

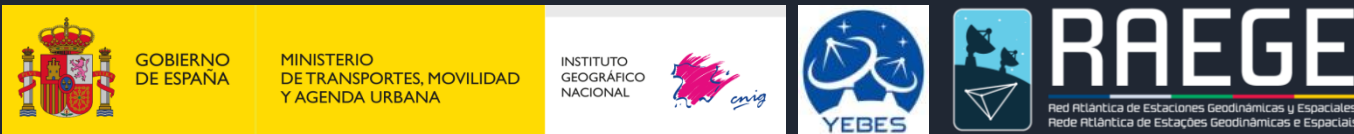
INSTRUMENTATION DEVELOPMENTS FOR VGOS AT YEBES OBSERVATORY

May 7th, 2025

13th TOW Meeting

MIT - Haystack Observatory (USA)

José A. López-Pérez
on behalf of Yebes Observatory staff



Miguel Gómez Garrido
astro.miguel photo

Introduction to Yebes Observatory, IGN-Spain



50th anniversary

Expertise

- 35+ years of expertise in receivers and LNAs developments.
- VLBI & SLR operations

National Role

- Key research facility of Spain's scientific infrastructure.
- **1st Spanish GGOS core site**

International Recognition

- Designated as IVS Technological Development Center since 2015.
- Collaborating with international partners in astronomy and geodesy.

Broadband Receiver Advancements

1

First Generation (S/X)

Legacy S/X receivers with limited bandwidth and slow antennas.

2

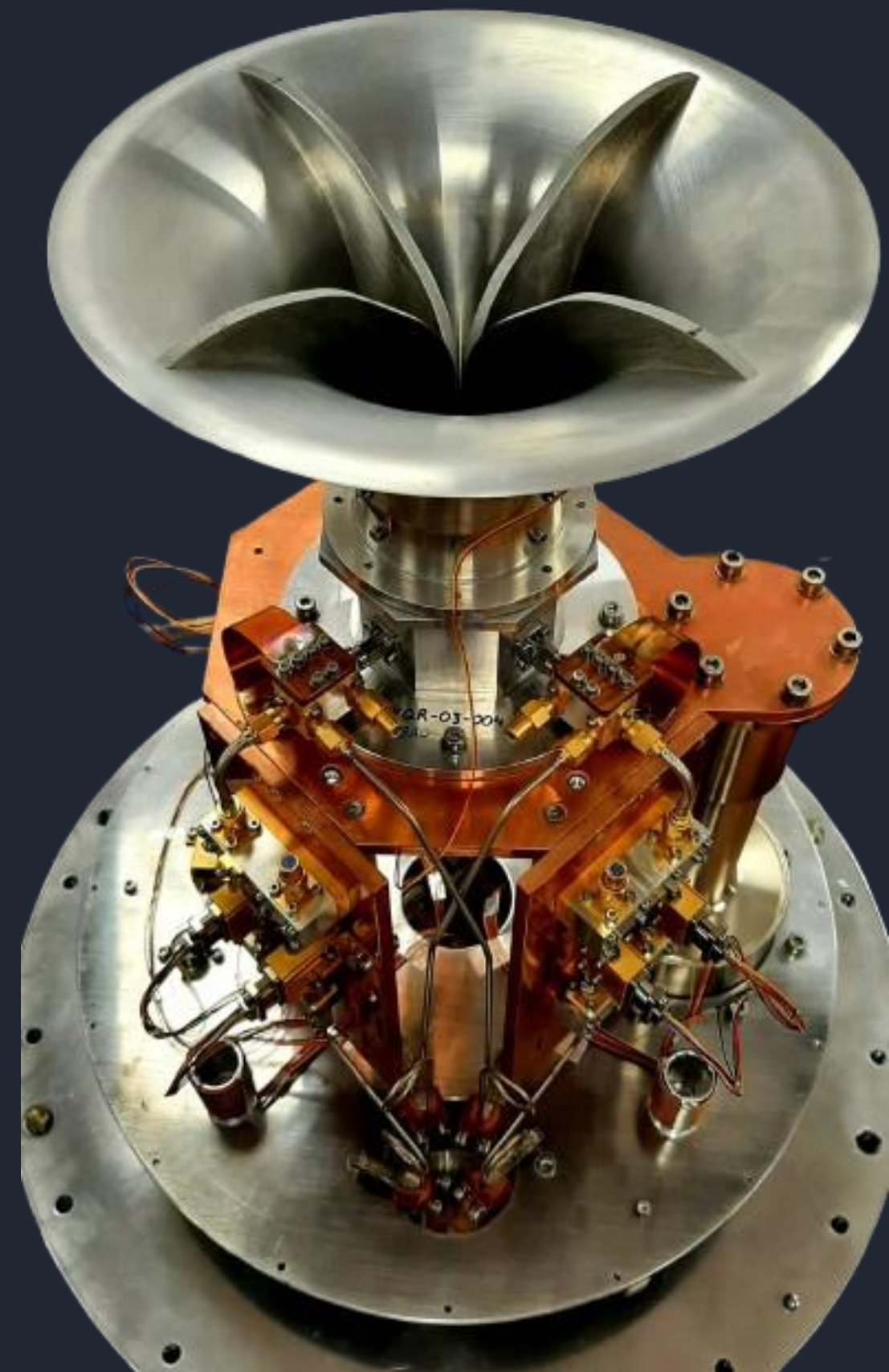
Current VGOS (2 – 14 GHz)

Wideband receivers with larger bandwidth, sampling and small high-speed antennas

3

Yebe's Innovations for VGOS

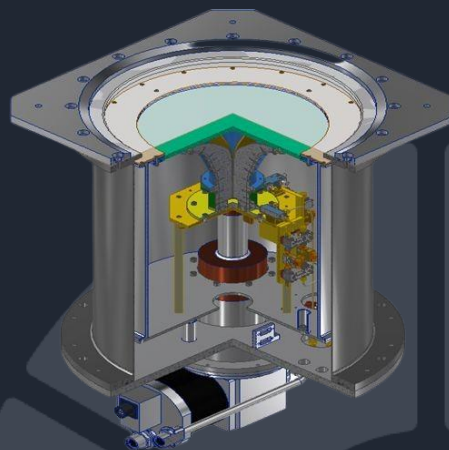
- Improved QRFH & dewar
- Cryo 30dB couplers
- Balanced LNAs
- FO CDMS
- PhaseCal equalizer
- HTS filters for RFI



Design Innovations

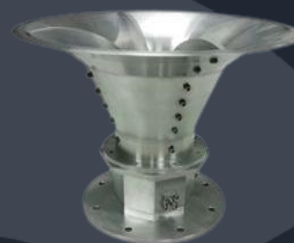
Cryogenic System

Compact dewar achieving 10K for critical components.



Feed Horn

Quad-ridge design enabling octave bandwidth with good matching and phase center stability.



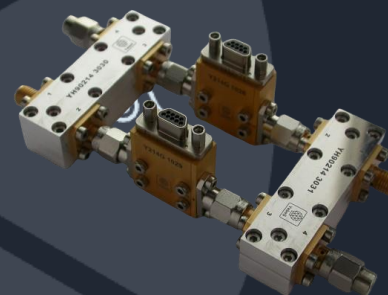
CDMS

Optimized CDMS using 5MHz fiber optic transceivers



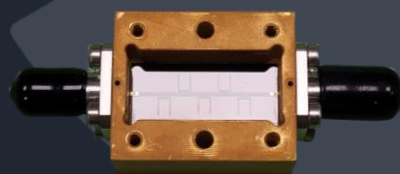
Low-Noise Amplifiers

Custom-designed InP HEMT LNAs with noise temperatures below 4K.



HTS Filters

Advanced HTS filter implementation to eliminate RFI signals.



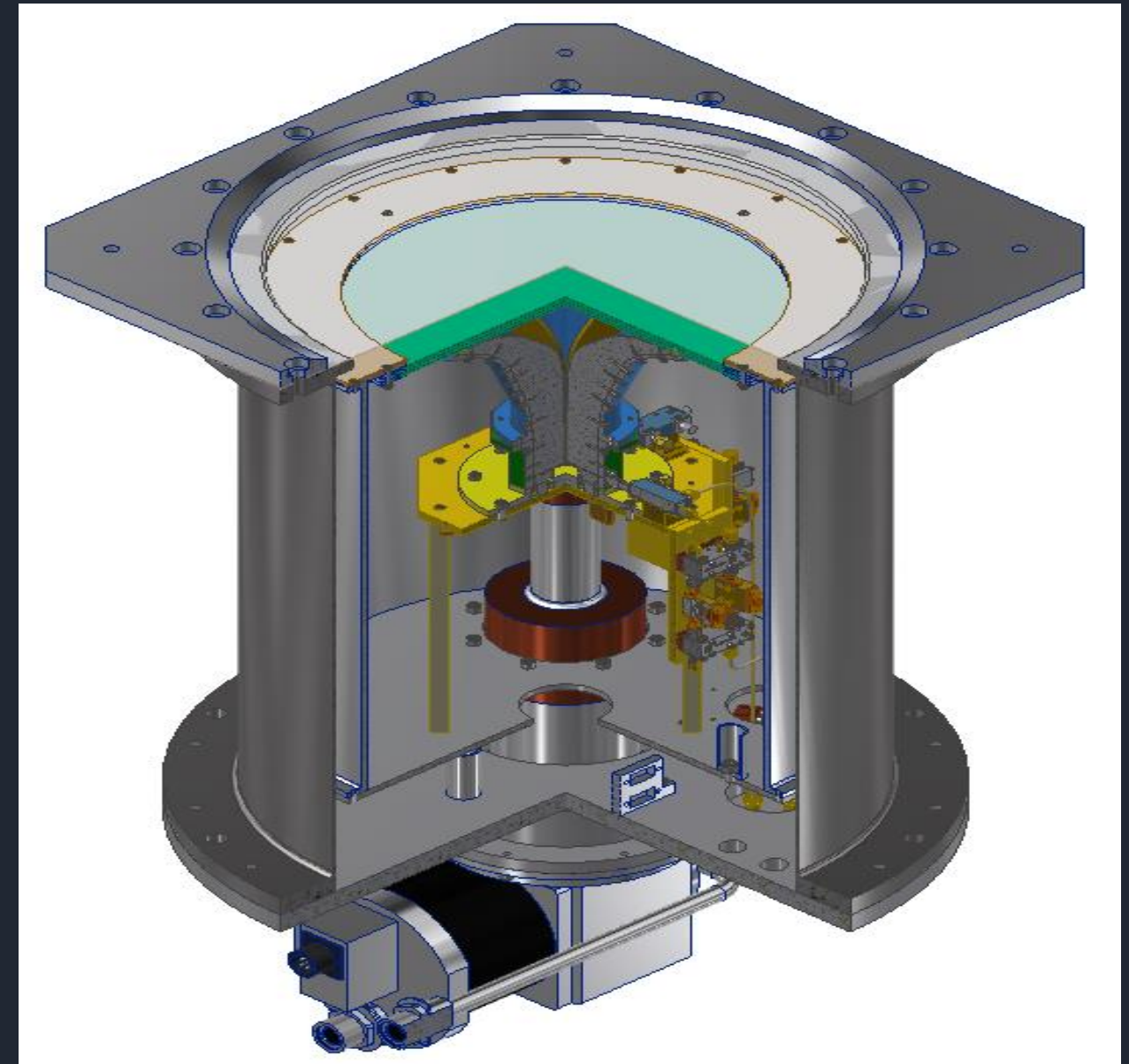
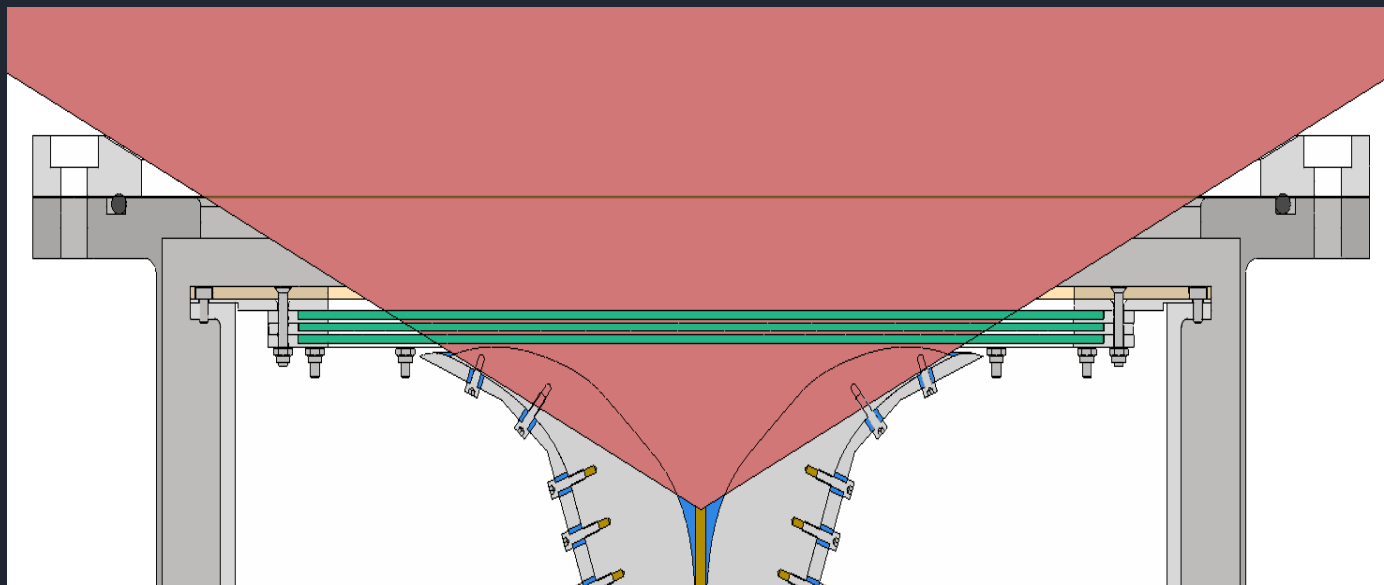
Calibration System

Equalized NoiseCal and PhaseCal injection via customized cryo coupler.



