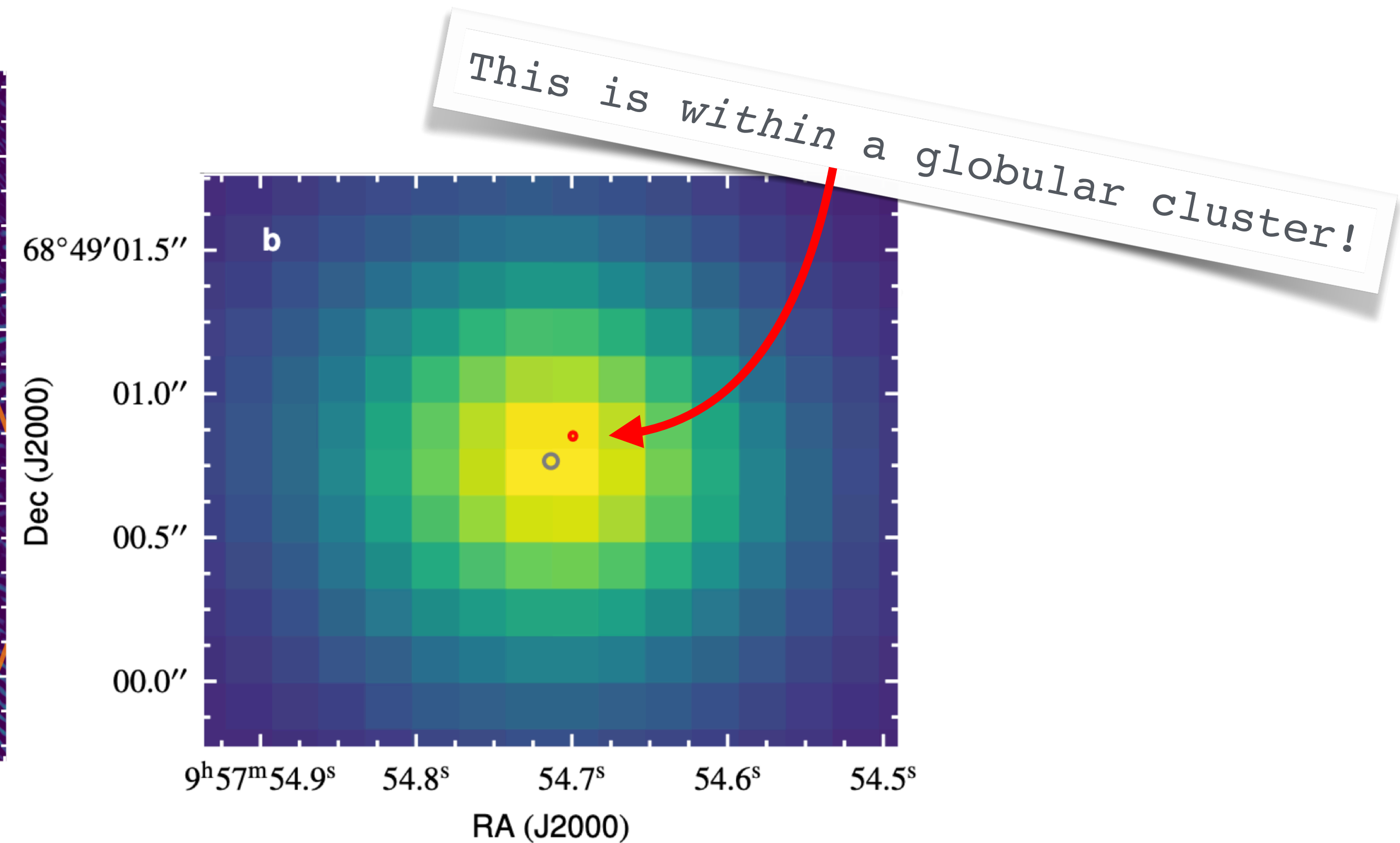