Optimized dewar

	1st gen.	2nd gen.
Vacuum window diameter	319 mm	335 mm
Feed distance to window	45 mm	37 mm
Clearance angle for feed	58°	65°
IR filter	Polystyrene	Polystyrene
Cryostat volume	59 L	49 L

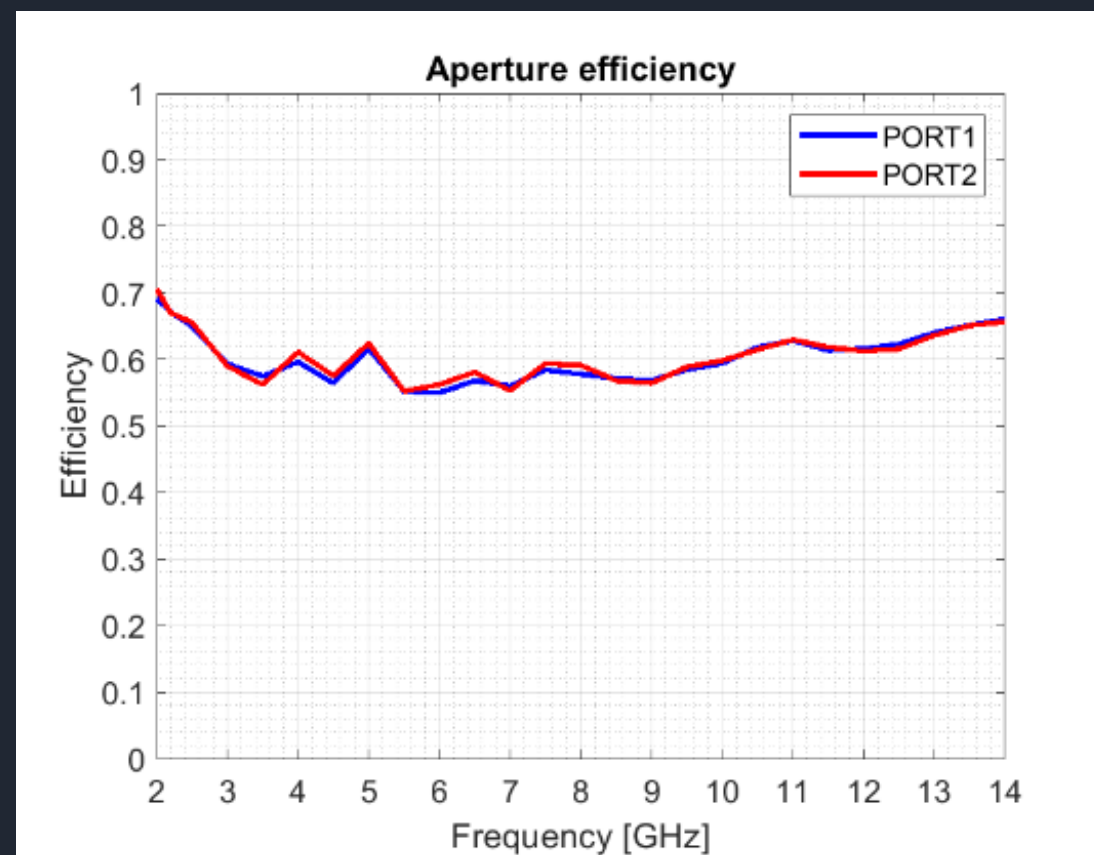
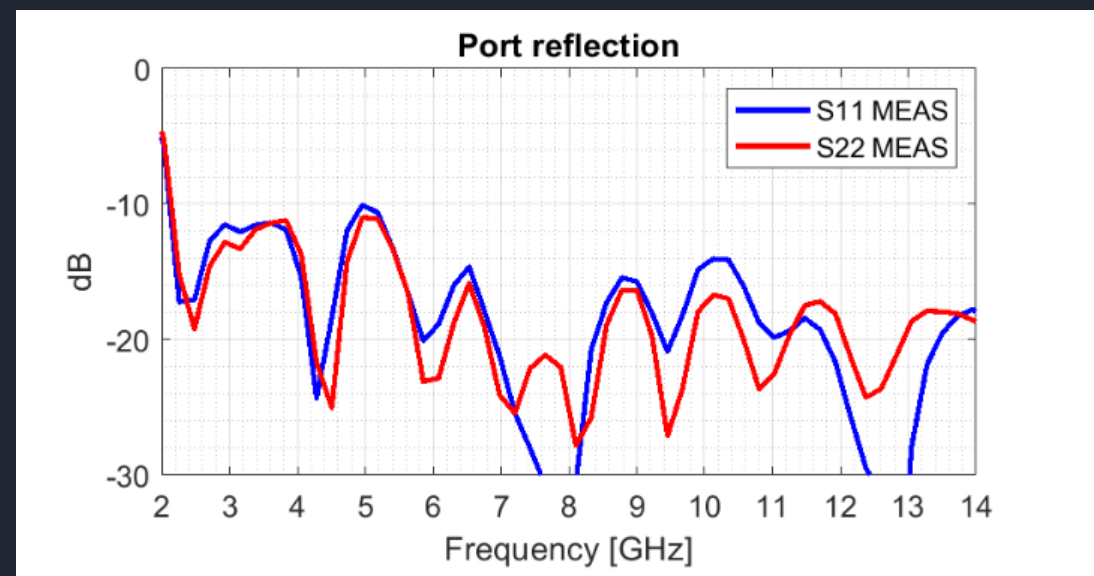


Improved QRFH

- YQR-03-00X Feed Horn
- Frequency range: 2.2-14 GHz
- Improved port reflection (return losses): Below -10dB across the whole bandwidth.
- Aperture efficiency in the 13.2m radio telescope: 60%



(IT-CDT-2020-26 <https://icts-yebes.oan.es/reports/doc/IT-CDT-2020-26.pdf>)



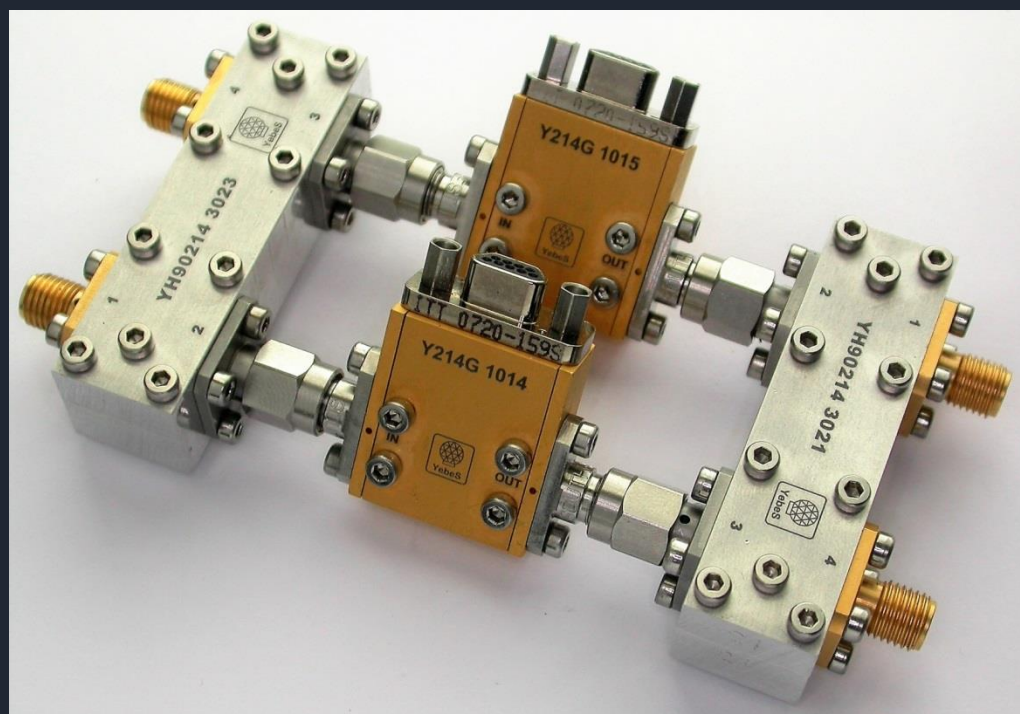
Cryogenic 30dB coupler

- Optimized in the range 3-14 GHz, usable down to 2 GHz.
- Specially designed to withstand thermal cycling and operate at cryo temps
- Port matching ≤ -20 dB
- Coupling = -29.2 ± 1 dB
- Insertion loss < 0.3 dB @ 14GHz & 15K
- Connectors contribution $\approx 2 \times 0.1$ dB
- Size: 21.3 x 14.5 x 17 mm
- Weight: 22 grams.
- Units available at 1k€

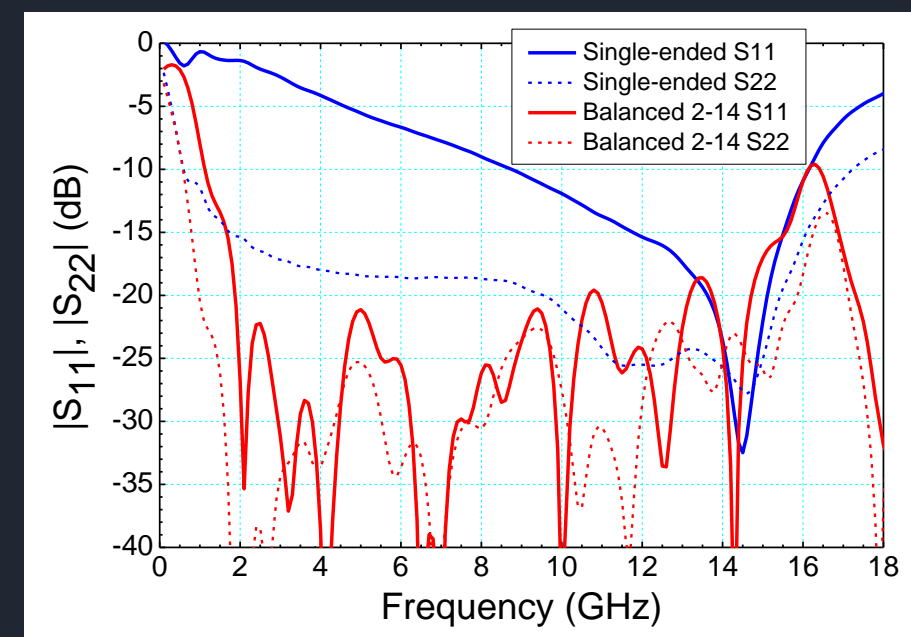
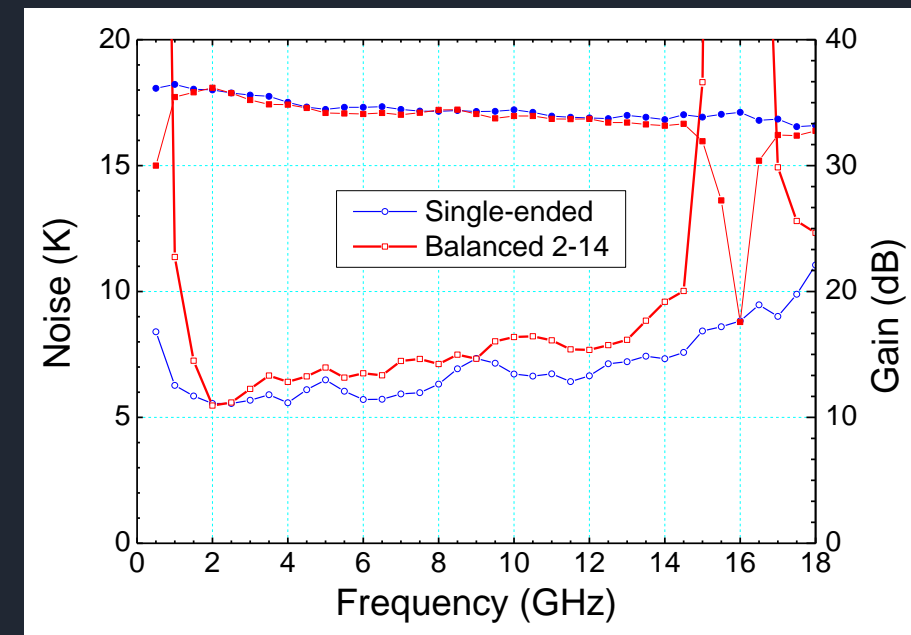


Balanced LNAs

- Using 3dB/90° hybrids
- Very low noise penalty
- Drastic improvement of input reflection
- Balanced versions for 2-14 GHz and 1.5-15.5 GHz available



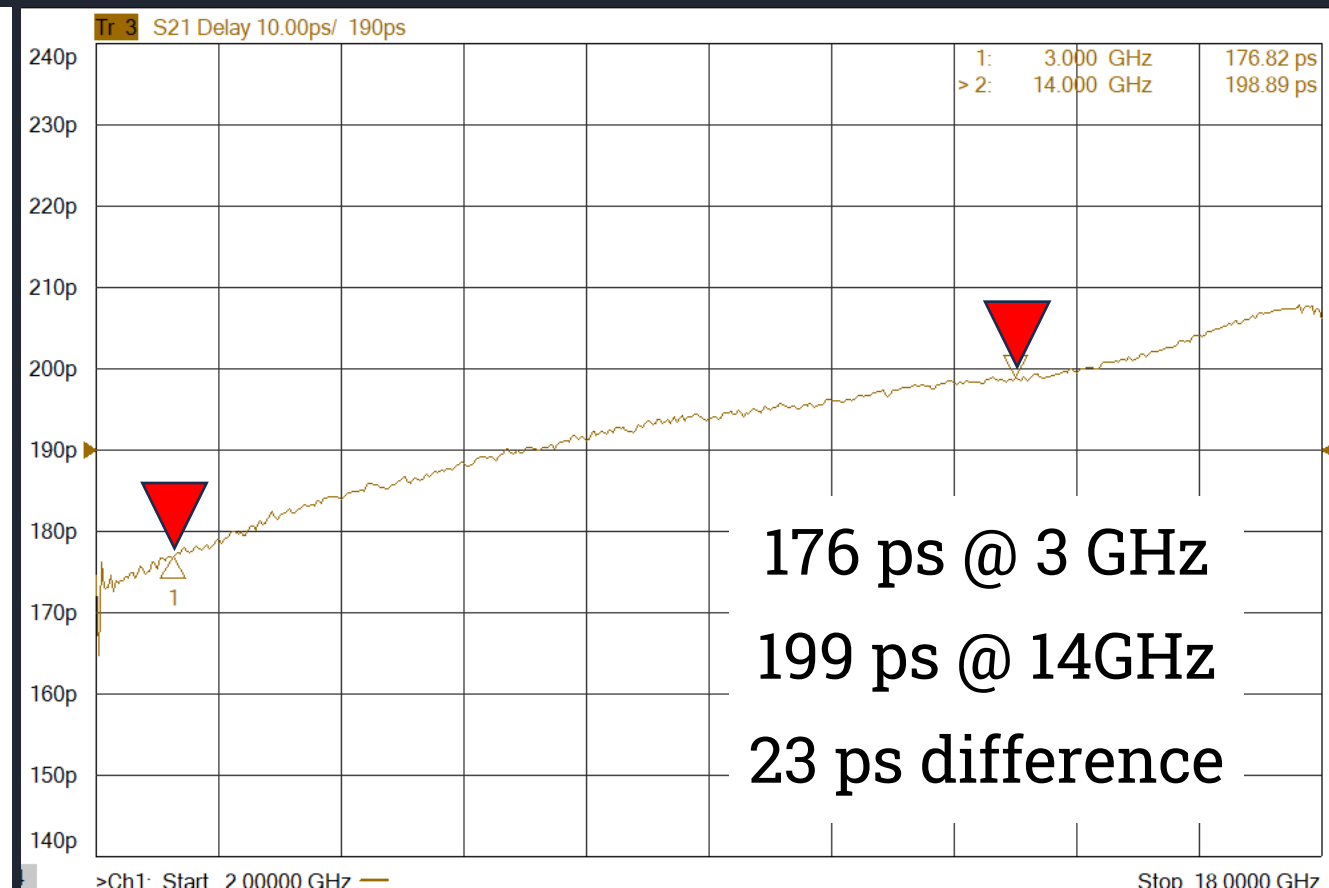
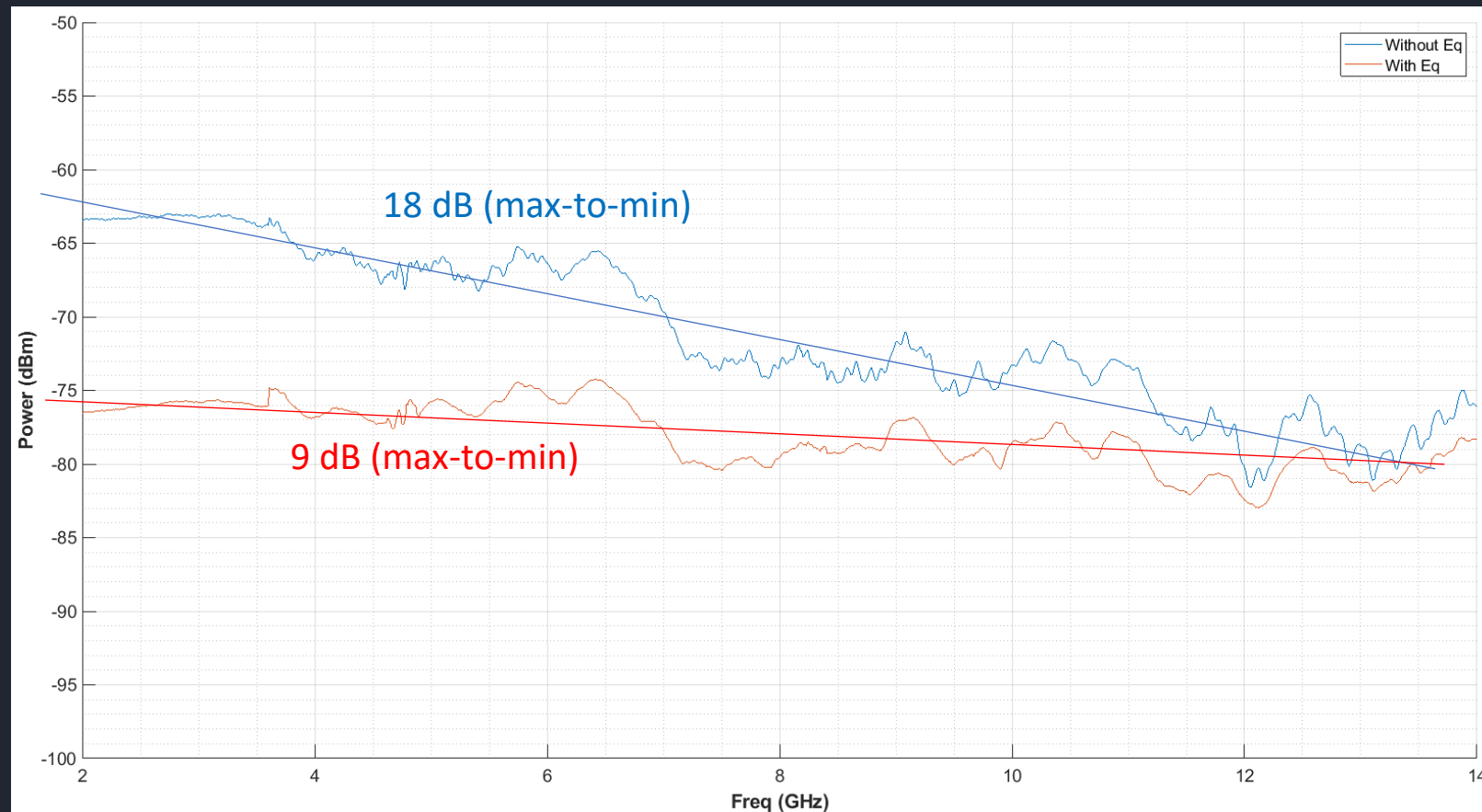
Band	2-14 GHz
Tn	7.6 K
Gain	33.8 dB
IRL	-21 dB
ORL	-23 dB



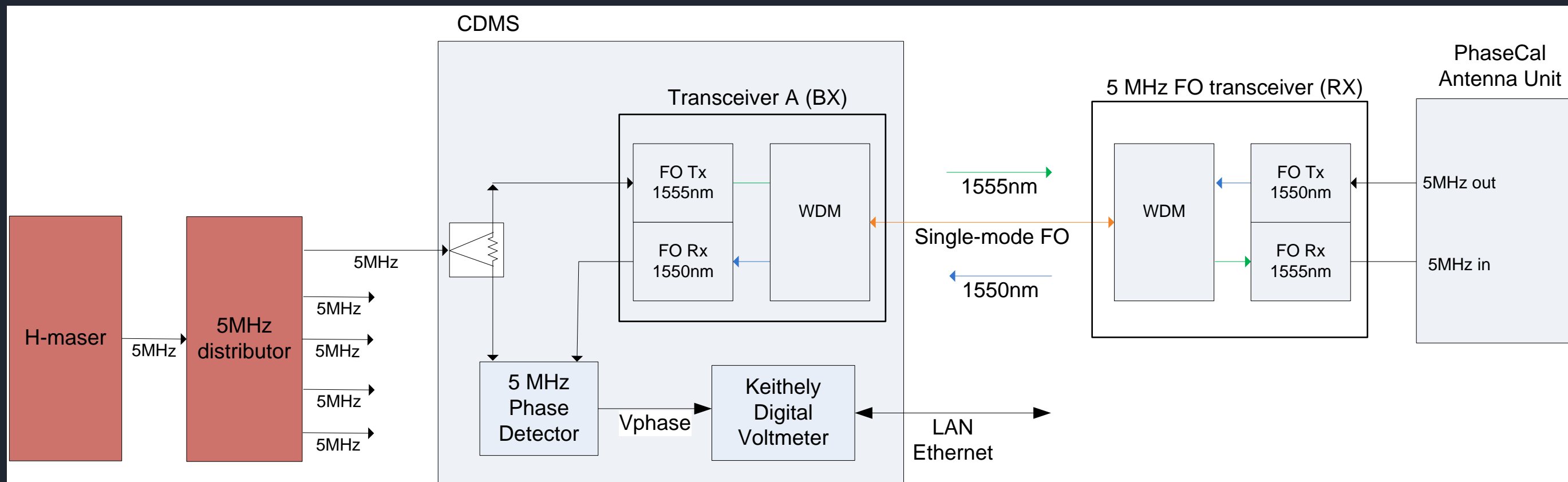
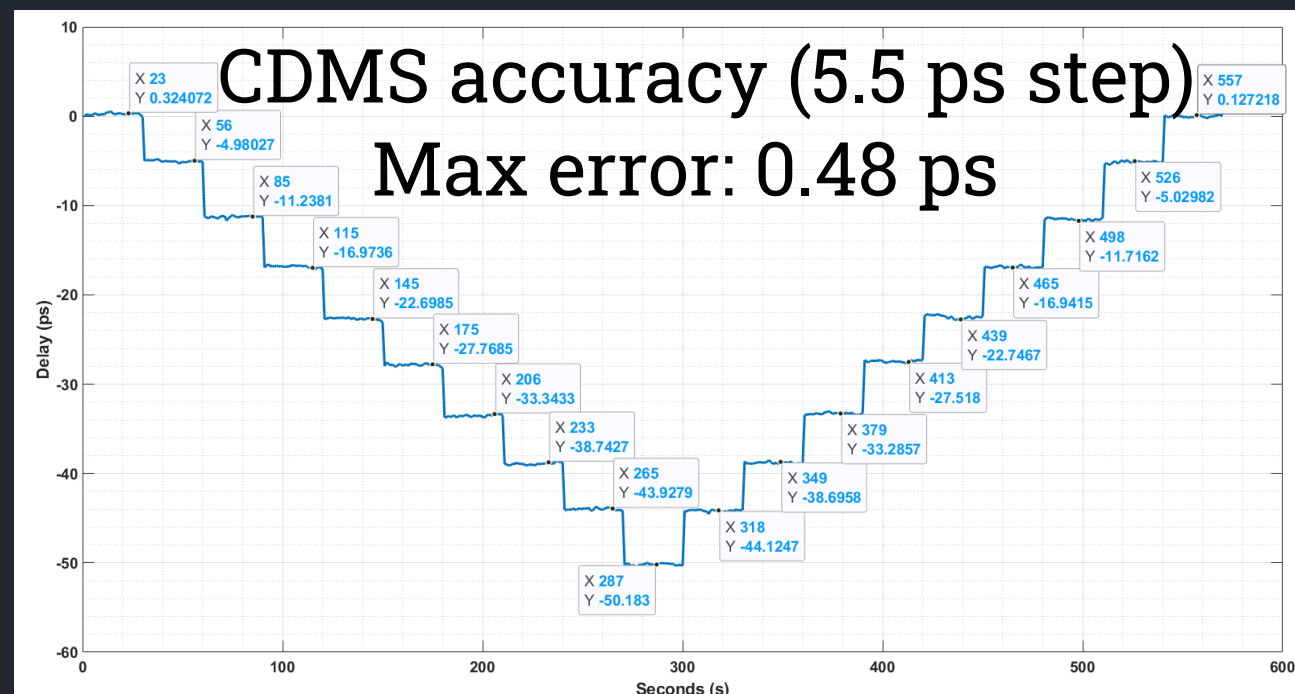
I. Malo-Gomez, J. D. Gallego-Puyol, C. Diez-Gonzalez, I. Lopez-Fernandez and C. Briso-Rodriguez, "Cryogenic Hybrid Coupler for Ultra-Low-Noise Radio Astronomy Balanced Amplifiers," in *IEEE TMTT*, Dec. 2009