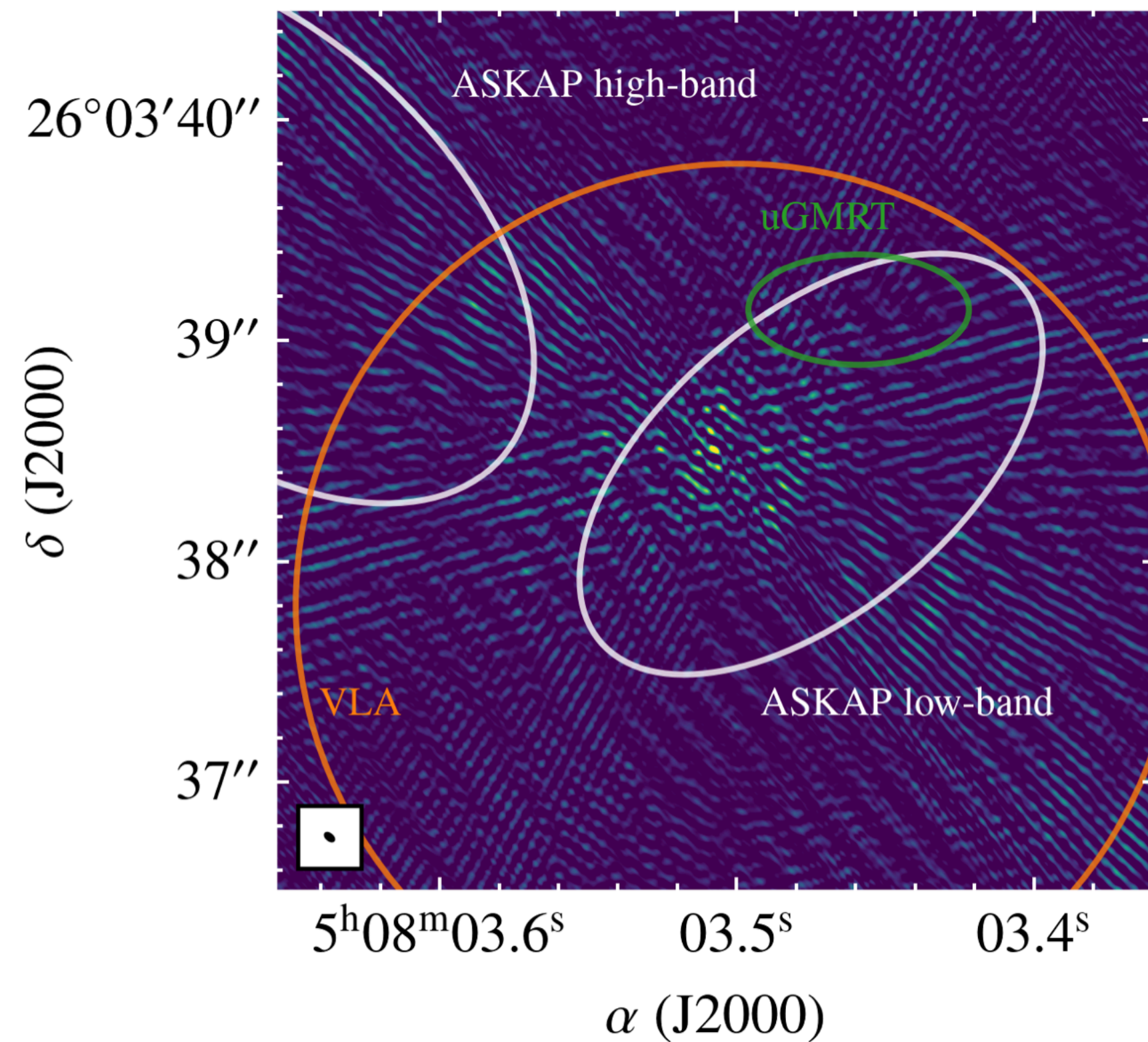