Equalized PhaseCal

Equalized spectrum of tones

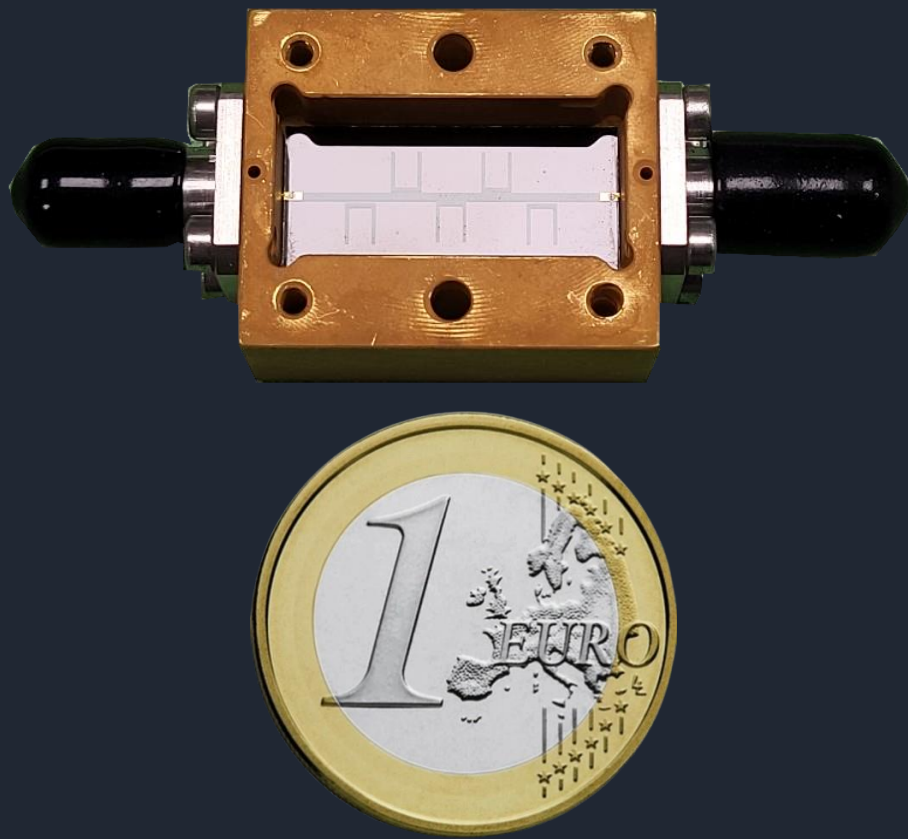


FO CDMS

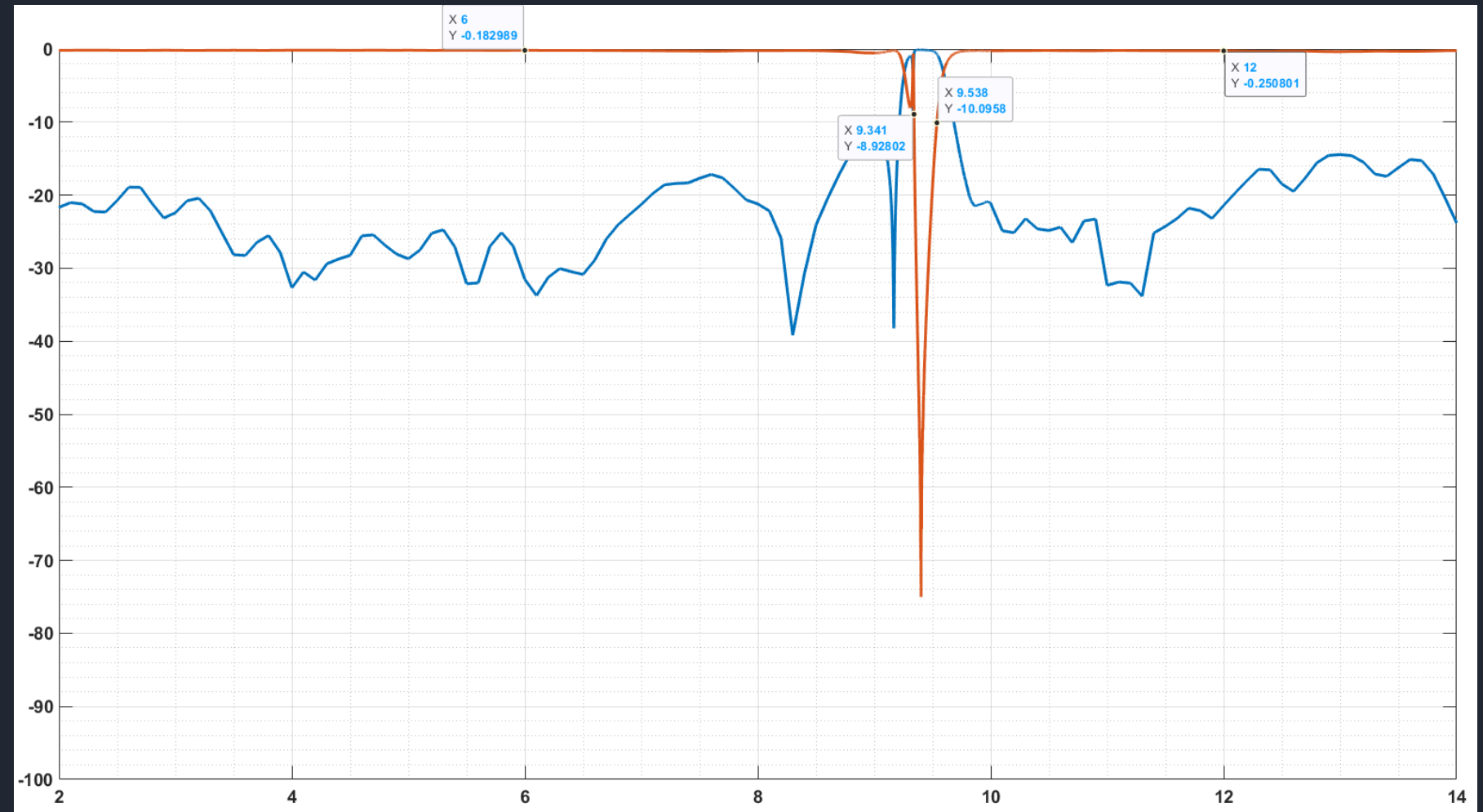


HTS filters

HTS notch filter
for SLR radar at 9.4 GHz



Measured performance @ 10K



See Pablo García-Carreño "Development of HTS filters at Yebes Observatory", 9th IVTW-2024, Haystack Observatory

HTS filters

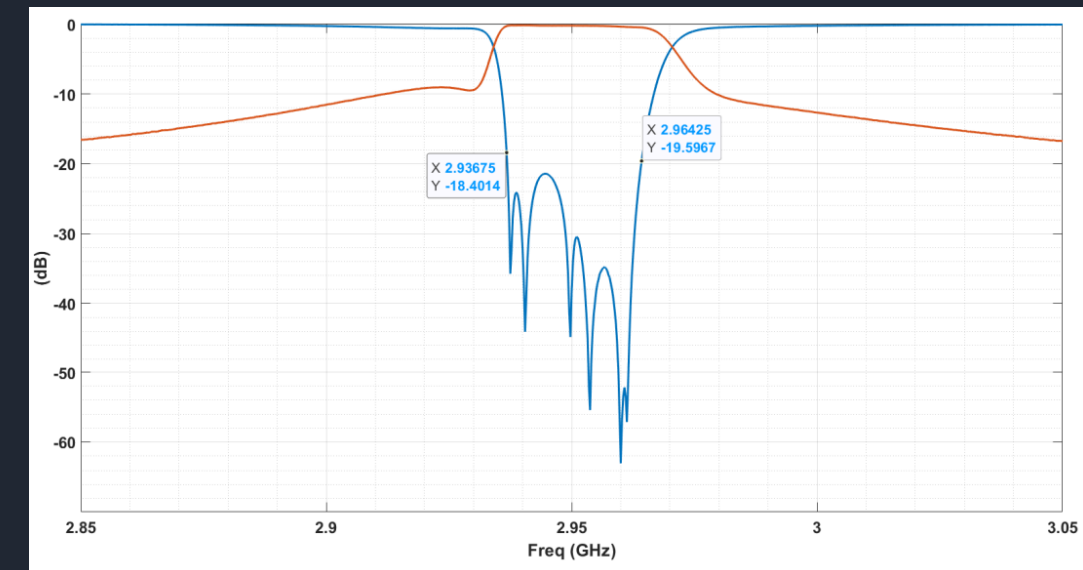
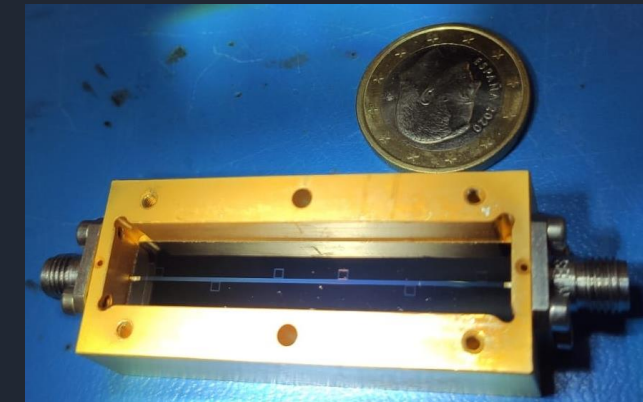
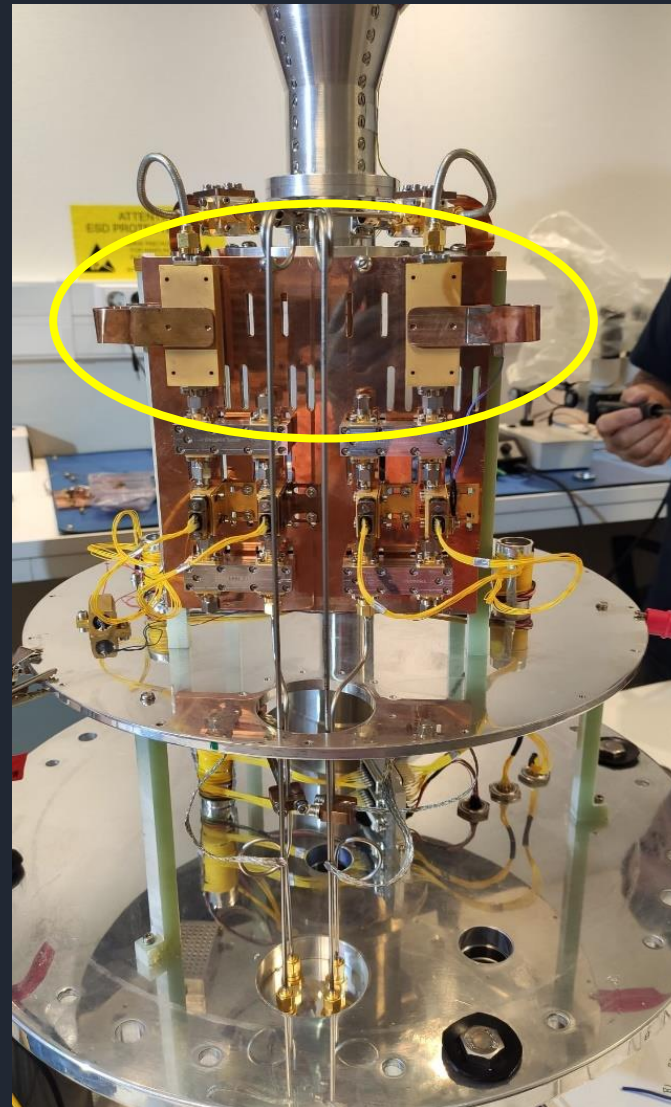
HTS notch filters for SD radar at 2.95 GHz

Upgraded VGOS receiver for RAEGE

Santa Maria:

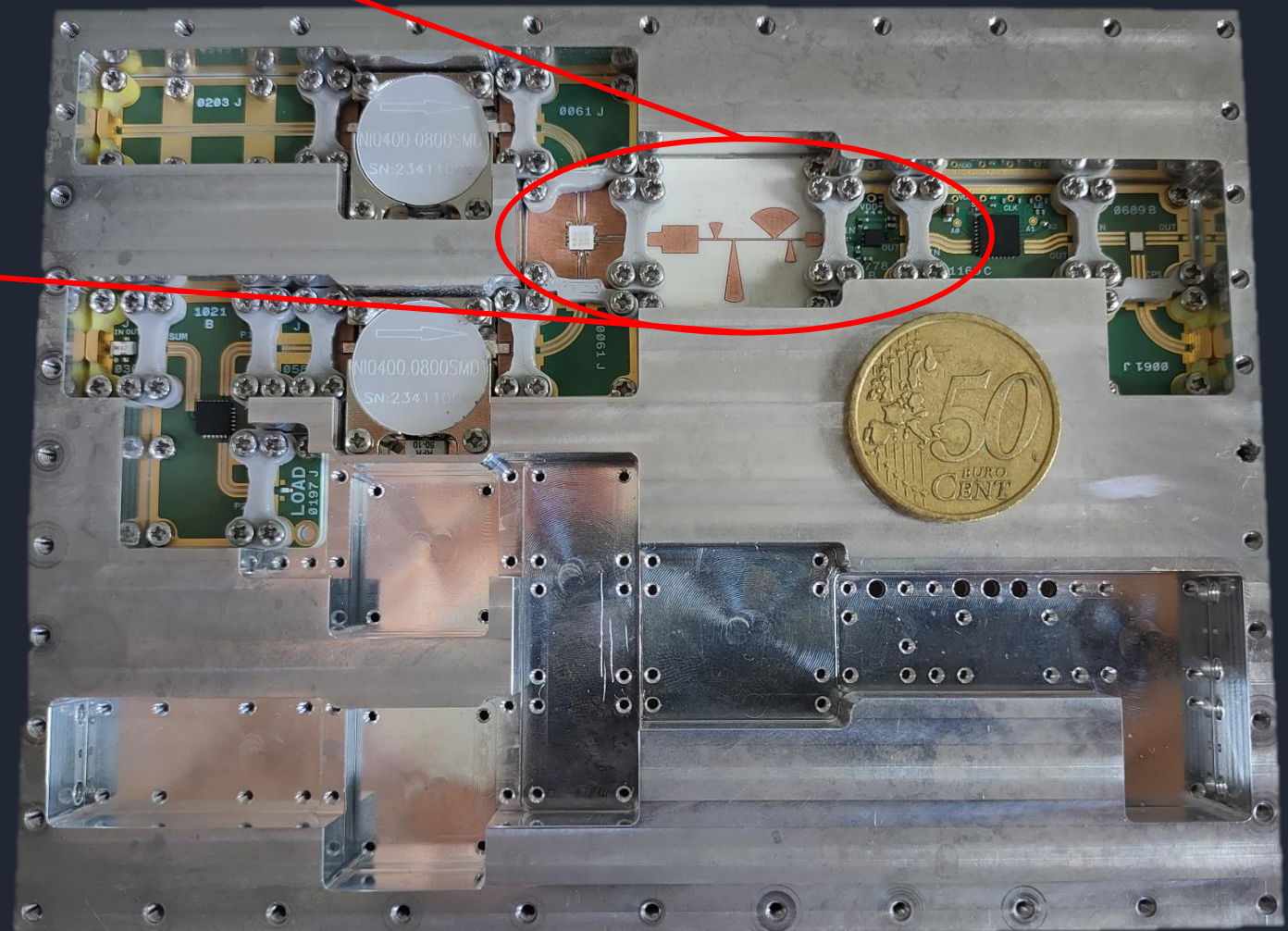
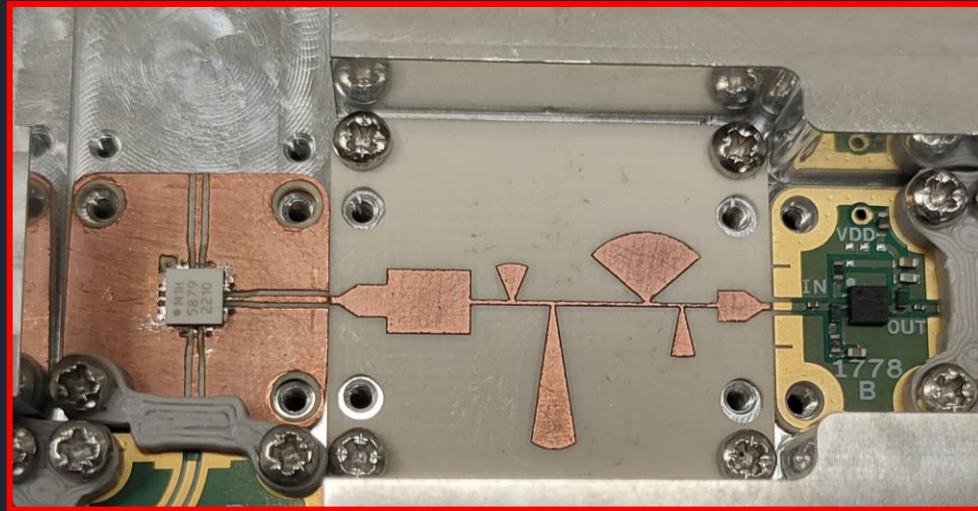
A high-power 2.95GHz radar caused receiver saturation in the VGOS band, leading to the design of HTS

- To reject the RFI.
- In front of the LNA to protect it.



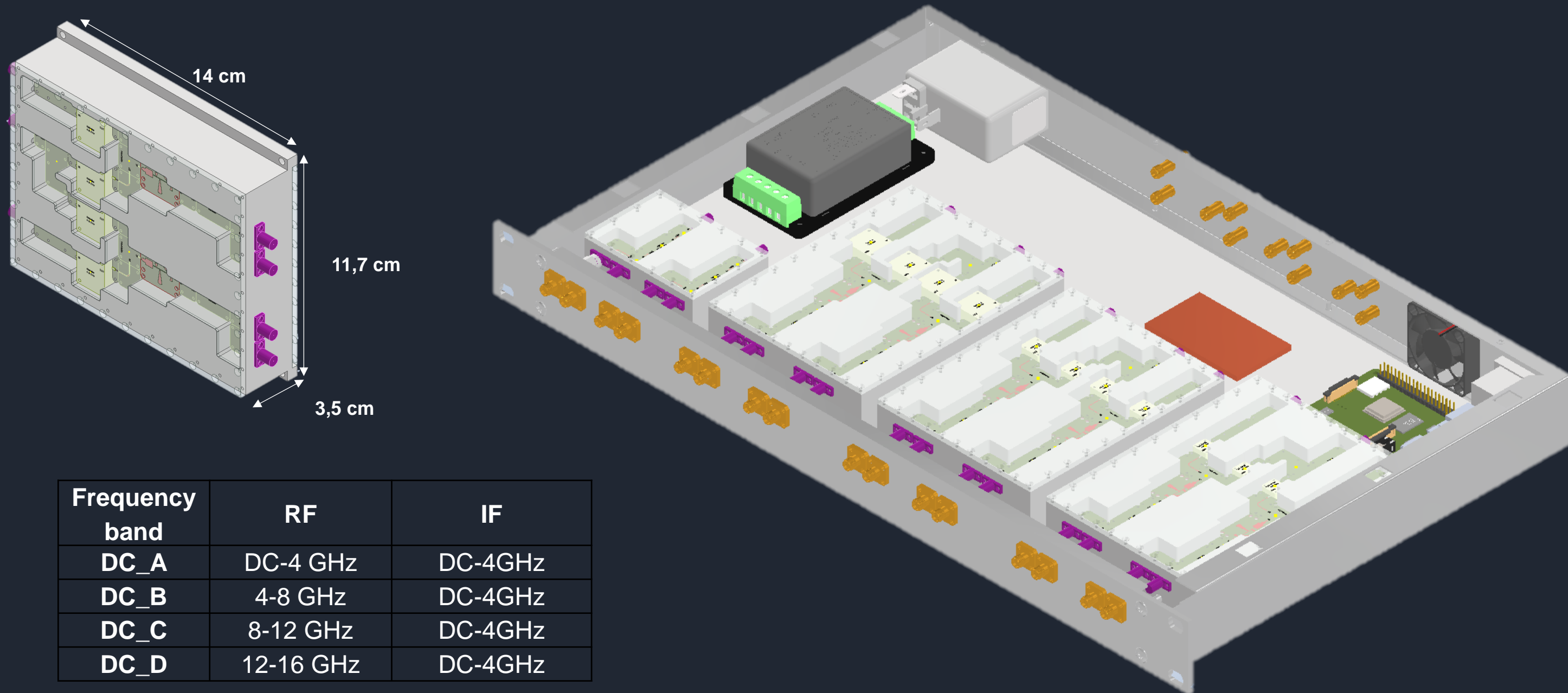
See Pablo García-Carreño "Development of HTS filters at Yebes Observatory", 9th IVTW-2024, Haystack Observatory

Compact down-converters

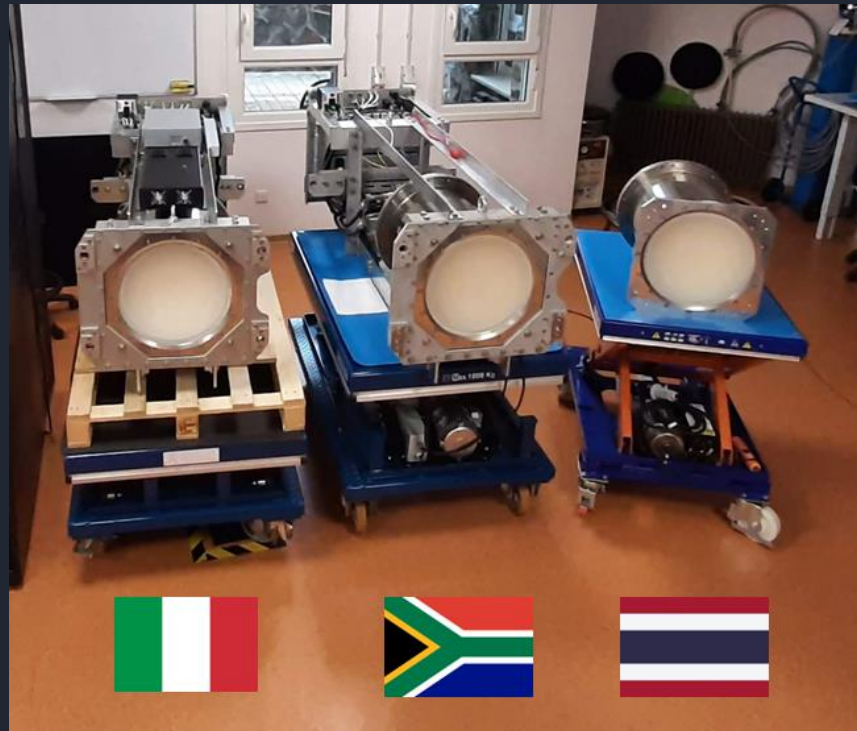


A new development centered on SMD components and microstrip circuits, involving either in-house designs manufactured in our workshops or components purchased from X-Microwave.

Compact down-converters



Latest VGOS receivers



VGOS NARIT

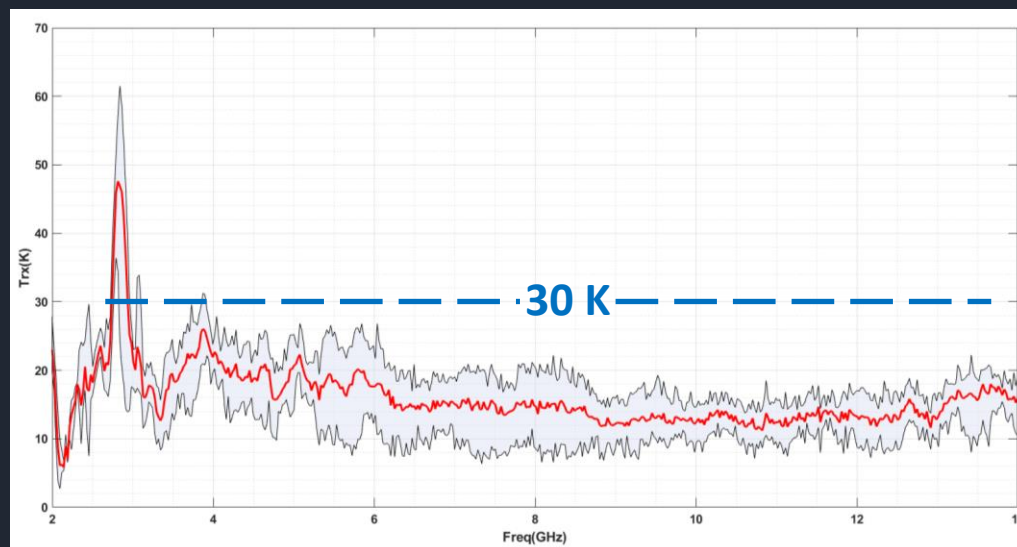
- Dewar + Frontend integrated, tested and shipped by Dec. 2023.
- NARIT is developing the room-temperature electronics

VGOS HartRAO - SARAO

- Full receiver (except backends) integrated, tested, shipped
- Installed in collaboration with HartRAO in 12-23 Feb. 2024
- Currently under tests to join the VGOS core network

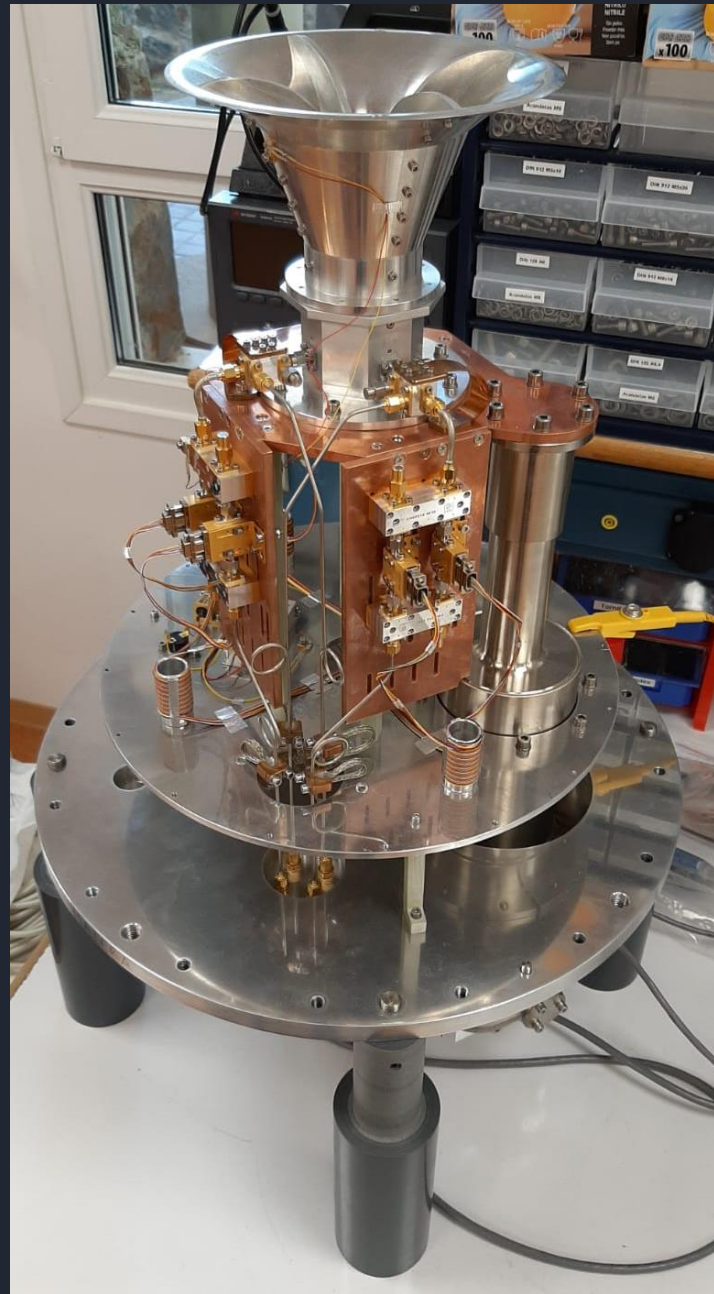
VGOS ASI - Matera

- Full receiver (incl. backends) integrated, tested, shipped
- Installed in collaboration with OHB/eGEOS in 2-12 Apr. 2024
- Currently under tests to join the VGOS core network