webserver running TOM

EVN-lite: subset of small EVN telescopes
offerening more observing hours to
large programs

PRECISE

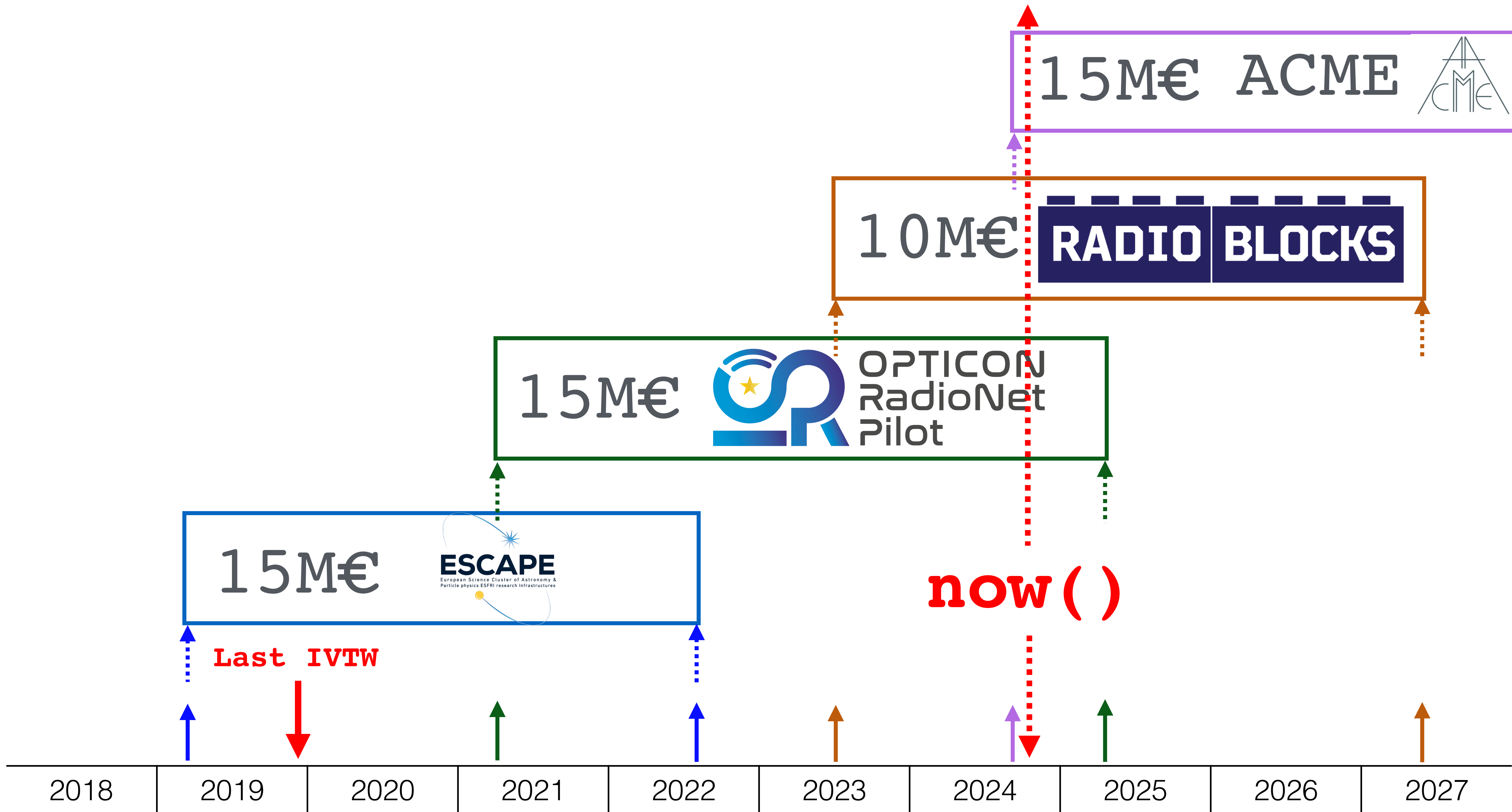
"**P**inpointing **RE**peating **Ch**ime **S**ources with **EVN** dishes"



Two different Fast Radio Burst localisations to mas precision

"PRECISE localizations of repeating Fast Radio Bursts" Marcote et al. <https://arxiv.org/abs/2202.11644>

?



Thanks for
attention!