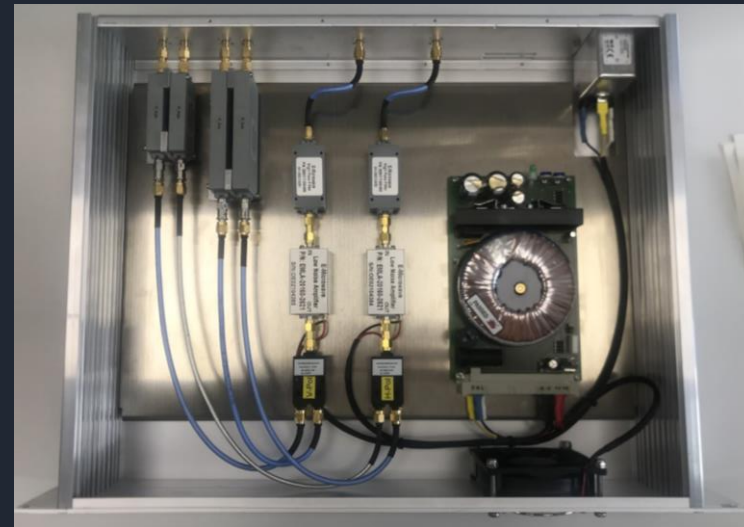


Receiver construction

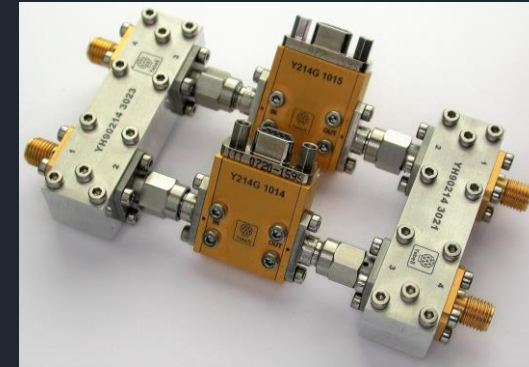
Cooled frontend



FPA: Preamp



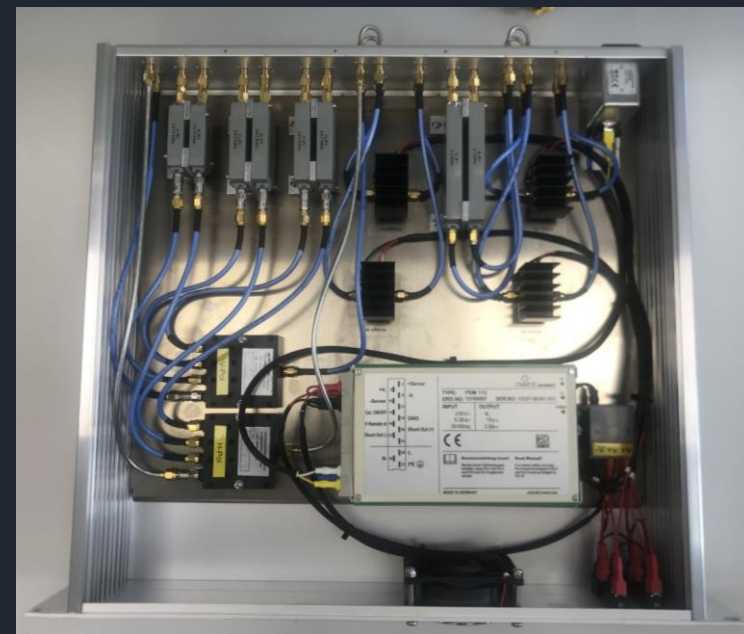
LNAs



Cryo couplers



BSCU: Filter bank

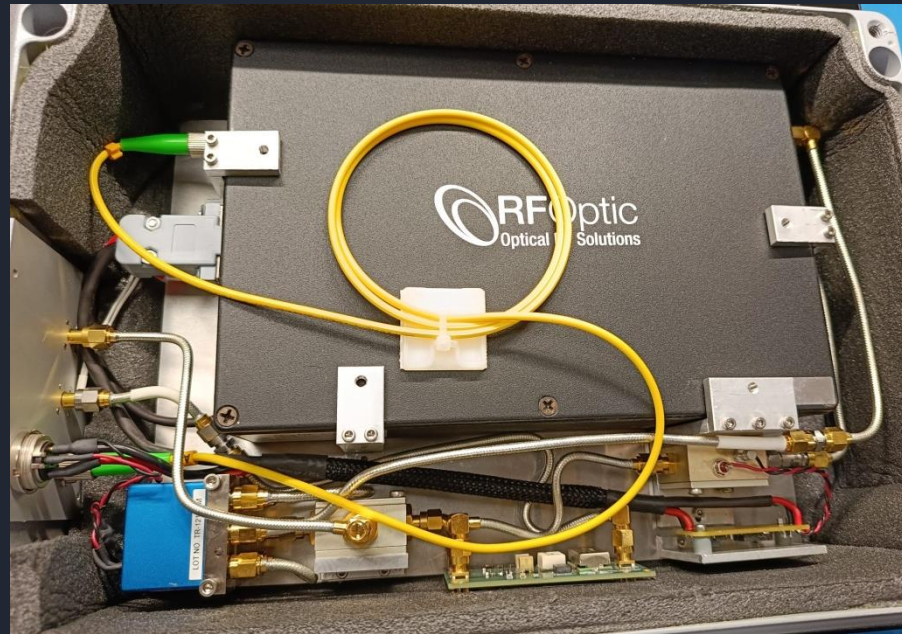


PhaseCal+NoiseCal

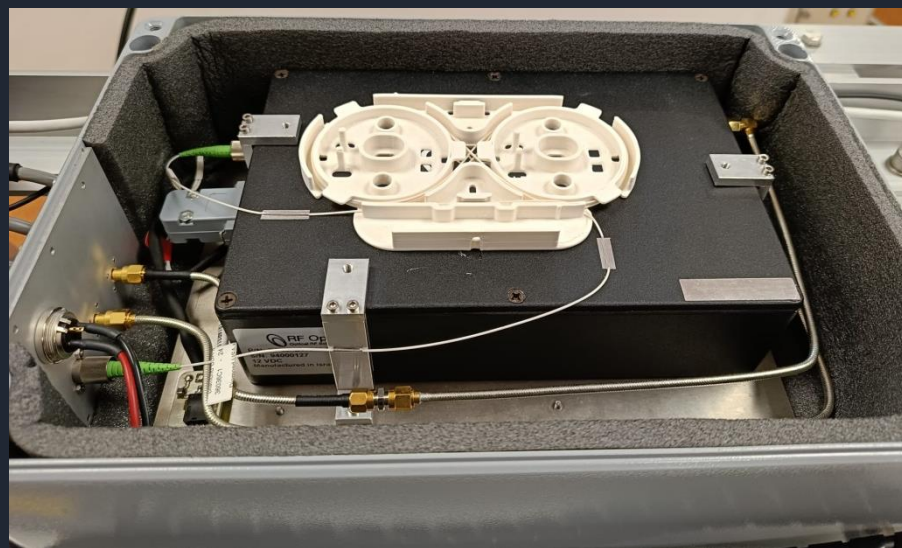


Receiver construction

CDMS GU



CDMS AU



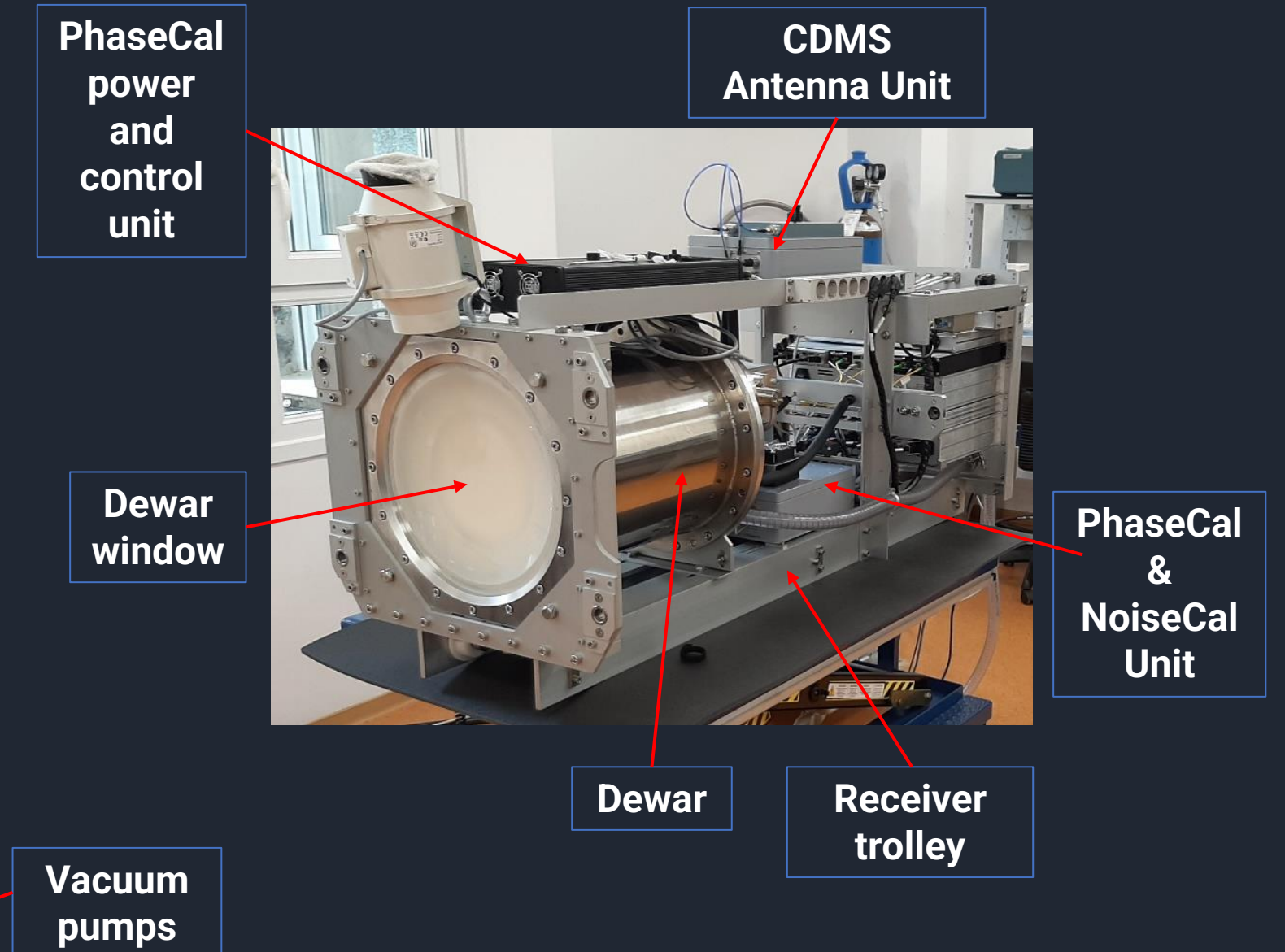
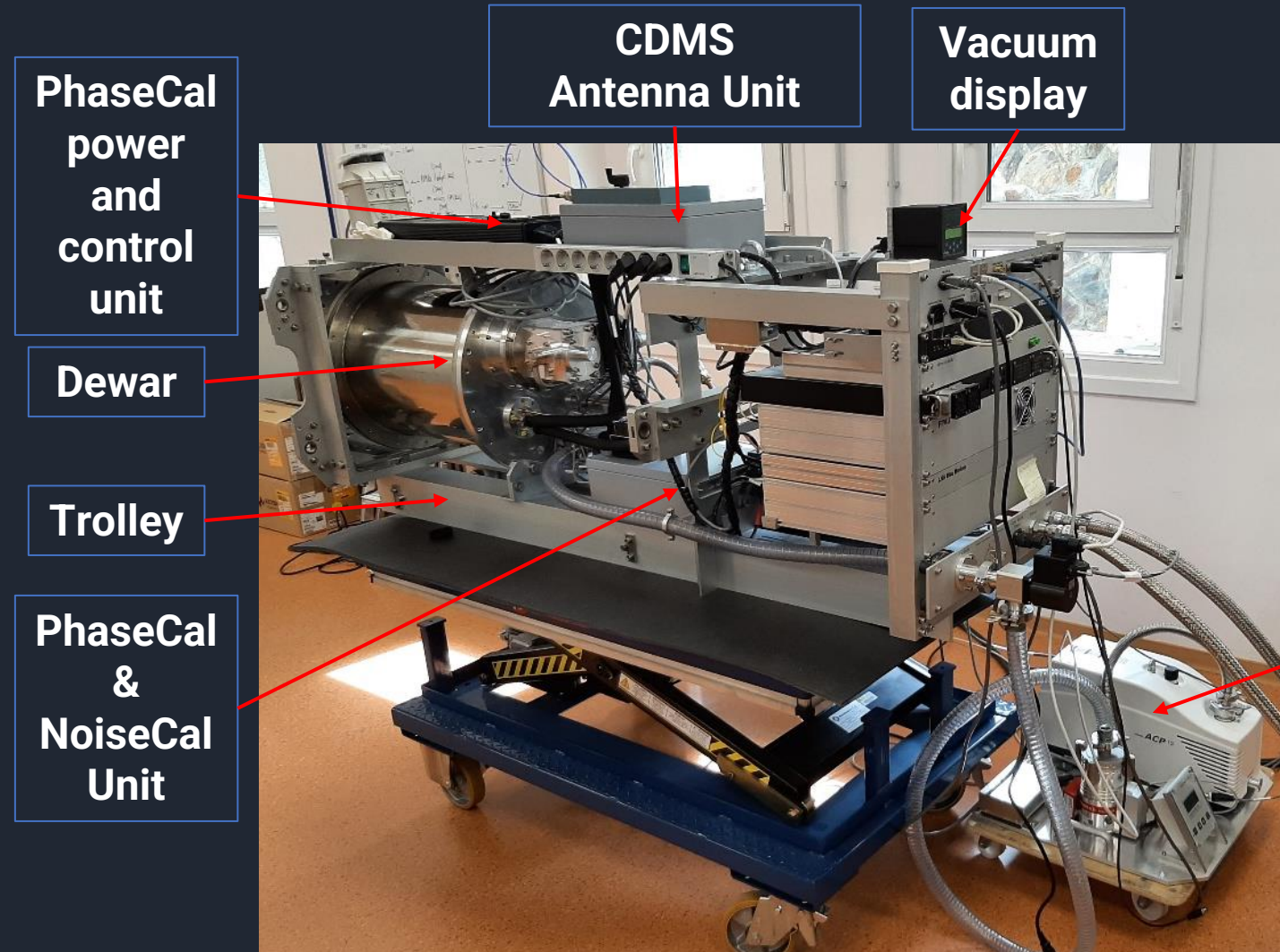
LNAs Biasing Unit



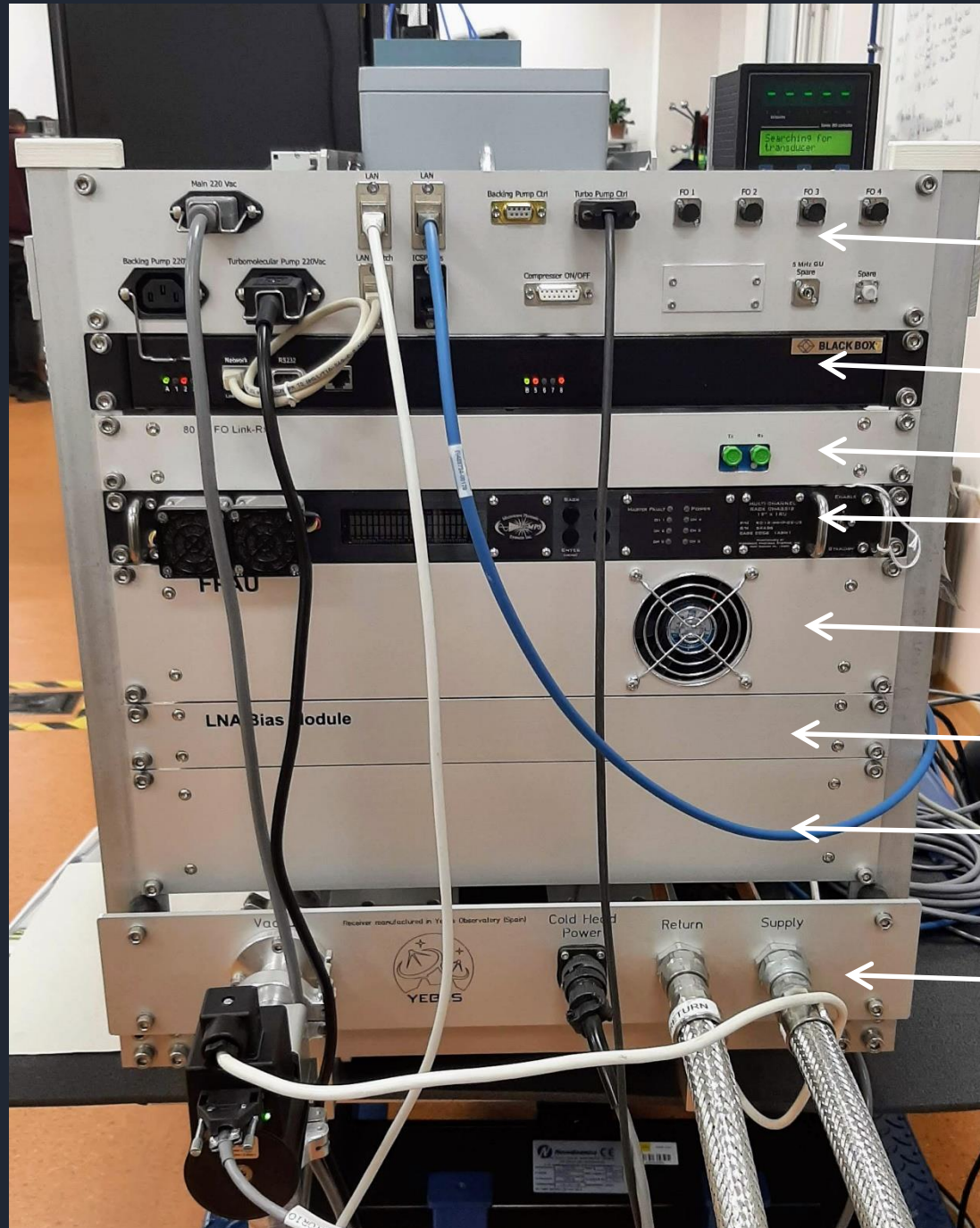
Trolley



Receiver integration



Rx interface plate



Interface plate
(connectors and FO)

Ethernet-controlled
socket strip

80Hz FO link transceiver

RFoF link transmitters

FPA module

LNA Bias module

CVCU

Cryo & Vacuum interface plate

Backends and T&F cabinets

2 x 19"/42U racks

CDMS voltmeter

CDMS GU

CISCO
switches

Mark 6

Power meter

80Hz link transceiver

RFoF link receivers

BSCU

dBBC3

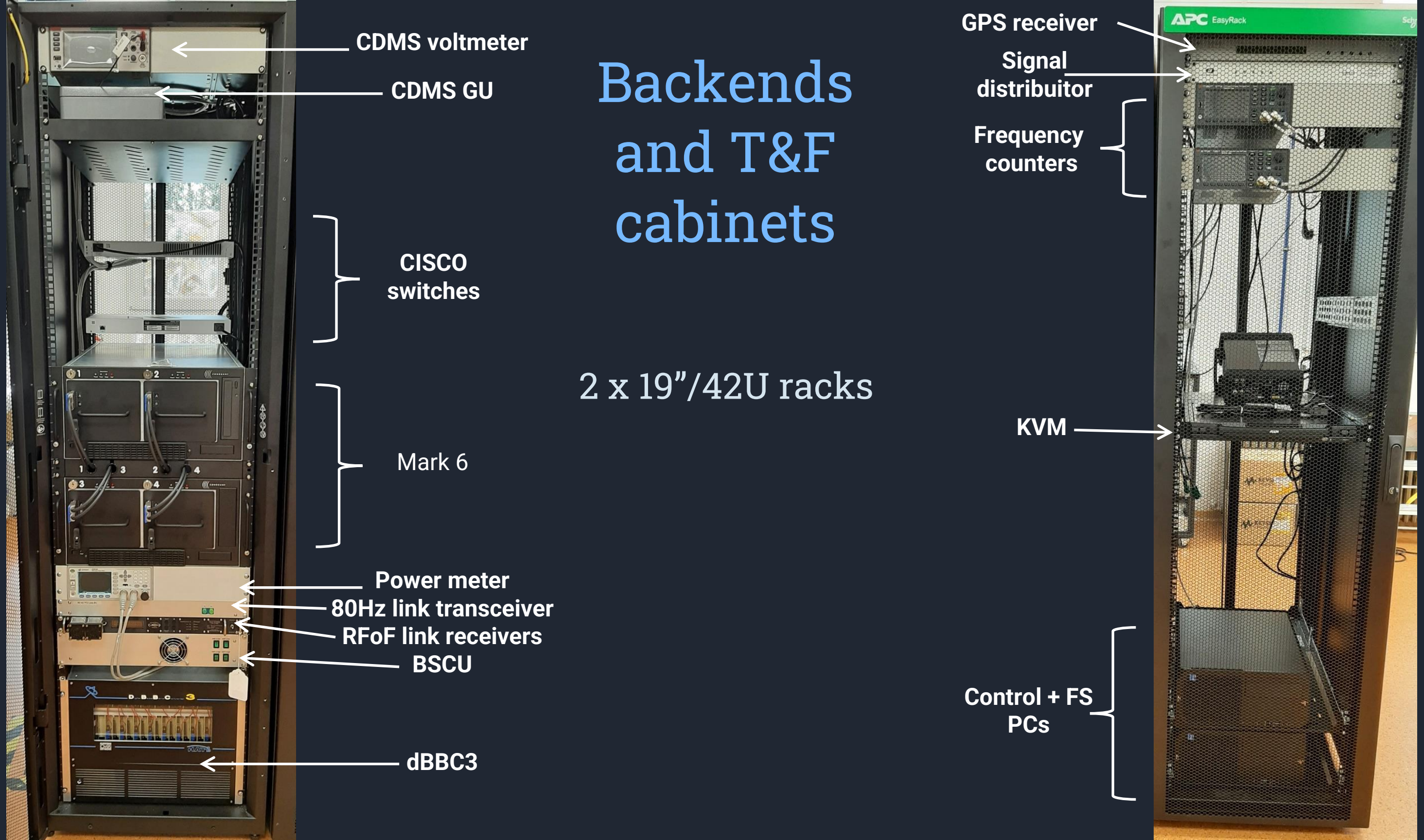
GPS receiver

Signal
distributor

Frequency
counters

KVM

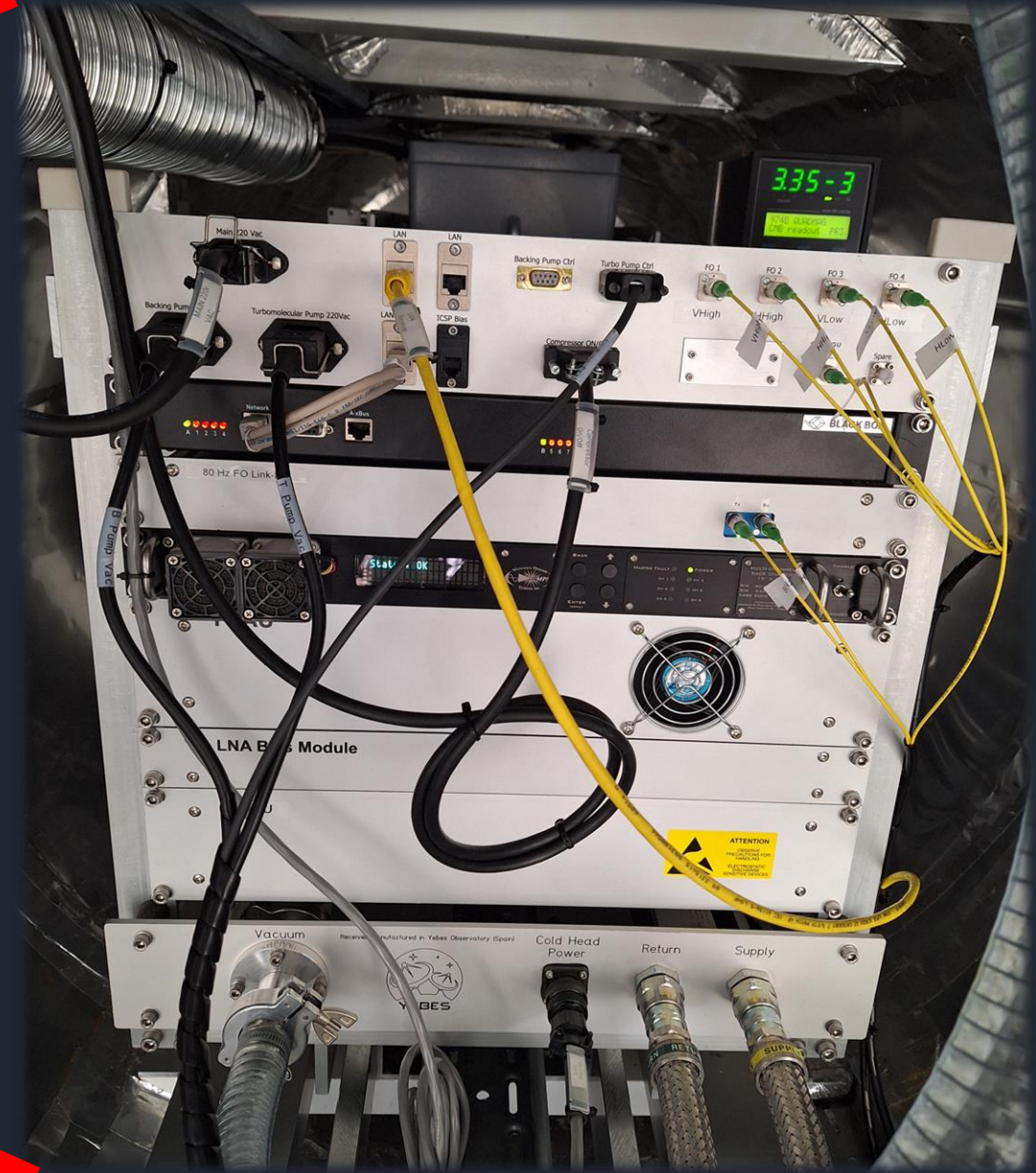
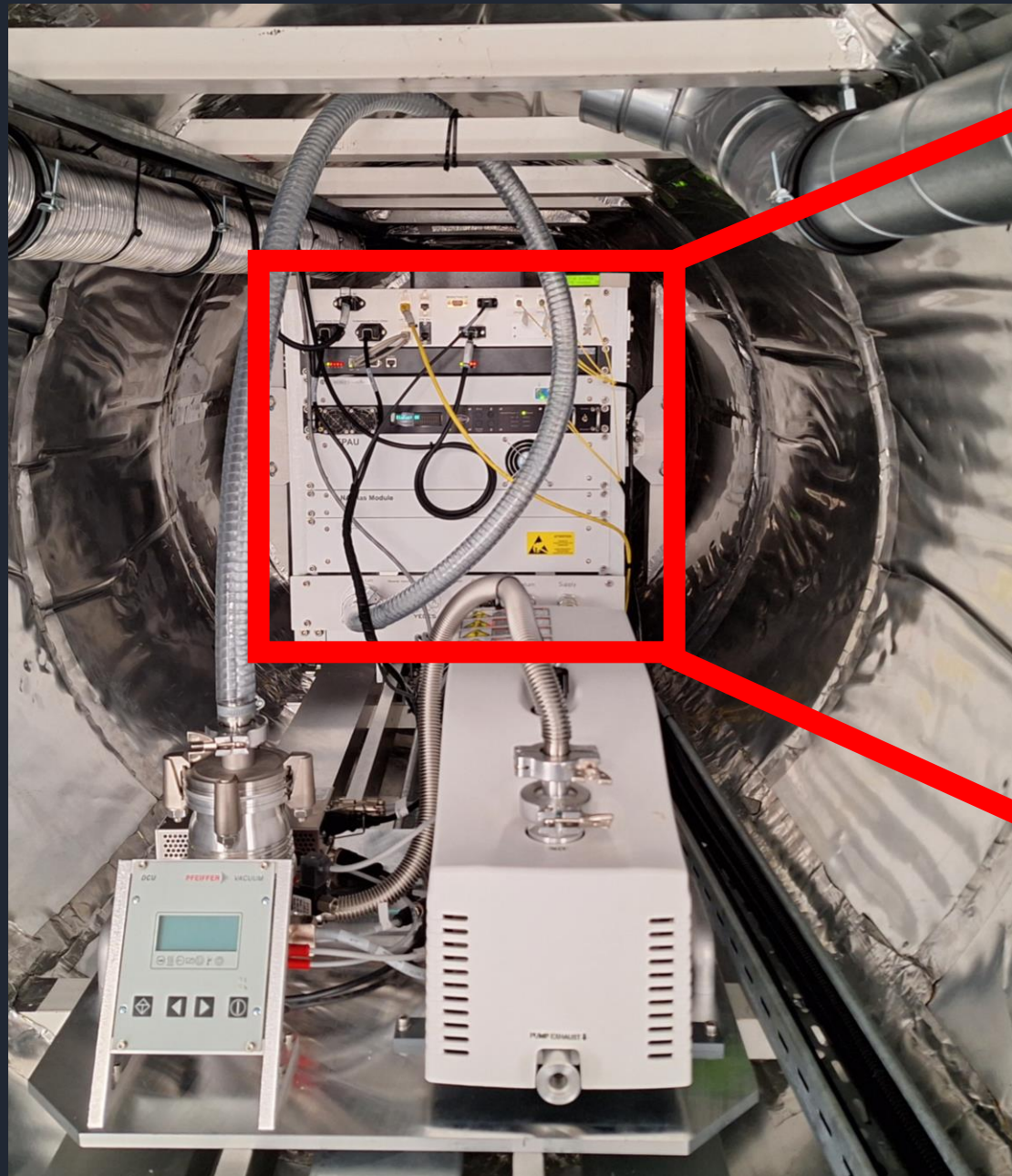
Control + FS
PCs



Receiver installation: April 3rd, 2024

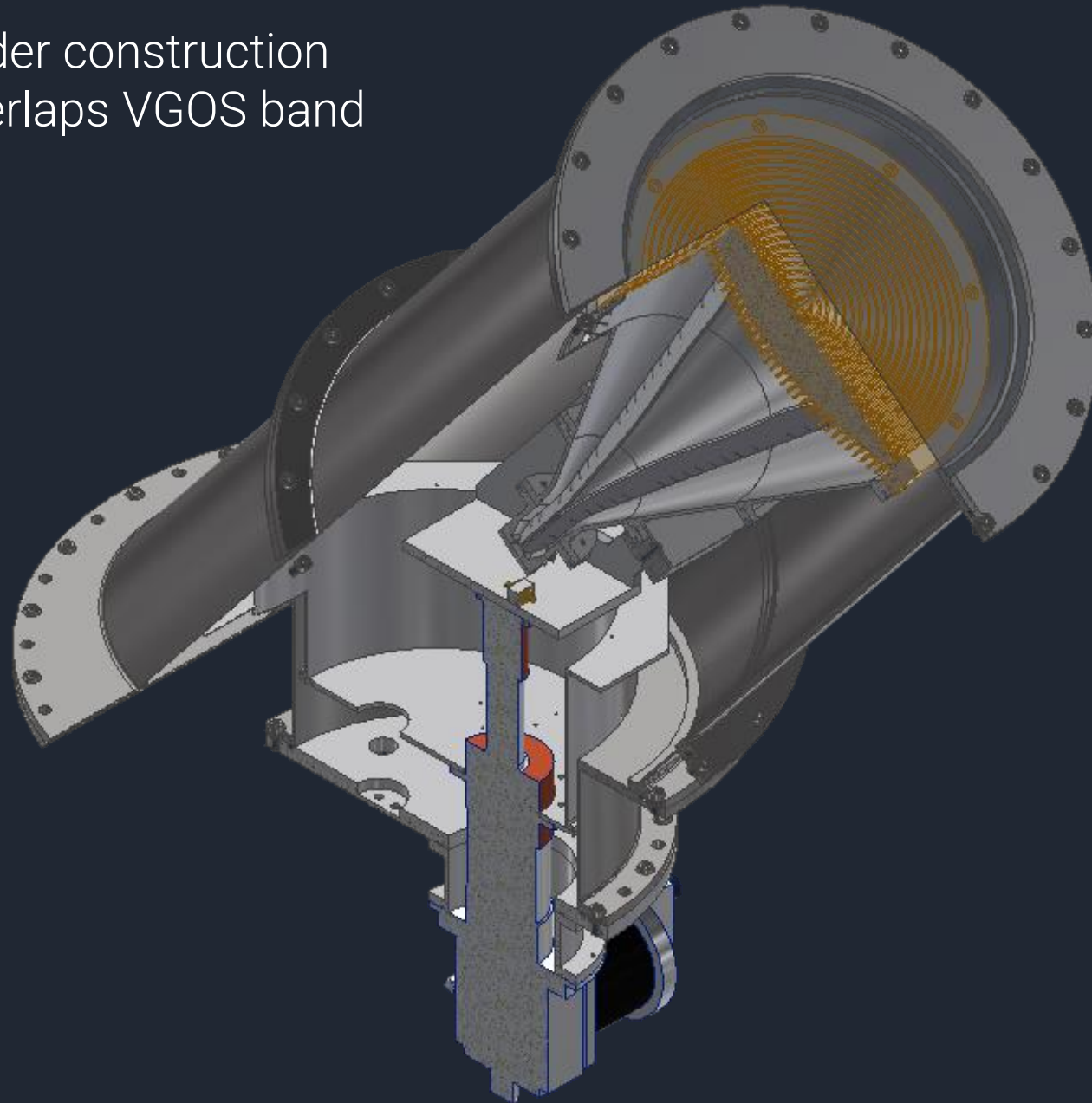


Receiver installation



YNART 4-18 GHz receiver for 40-m telescope

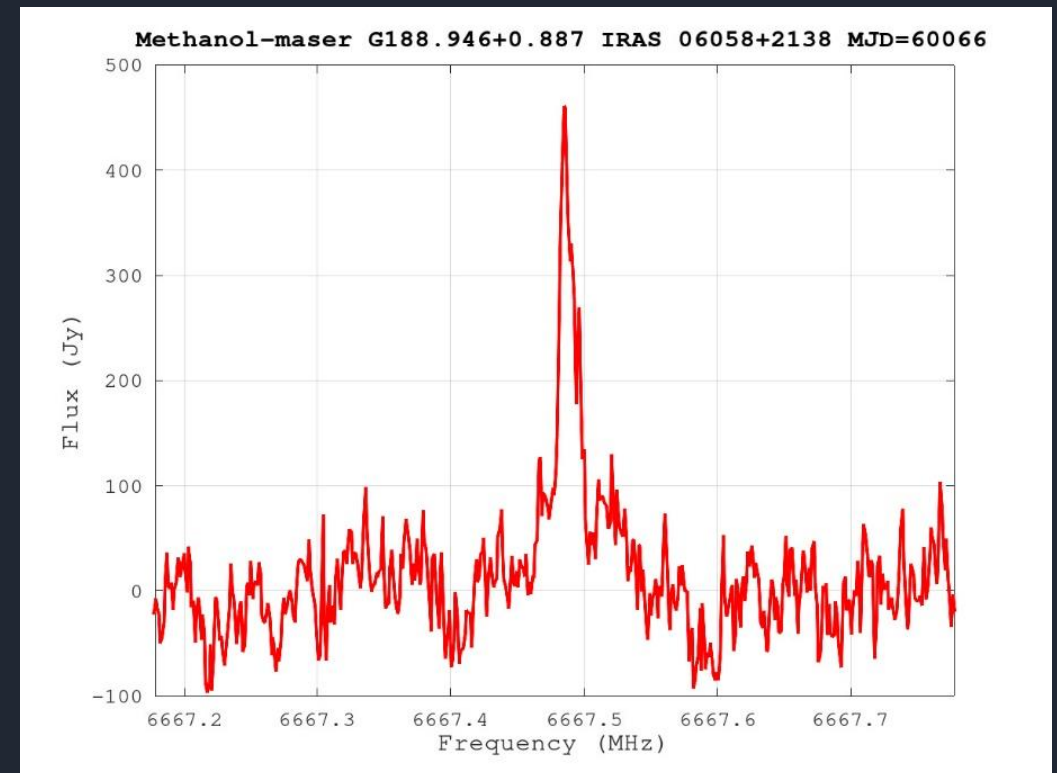
- Under construction
- Overlaps VGOS band



Science cases: astro-VLBI and methanol masers

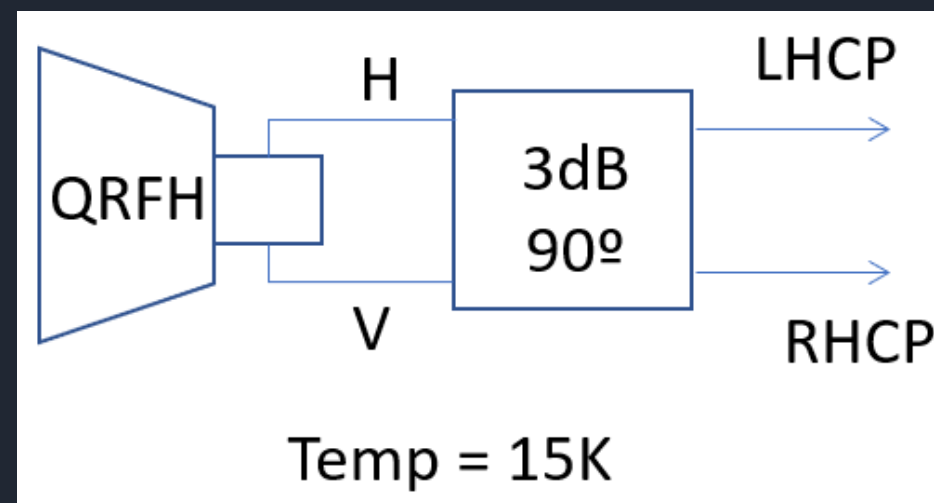
Specifications:

- Operating band: 4 – 18 GHz
- $T_{rx} < 20$ K
- Dual linear polarization
- Noise calibration and phase calibration injection



Linear-to-circular hardware converter

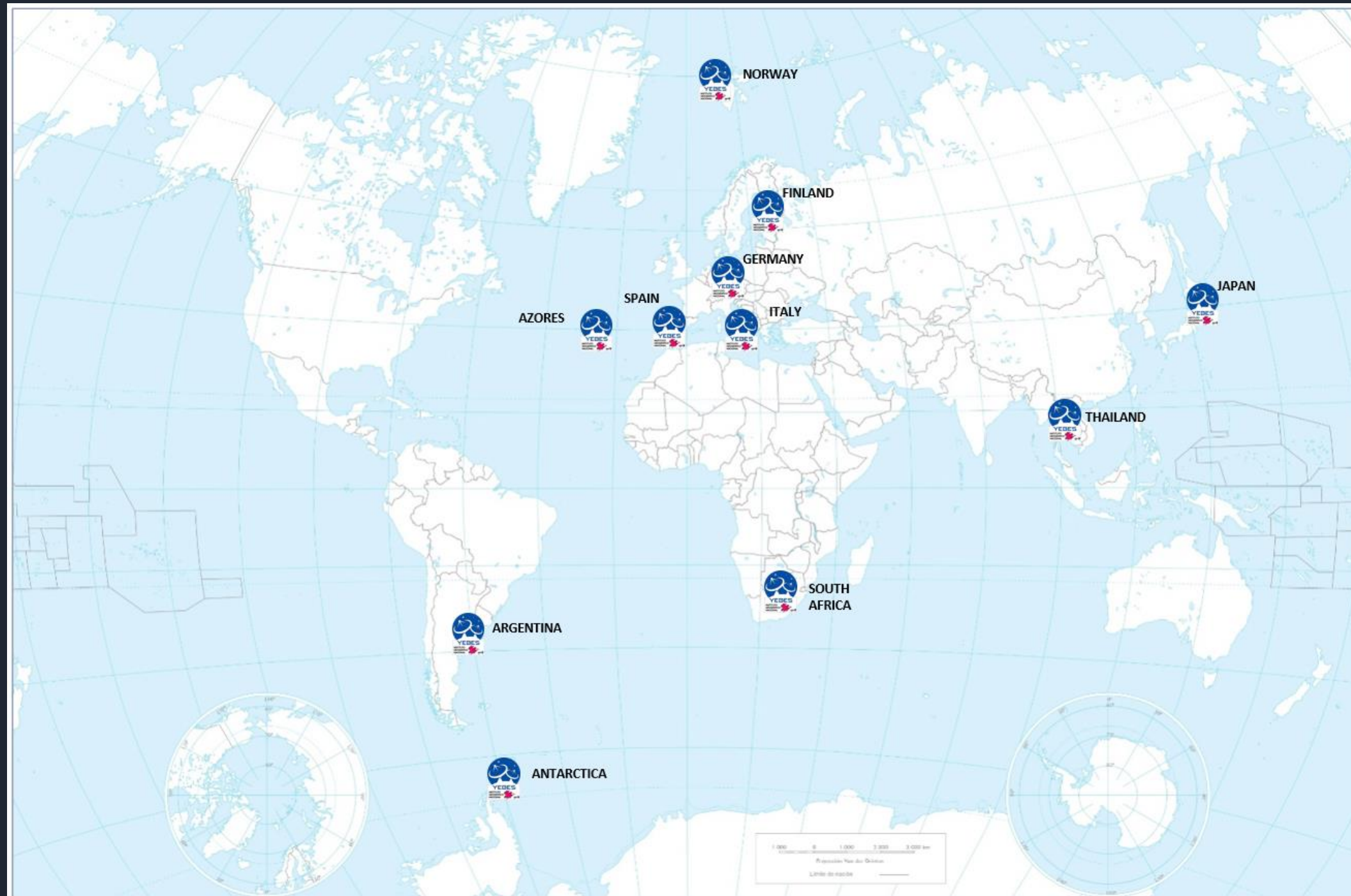
- Adding a cryogenic 3dB/90° hybrid circuit, circular polarizations can be easily obtained by hardware
- Degradation of LNA noise by 1-3 Kelvin
- Concept already tested in EVN-BRAND 1.5-15.5 GHz receiver



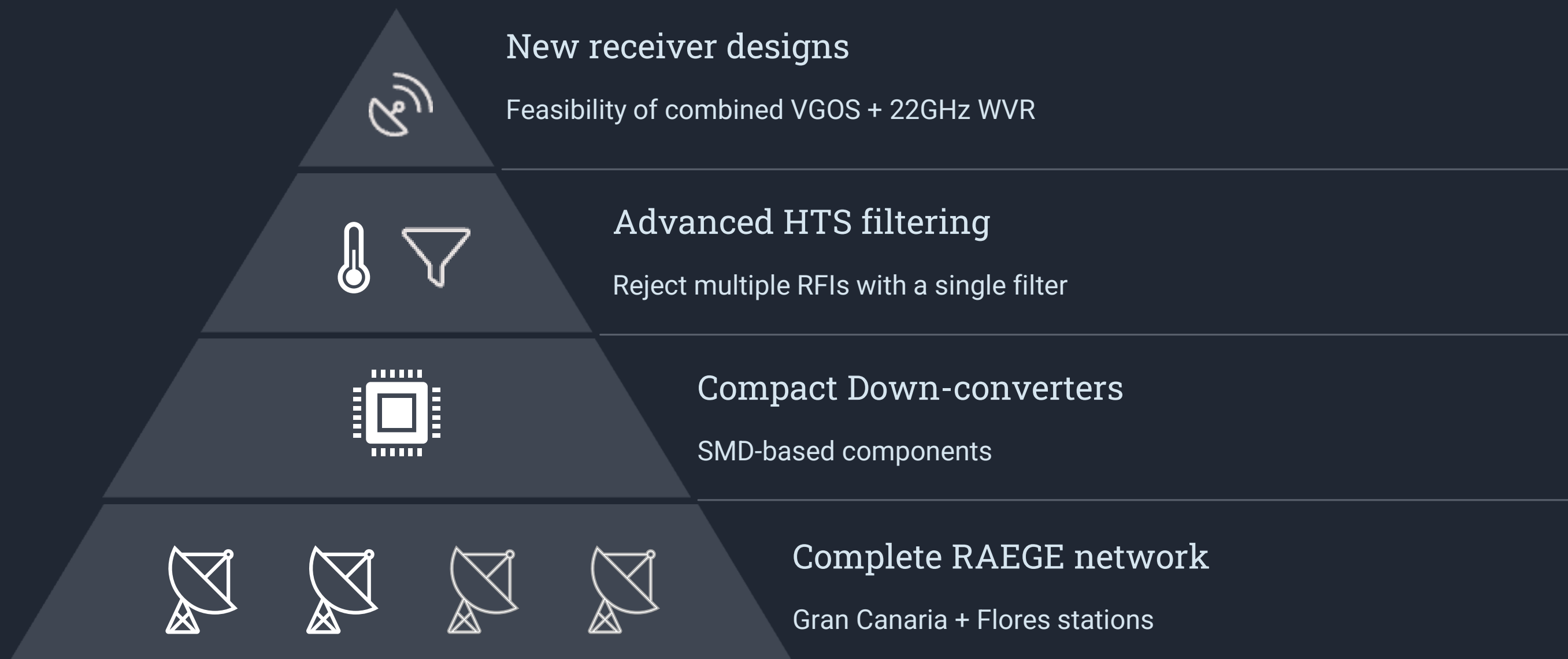
O. García-Pérez, F. Tercero, I. Malo, J. A. López-Pérez: “**Linear to circular polarization conversion using microwave hybrids for VGOS (2-14 GHz)**”, CDT Technical Report 2018-13.

I. Malo-Gómez, J. D. Gallego-Puyol, C. Díez-González, et al., “**Cryogenic hybrid coupler for ultra-low-noise radio astronomy balanced amplifiers,**” *IEEE Trans. Microw. Theory Tech.*, vol. 57, pp. 3239-3245, Nov. 2009.

Yebes receivers around the world



Future Works



Thank you for your attention



Remember: "Talent win games, but teamwork win championships"

(Michael Jordan)